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SPACE THE FINITE FRONTIER

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SPACE, THE FINITE FRONTIER

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Reachable and controllable sets for electric propulsion spacecraft are important to many problems including: dynamic replanning, robust mission design, space situational awareness, assessing advanced concepts, and threat assessments. Current methods result in a two-point boundary value problem or are limited in their application. We solve the reachable and controllable problem by formulating it as a multi-stage indirect approach. We demonstrate that this enables rapid, reliable, and autonomous estimates of the reachable and controllable set. We show that our approach works in strong multi-body environments (i.e., flybys).

INTRODUCTION

Range as a function of fuel is an important concept. It is needed for dynamic fleet replanning, context-based mission orders, robust mission design, situational awareness, autonomous operations, and designing advanced concepts such as basing versus refueling trades. In addition, it has important implications for human capital and operations.

Air, land, and sea have long used fuel-range calculations. Electric propulsion (EP) space systems do not have an equivalent analog. Work by Holzinger¹ derived the general equations for a continuous indirect formulation and showed that these problems can be solved. The indirect approach results in a two-point boundary value problem which is difficult to solve. Bryson and Ho characterized two-point boundary value problems as:²

The main difficulty with these methods is getting started; i.e., finding a first estimate of the unspecified conditions at one end that produces a solution reasonably close to the specified conditions at the other end. The reason for this peculiar difficulty is that external solutions are often very sensitive to small changes in the unspecified boundary conditions. This extraordinary sensitivity is a direct result of the nature of the Euler-Lagrange equations,...

Bando et. al³ use a power series expansion of the generating function to identify low thrust transfers in the Hill frame. Generating functions provide the solution flow; however, current techniques only work on dynamical systems with polynomial forms (e.g., Hill problems).⁴ Thorne⁵ derived a set of formulas that produce approximate values for both the initial Lagrange co-states and the associated optimal flight time needed to solve the minimum-time, continuous-thrust orbital trajectory design

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problem. Thorne’s formulas are especially accurate for either high-thrust or short-duration transfers starting from circular conditions, but they have also been used effectively to solve low-thrust, noncircular, three-dimensional problems with the continuation method.

Our contribution is to develop a fast, scalable, and automated algorithm that can rapidly and reliably estimate the reachable and controllable set for EP spacecraft. Our algorithm does not require the need for any initial guess by users because we structure the problem to avoid the need to solve a two-point boundary value problem. We test the algorithm in multi-body environments including near-lunar flybys. This demonstrates that our approach works in highly challenging dynamic environments.

We achieve these results by:

- Formulating the reachable problem as an indirect multi-stage problem
- Writing the cost function in a linear form
- Structuring the optimal control problem so that a non-thrusting trajectory is a valid solution

We begin by reviewing the multi-stage formulation and discuss notation. Next, we derive the specific equations used in our formulation of the reachable problem. We then show how this formulation can be rapidly and automatically solved with several examples on consumer hardware.

INDIRECT MULTI-STAGE FORMULATION AND NOTATION

We consider a controlled dynamical system defined by:

$$\dot{\mathbf{x}} = \mathbf{f}(\mathbf{x}, t, \mathbf{u}, \mathbf{p}) \quad (1)$$

Where a ***bold*** indicates a vector or matrix while a *normal* font indicates a scalar quantity. $\mathbf{x}$ represents the state vector; $\mathbf{u}$ is the control vector, which is in general a function of time; t is time, and $\mathbf{p}$ are a set of parameters that are constant in the system.

This system can be transformed into a set of indirect multi-stage formulation equations.² A trajectory segment i is governed by:

$$\mathbf{x}^{i+1} = \mathbf{F}^i(\mathbf{x}^i, t^i, \mathbf{u}^i, \mathbf{p}) \quad (2)$$

where

$$\mathbf{F}^i = \mathbf{x}^i + \int_{t_i}^{t_{i+1}} \dot{\mathbf{x}} dt = \mathbf{x}^i + \int_{t_i}^{t_{i+1}} \mathbf{f}(\mathbf{x}, t, \mathbf{u}, \mathbf{p}) dt \quad (3)$$

A trajectory is then composed of a series of N trajectory segments, where $N \geq 1$. For a multi-segment trajectory we can integrate Eq. (2) from i to $i + 1$, then recursively use $i + 1$ to obtain $i + 2$ and so on. In our current formulation, within each segment the controls $\mathbf{u}(i)$ are fixed. This is not a significant limitation as time varying inputs can be modeled by having $\mathbf{u}(i)$ specified by constant coefficients to a time varying function.

The cost function for a multi-stage formulation is given by:

$$J = \phi(\mathbf{x}^N, N) + \sum L(i, \mathbf{x}^i, \mathbf{u}^i) \quad (4)$$

where $\phi(\mathbf{x}_N, N)$ is the terminal cost function and $L(i, \mathbf{x}, \mathbf{u}) = L^i$ are the integral cost function evaluated at the i th step. This leads to a Hamiltonian of:

$$H = \sum H^i \quad (5)$$

$$H^i = \boldsymbol{\lambda}^{i+1} \cdot \mathbf{F}^i + L^i + \sum_k \nu_k C_k \quad (6)$$

For notational convenience later, we specify the adjoints with an index $i + 1$, and C represents the per stage constraints.

Taking the partial with respect to $\mathbf{x}$ yields the dynamics equations for the co-states.

$$H_{\mathbf{x}}^i = \boldsymbol{\lambda}^i = \mathbf{F}_{\mathbf{x}}^{iT} \boldsymbol{\lambda}^{i+1} + L_{\mathbf{x}}^i + \sum_k \nu_k C_{k,\mathbf{x}} \quad (7)$$

where T is the transpose operator. The control law is then given by Pontryagin's principle, here by taking the partial of the Hamiltonian with respect to the control and solving for when this is zero and hence an extremal. This is:

$$H_{\mathbf{u}}^i = \mathbf{0} = \mathbf{F}_{\mathbf{u}}^{iT} \boldsymbol{\lambda}^{i+1} + L_{\mathbf{u}}^i + \sum_k \nu_k C_{\mathbf{u}}^k \quad (8)$$

The initial and terminal conditions are given as:

$$\mathbf{C}^0(\mathbf{x}) = \mathbf{0} \quad \mathbf{C}^N(\mathbf{x}) = \mathbf{0} \quad (9)$$

which result in the initial and terminal co-states being:

$$\boldsymbol{\lambda}^0 = \nu^0 \cdot \mathbf{C}_{\mathbf{x}}^0 \quad \boldsymbol{\lambda}^N = \phi_{\mathbf{x}} + \nu^N \cdot \mathbf{C}_{\mathbf{x}}^N \quad (10)$$

Taken together, these represent the general equations for the multi-stage optimization problem. Next, we specify our particular formulation.

REACHABLE EQUATIONS

The equations of motion we use in this paper are stated in an inertial frame with an arbitrary origin:

$$\dot{\mathbf{x}} = \begin{bmatrix} \dot{\mathbf{r}} \\ \dot{\mathbf{v}} \\ \dot{n} \end{bmatrix} = \begin{bmatrix} \mathbf{v} \\ \sum_j -\frac{\mu_j}{|\mathbf{r} - \mathbf{r}^j|^3} (\mathbf{r} - \mathbf{r}^j) + n\mathbf{T} \\ n^2 \frac{T_{mag}}{c} \end{bmatrix} \quad (11)$$

where μ_j is the gravitational parameter of the attracting body located at $\mathbf{r}^j$, $\mathbf{T}$ is the thrust vector, T_{mag} is the magnitude of thrust, $\mathbf{r}$ is the position vector, $\mathbf{v}$ is the velocity, n is the inverse mass, and c is one Earth gravity times the specific impulse. The cost function is defined as:

$$J = -min \quad \hat{\mathbf{z}} \cdot \mathbf{x}^N \quad (12)$$

This represents maximizing $\mathbf{x}^N$ along the direction $\hat{\mathbf{z}}$, which is a specified vector which will be swept over to sample the reachable set. We also require that $|\mathbf{T}| = T_{mag}$ and $0 \leq |\mathbf{T}| \leq T_{max}$. These conditions require the thrust must be less than the maximum allowable thrust, T_{max} , and that

the total thrust and T_{mag} are equal. The latter simply enforces the constraint that the thrust used to generate acceleration is equal to the thrust supplied by the engine.

The Hamiltonian is then:

$$H^i = \mathbf{F}^{iT} \boldsymbol{\lambda}^{i+1} + \nu_1(|\mathbf{T}| - T_{mag}) + \nu_2\left(\frac{1}{2}T_{mag}^2 - \frac{1}{2}T_{max}^2\right) \quad (13)$$

The co-state equation is:

$$H_x^i = \boldsymbol{\lambda}^i = \mathbf{F}_x^{iT} \boldsymbol{\lambda}^{i+1} \quad (14)$$

and is formulated to move backwards in time. The term $\mathbf{F}_x^i$ is the state transition matrix at the end of the segment. It is given by

$$\mathbf{F}_x^i = \frac{\partial \mathbf{x}^{i+1}}{\partial \mathbf{x}^i} = \Phi(t^{i+1}, t^i) = \int_{t^i}^{t^{i+1}} \begin{bmatrix} \mathbf{0} & \mathbf{I} & \mathbf{0} \\ \frac{\partial \dot{\mathbf{v}}}{\partial \mathbf{r}} & \mathbf{0} & \mathbf{T} \\ 0 & 0 & 2n\frac{T_{mag}}{c} \end{bmatrix} \Phi(t, t^i) dt \quad (15)$$

where Φ is the state transition matrix. The initial conditions for $\Phi(i, i) = \mathbf{I}$

The control law is defined by:

$$H_u = 0 = \mathbf{F}_u^{iT} \boldsymbol{\lambda}^{i+1} + \nu_1 \begin{bmatrix} \frac{\mathbf{T}}{|\mathbf{T}|} \\ -1 \end{bmatrix} + \nu_2 \begin{bmatrix} 0 \\ T_{mag} \end{bmatrix} \quad (16)$$

and

$$\mathbf{F}_u^i = \frac{\partial \mathbf{x}^{i+1}}{\partial \mathbf{u}^i} = \Omega(t^{i+1}, t^i) = \int_{t^i}^{t^{i+1}} \begin{bmatrix} \mathbf{0} & \mathbf{I} & \mathbf{0} \\ \frac{\partial \dot{\mathbf{v}}}{\partial \mathbf{r}} & \mathbf{0} & \mathbf{T} \\ 0 & 0 & 2n\frac{T_{mag}}{c} \end{bmatrix} \Omega(t, t^i) + \begin{bmatrix} \mathbf{0} & \mathbf{0} & \mathbf{0} & \mathbf{0} \\ - & \text{diag}(\mathbf{n}) & - & \mathbf{0} \\ 0 & 0 & 0 & \frac{n^2}{c} \end{bmatrix} dt \quad (17)$$

The term Ω is analogous to the impulse response matrix for linear systems. The initial condition for the integral is given by $\Omega(i, i) = 0$. Solving Eq. (16) and satisfying the constraints results in:

$$T_{mag} = T_{max}, \quad \text{and} \quad \mathbf{T} = -T_{max} \frac{\mathbf{F}_T^T \boldsymbol{\lambda}}{|\mathbf{F}_T^T \boldsymbol{\lambda}|} \quad (18)$$

where

$$\nu_1 = |\mathbf{F}_T^T \boldsymbol{\lambda}^{i+1}| \quad (19)$$

$$\nu_2 = \frac{\nu_1 - \mathbf{F}_{T_{mag}}^T \boldsymbol{\lambda}^{i+1}}{T_{max}} \quad (20)$$

The initial and terminal conditions are then:

$$\mathbf{x} = \mathbf{x}^0 \quad \boldsymbol{\lambda}^N = -\hat{\mathbf{z}} \quad (21)$$

In these equations $\mathbf{F}_T$ is the columns of $\mathbf{F}_u$ associated with the thrust direction $\mathbf{T}$. Similarly, $\mathbf{F}_{T_{mag}}$ is the column of $\mathbf{F}_u$ associated with T_{mag} .

These equations now provide sufficient information to estimate the reachable and controllable set. One interesting feature is that T_{max} , see Eq. (18), does not affect the direction of thrust. This allows us to integrate along the unpowered trajectory then conduct trade studies by sweeping over various thrust levels.

FAST FIRST ORDER ESTIMATE (FFOE) SOLUTION

We first develop a Fast First Order Estimate (FFOE) of the reachable set. This solution provides a good estimate of the reachable set. We begin by recognizing that the non-thrusting solution with $T = 0$ and $T_{mag} = 0$ can be considered an optimal solution with a zero reachability set. We can then compute the unpowered trajectory and partials (F_x and F_u) for this trajectory. We then select the values of $\hat{z}$ from a unit ball and compute the control law from Eq. (18). The final step is to integrate Eq. (11) using the derived control law to compute the reachable set under this linearization assumption. This is more formally shown in Alg. (1)

Starting with an unpowered trajectory has several practical benefits. It eliminates the need for a user supplied initial guess. The updated thrust laws can be computed using a series of matrix multiplications without requiring additional integrations. As we will show in the Example section, this method is very fast as it reduces the number of first order integrations required to 1 and requires N integrations of the equations of motion, where N is the number of points on the reachable set.

OPEN-LOOP EXACT (OLE) SOLUTION

The open-loop exact (OLE) solution is an open-loop method that iteratively updates the control law until the change in cost function is small. The OLE algorithm is initialized using the FFOE for a given $\hat{z}$. Then the trajectory and partials are numerically integrated and recomputed along the previous control law, using Eq. (18). The difference between the prior and current cost function is then compared. If the change in cost function between iterations is less than a specified precision the algorithm terminates. This is equivalent to terminating when the gradient along the constraint is small.

EXAMPLES

In this section we provide various examples to highlight the performance and utility of the algorithms. For these examples we code the equations in C++ on an M1 MacBook Pro from 2020. We use the Boost Libraries for integration. In particular, we use a Fehlberg 78 order integrator with error tolerances $\leq 10^{-13}$. We use Armadillo^{6,7} for matrix/vector support. We also use the JPL SPICE libraries^{8,9} and DE440 ephemerides.¹⁰ Plotting is done in spacekit.js*. Parallel processing is done by leveraging the BOOST ASIO library. The reference frame is ECLIPJ2000, the specific impulse is set to 3000 seconds, and the Sun, Earth, and Moon are gravitating.

We show three examples to demonstrate the uses of the algorithm. Case 1 is a reachability case with an elliptic orbit around Earth with no significant multi-body effects. The goal of case 1 is to show how the FFOE and OLE compare to one another. Case 2 is also a reachability case with a lunar flyby. This shows that the FFOE algorithm works well with strong multi-body interactions. For case 3, we run case 2 as a controllability problem. This highlights the initial envelope that a spacecraft can start in and still reach the terminal states.

Case 1

The first case represents a relatively straightforward reachability case. The initial orbit is an elliptical trajectory with no significant multi-body effects. The simulation parameters are a max thrust of 0.01 N, 200 segments, and segment duration of 7200 seconds.

*<https://typpo.github.io/spacekit/>

Algorithm 1 FFOE algorithm

```
 $x^0 \leftarrow$  initial conditions  
 $t^0 \leftarrow$  start time  
 $dt \leftarrow$  segment duration  
 $Seg \leftarrow$  number of segments  
 $T \leftarrow \mathbf{0}$   
 $i \leftarrow 0$   
while  $i < Seg$  do  
     $\mathbf{x}^{i+1}, \mathbf{F}_x^i, \mathbf{F}_u^i \leftarrow \mathbf{F}^i(\mathbf{x}^i, \mathbf{T}^i, t_0, t_0 + dt)$   
     $t_0 \leftarrow t_0 + dt$   
     $i \leftarrow i + 1$   
end while  
 $j \leftarrow 0$   
 $N \leftarrow$  Number of sample points  
while  $j < N$  do  
     $i \leftarrow Seg - 1$   
     $\lambda^{Seg} \leftarrow$  from unit ball  
    while  $i \geq 0$  do  
         $\mathbf{T}^{i,j} \leftarrow$  Eq. (18)  
         $\lambda^i \leftarrow$  Eq. (14)  
         $i \leftarrow i - 1$   
    end while  
     $j \leftarrow j + 1$   
end while  
 $j \leftarrow 0$   
while  $j < N$  do  
     $i \leftarrow 0$   
    while  $i < Seg$  do  
         $\mathbf{x}^{i+1} \leftarrow \mathbf{F}^i(\mathbf{x}^i, \mathbf{T}^{i,j}, t_0, t_0 + dt)$   
         $t_0 \leftarrow t_0 + dt$   
         $i \leftarrow i + 1$   
    end while  
     $j \leftarrow j + 1$   
end while
```

We compute the reachable set using the first order methods and the open-loop method. We see both the FFOE solution and OLE solution produce similar results, see Fig. (1). The computational time is 0.27 seconds and 4.32 seconds respectively.

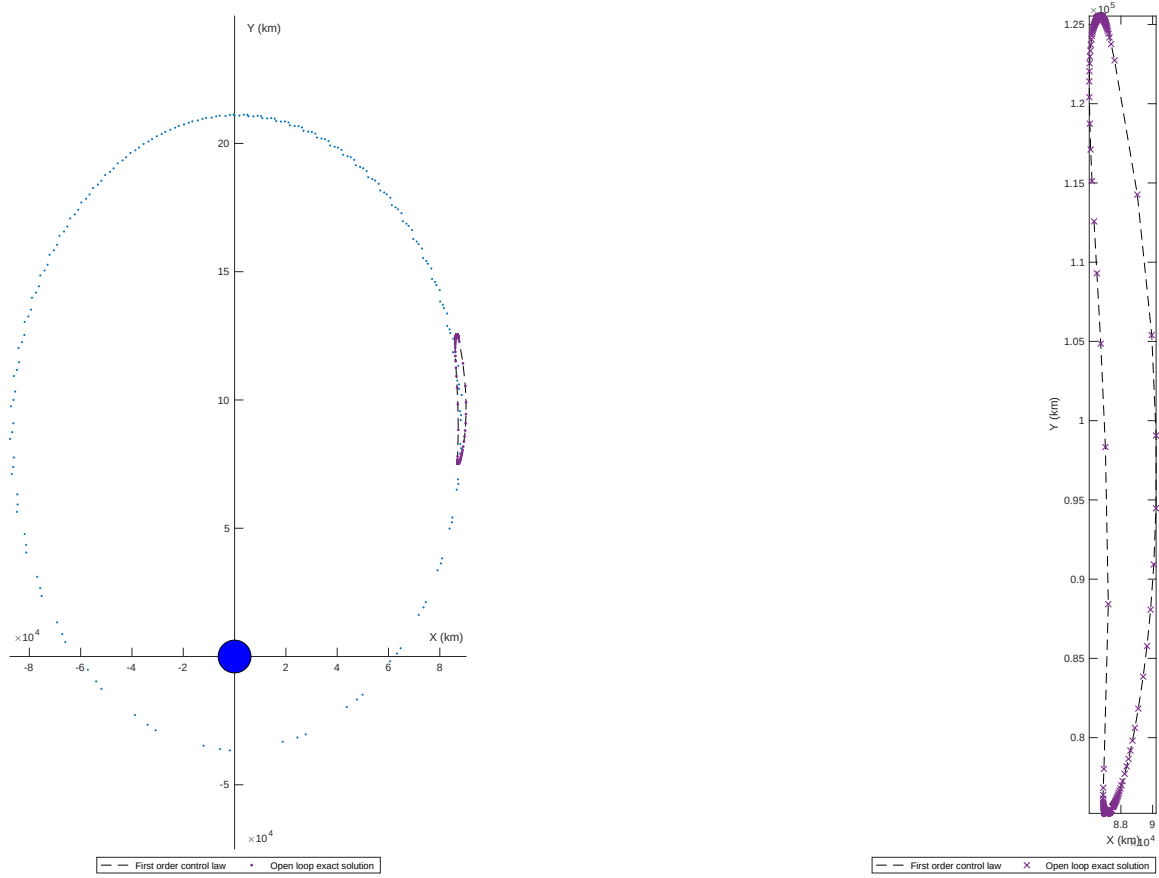


Figure 1. Elliptic trajectory around Earth. No significant multi-body interaction. The epoch and initial conditions are 64.1 seconds and $x^0 = [0, 210820, 0, 0.753136347, 0, 0, 1000]$ Units are km, km/s, and kg.

Case 2

The second case is a multi-body reachability case. The trajectory has a close approach to the moon. This results in strong multi-body effects and an inclination and energy increase. The first order method does a fairly good job; however, we see that the exact method captures the edges better, see Fig. (2). In Fig. (3) we can see the entire trajectory projected onto the X-Y plane relative to the Earth and Moon.

The simulation parameters are a max thrust of 0.02 N, 200 segments, and segment duration of 7200 seconds. The computation time of the FFOE and OLE algorithms is 0.25 seconds and 6.42 seconds respectively. Fig. (4) and Fig. (5) highlight the three dimensional nature of the trajectories by showing the close approach and inclination change.

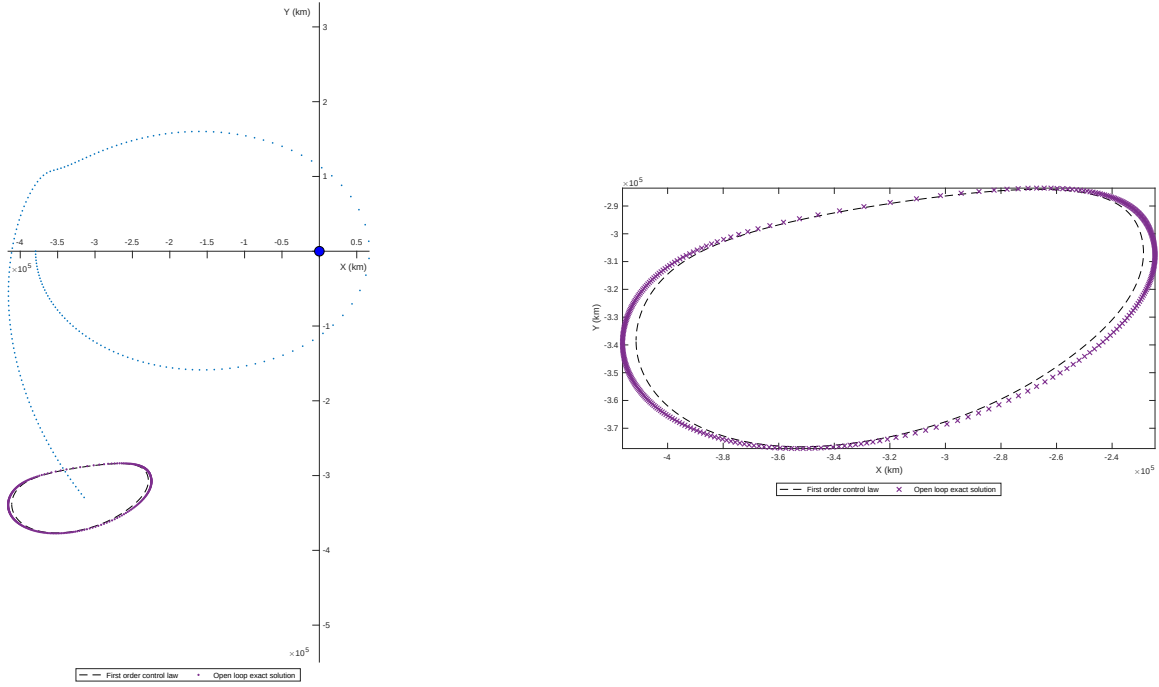


Figure 2. Earth-Sun-Moon gravitating. We show good agreement between the FFOE and OLE algorithm. The OLE solution shows that the extreme edges are underestimated by the FFOE algorithm. The epoch and initial conditions are 1072864.169 seconds and $x^0 = [-379476, 0, 0, 0, -0.561354689, 0, 1000]$. Units are km, km/s, and kg.

Case 3

For the third case, we start at the end of case 2 and run it backwards. This is a controllability case and shows the range of initial conditions that would allow the spacecraft to still reach the terminal state. The simulation parameters are a max thrust of 0.02 N, 200 segments, and segment duration of -7200 seconds. Fig. (6) shows the simulation results. We see that the trajectory is the same as case 2. For the controllability problem, we identified the envelope of initial conditions that enable the final states in case 2 to be reached.

ADDITIONAL APPLICATIONS

We are currently extending the FFOE algorithm to incorporate initial state uncertainty¹ and ΔV impulses. In addition, the authors are exploring how the algorithm can be used to solve fuel-minimum trajectory optimization problems.

CONCLUSION

We demonstrate a fast and automated algorithm to estimate the reachable and controllable set for EP spacecraft. We show that the algorithm works in multi-body cases and scales through the use of parallel/concurrent processing.

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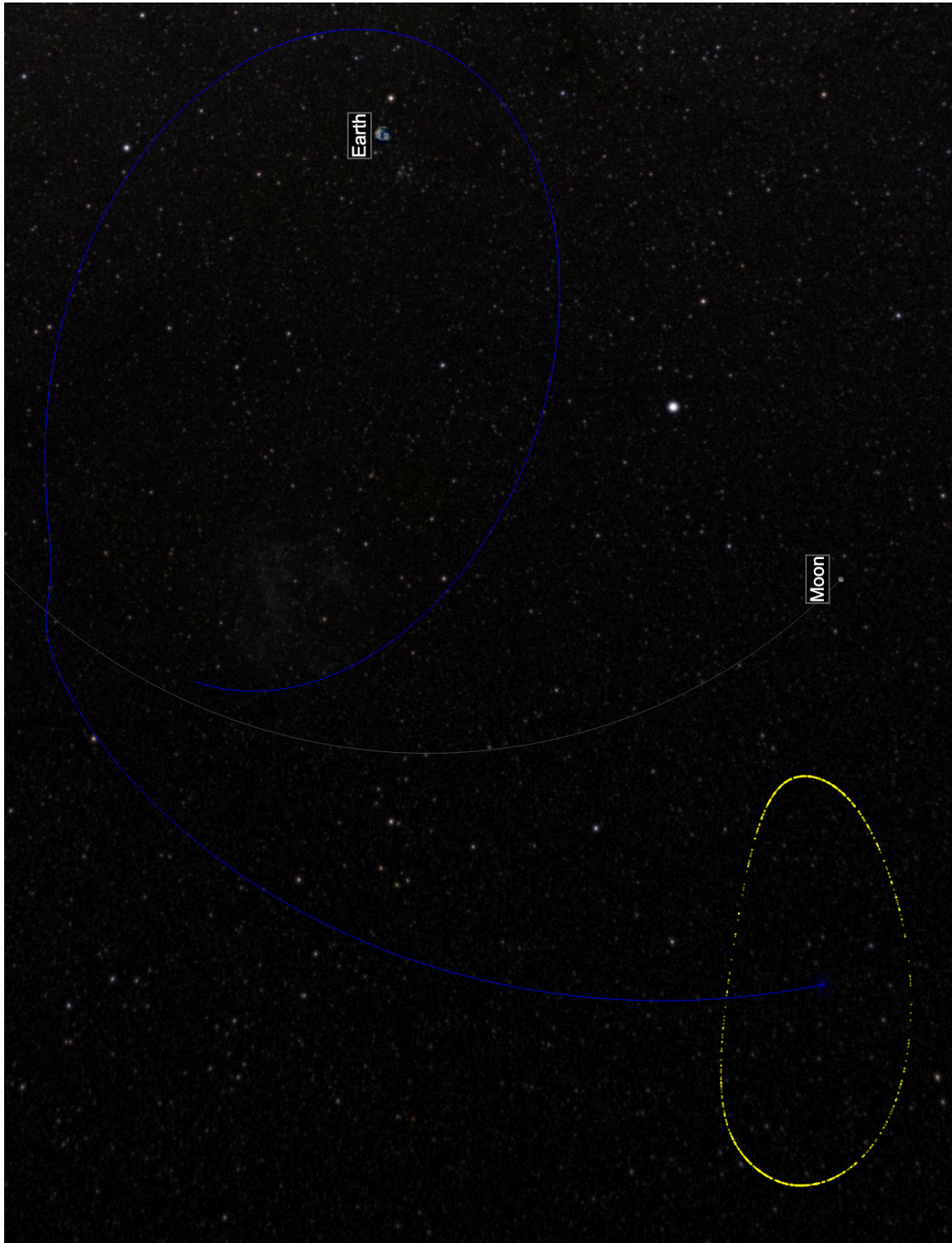


Figure 3. Visualization of simulation results in X-Y plane with Earth, Moon, and spacecraft reachable set.

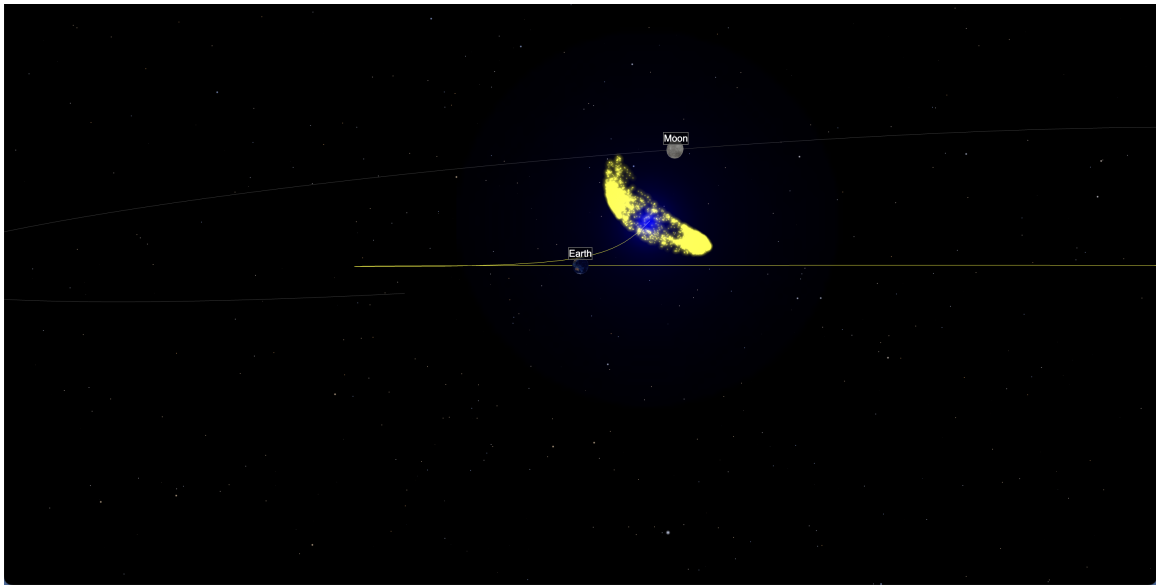


Figure 4. Near moon interaction during flyby for case 2.

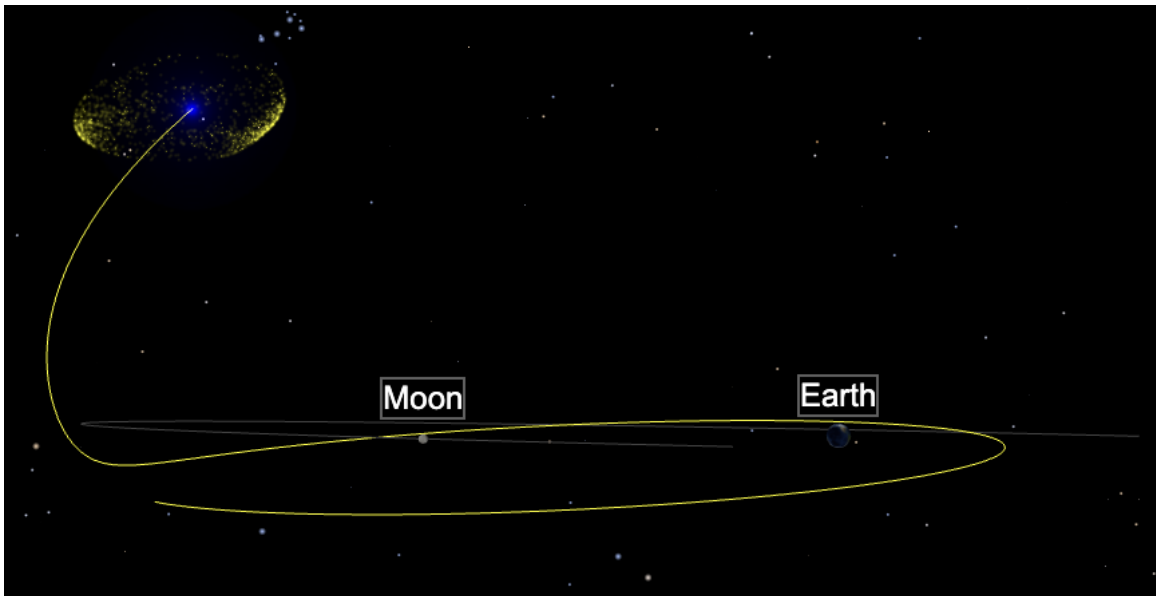


Figure 5. 3D view to highlight inclination change of case 2 post flyby.

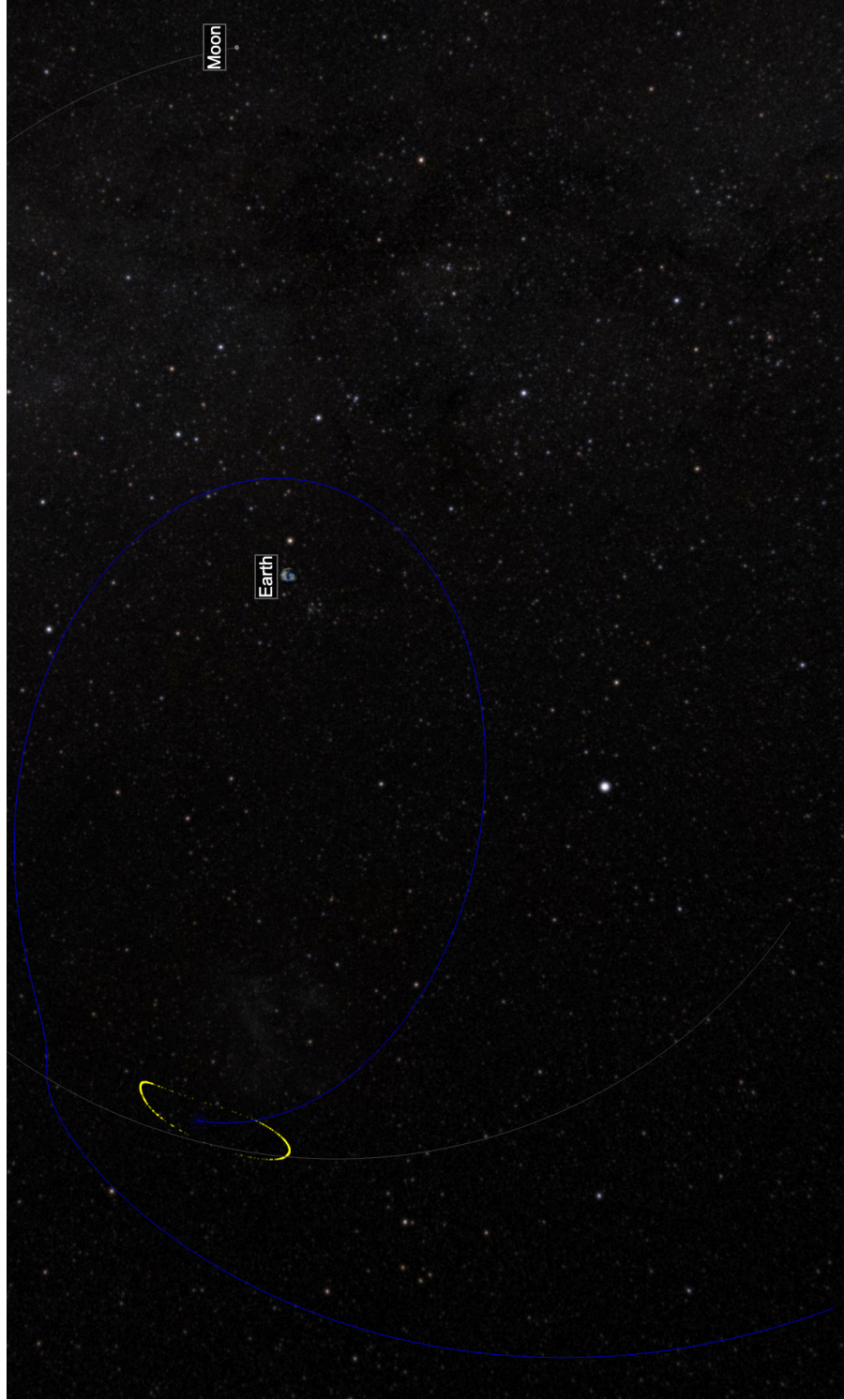


Figure 6. Earth-Sun-Moon gravitating. Controllability case that shows the initial conditions that can result in the terminal conditions for case 2. The epoch and initial conditions are 2512864.169 seconds and $x^0 = [-314562.64, -329219.31, 237676.53, 0.5262560, -0.65552567495712999, 0.1763679, 1000.0]$. Units are km, km/s, and kg.

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