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# PROGRAM KMEC: THE COMPUTATION OF KOZAI MEAN ORBITAL ELEMENTS USING A NONSINGULAR FORMULATION

BY A. D. PARKS

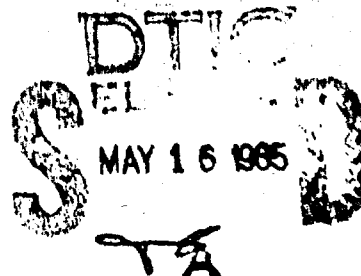
STRATEGIC SYSTEMS DEPARTMENT

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| 20. ABSTRACT (Continue on reverse side if necessary and identify by block number)<br>A geopotential expressed in terms of a nonsingular set of orbital elements is used to iteratively compute Kozai mean orbital elements from either Brouwer mean orbital elements or inertial osculating Cartesian position and velocity vectors. Although the computational procedure is general in nature, KMEC has been tailored to generate Kozai mean elements in a format that is directly applicable for use in the MX 1502-DS and modified TRANET II geocievers. <i>Additional keywords: KMEC computer program; KMEC (Kozai mean element converter); Kalman filtering</i> |                                     |                                                                                  |

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## FOREWORD

This document provides a detailed description of the formulation and computational processes contained within the Kozai Mean Element Converter (KMEC) software package. The KMEC program was developed by the Naval Surface Weapons Center under the auspices of the Defense Mapping Agency to specifically provide operational support to the new MX 1502-DS and modified TRANET II Doppler beacon satellite tracking receivers. This report has been reviewed and approved by Dr. R. J. Anderle.

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## INTRODUCTION

The primary function of the Kozai mean element converter (KMEC) is to generate mean Kozai orbital element sets for satellites tracked by the MX 1502-DS and modified TRANET II geoceivers. These geoceivers will use the mean element sets in conjunction with a Kalman filter to provide autonomous orbit updates and to predict station-satellite inview periods. To avoid problems associated with mathematical singularities, such as those that occur with near-circular orbits, the following nonsingular element set has been selected to perform the mean element transformation in KMEC:

$$\left. \begin{aligned} a &= \text{semimajor axis} \\ \lambda &= \ell + g + h \\ \xi &= e \cos \tilde{\omega} \quad (\tilde{\omega} = g + h) \\ \eta &= e \sin \tilde{\omega} \\ P &= \sin \left( \frac{i}{2} \right) \cos h \\ Q &= \sin \left( \frac{i}{2} \right) \sin h \end{aligned} \right\} \quad (1)$$

where  $a$ ,  $e$ ,  $i$ ,  $\ell$ ,  $g$ , and  $h$  are the usual Keplerian elements.

KMEC is comprised of seven basic computational functions:

1. The process flow supervisor (PFS)
2. The Cartesian input section (CIS)
3. The Brouwer input section (BIS)
4. The Walter mean element interator (WMI)
5. The nonsingular orbital element builder (NEB)
6. The geoceiver format section (GFS)
7. The Keplerian element builder (KEB)

Figure 1 shows the functional overview of KMEC and the interfunctional data flow. A detailed description of each of these functions is presented in the following sections.

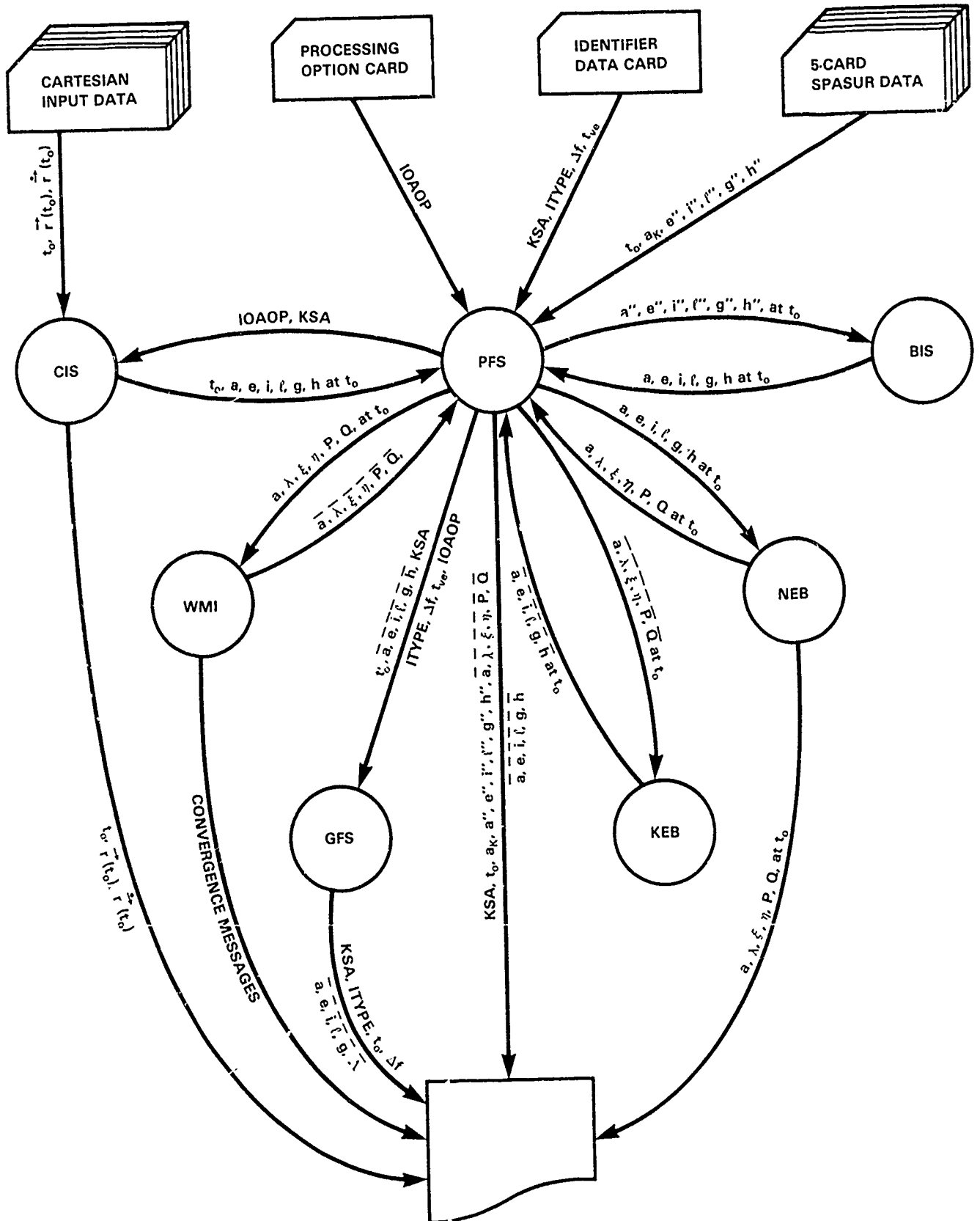


FIGURE 1. KMEC FUNCTIONAL OVERVIEW AND DATA FLOW

As user-supplied input, KMEC requires the selection of a processing option (IOAOP). If IOAOP = 0, KMEC will use a nonsingular transformation theory to convert a Brouwer mean element set obtained from Space Surveillance (SPASUR) input data to a Kozai mean element set. If IOAOP = 1, KMEC will use the same nonsingular transformation theory to convert osculating Cartesian position and velocity vectors to an associated Kozai mean element set. Also required as input are the satellite number (KSA), the satellite type (ITYPE), the satellite frequency offset ( $\Delta f$ ), and the universal time of transit of the vernal equinox ( $t_{VE}$ ). These data are included on the identifier data card. The Kozai mean elements and associated information are written to hard copy during the computational cycle.

## THE PROCESS FLOW SUPERVISOR (PFS)

### FUNCTIONAL DESCRIPTION

The principal tests performed by the PFS are to receive input data, direct the processing flow, and output computed results. Specifically, the PFS:

1. Receives from input the user-selected processing option
2. Receives from input the identifier card data
3. Receives from input Brouwer mean elements from five-card SPASUR data when IOAOP = 0
4. Directs processing through the CIS, BIS, WMI, NEB, KEB, and GFS functions
5. Converts input data to the proper computational units
6. Writes to hard copy the input, intermediate, and output mean element sets

The flow of the PFS function is presented in Figure 2.

When IOAOP = 1, the CIS function is entered and osculating Cartesian position and velocity vectors are received along with the vector epoch. These vectors are converted to an osculating Keplerian element set and are used to initiate the Kozai transformation process.

### PROCESSING EQUATIONS

The semimajor axis read from the SPASUR data is the Kaula semimajor axis  $a_k$  expressed in earth radii. This is converted in the PFS to the Brouwer mean semimajor axis  $a''$  via the transformation

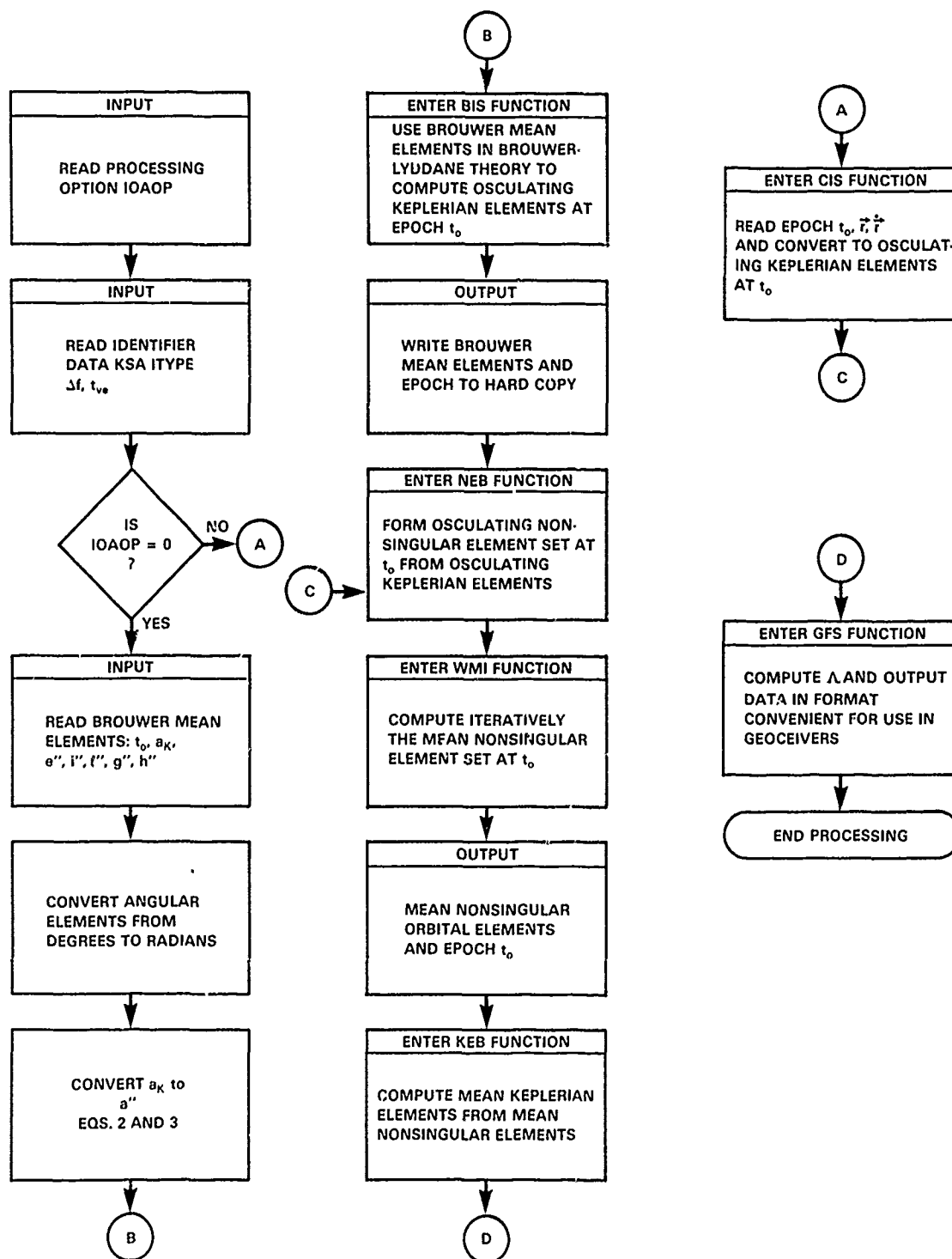


FIGURE 2. PROCESS FLOW SUPERVISOR LOGIC FLOW

$$a'' = a_k a_e \left( \frac{1 + 2X}{1 - X} \right)^{2/3} \quad (2)$$

where

$$X = \frac{3 J_2 (1 - 3/2 \sin^2 i'')}{4 a_k^2 (1 - e''^2)^{3/2}} \quad (3)$$

In the above expressions,  $a_e$  is the earth's semimajor axis,  $J_2$  is a zonal harmonic gravitational constant, and  $i''$  and  $e''$  are the Brouwer mean inclination and eccentricity, respectively.

### BROUWER INPUT SECTION (BIS)

#### FUNCTIONAL DESCRIPTION

The Brouwer input section accepts the Brouwer mean element set from the PFS and converts it into an associated osculating element set at epoch  $t_0$ . This is accomplished through the application of the Brouwer-Lyddane theory,<sup>1,2</sup> which has been modified to include the effects of atmospheric drag (the atmospheric drag decay rates are nulled during this computation). These osculating elements are then used to formulate the osculating nonsingular element set that initializes the mean element iteration algorithm. The BIS processing logic flow is shown in Figure 3.

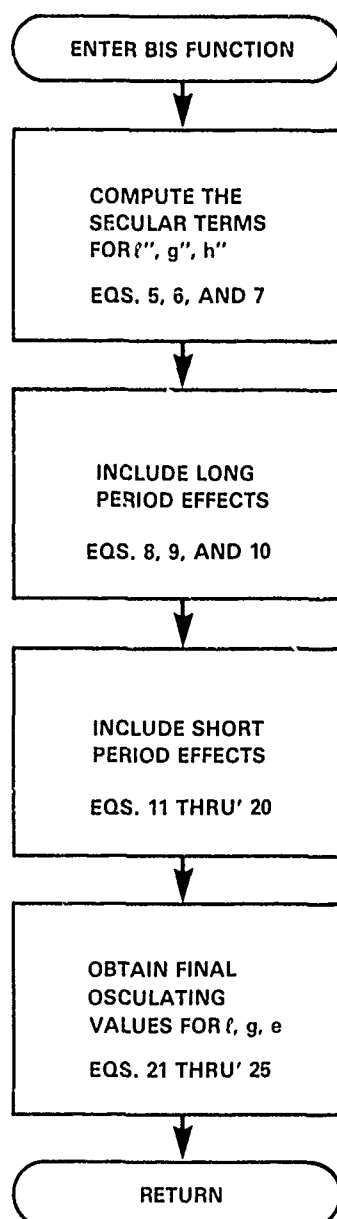


FIGURE 3. BROUWER INPUT SECTION PROCESS FLOW

## PROCESSING EQUATIONS FOR THE BROUWER-LYDDANE METHOD

The equations used to compute osculating orbital elements from mean Brouwer elements and associated decay rates are delineated in this section. First define the following:

$$\begin{aligned}
 \dot{a}'' &= \text{semimajor axis decay rate} \\
 \dot{e}'' &= \text{eccentricity decay rate} \\
 \dot{n} &= \text{time rate of change of mean motion} \\
 t &= \text{time from epoch} \\
 n_0 &= (\mu/a''^3)^{1/2} \\
 \eta &= (1 - e''^2)^{1/2} \\
 \theta &= \cos i'' \\
 \gamma_2 &= 1/2 C_{20} a_e^2/a''^2 \\
 \gamma_2' &= \gamma_2 \eta^{-4} \\
 \gamma_3' &= -C_{30} a_e^3 a''^{-3} \eta^{-6} \\
 \gamma_4' &= -3/8 C_{40} a_e^4 a''^{-4} \eta^{-8} \\
 \gamma_5' &= -C_{50} a_e^5 a''^{-5} \eta^{-10} \\
 \alpha &= 1 - 5\theta^2 \\
 \beta &= 1 - 11\theta^2 - 40\theta^4 \alpha^{-1} \\
 \gamma &= 1 - 3\theta^2 - 8\theta^4 \alpha^{-1} \\
 \delta &= 1 - 9\theta^2 - 24\theta^4 \alpha^{-1} \\
 \lambda &= 1 - 5\theta^2 - 16\theta^4 \alpha^{-1}
 \end{aligned} \tag{4}$$

where the  $C_{j0}$  ( $j = 2, 3, 4, 5$ ) are the zonal harmonic gravitational expansion coefficients. Then the secular terms are computed from

$$\begin{aligned}
 \varrho'' = n_0 t \left\{ 1 + \frac{3}{2} \gamma_2' \eta (3\theta^2 - 1) + \frac{3}{32} \gamma_2'^2 \eta \left[ -15 + 16\eta + 25\eta^2 \right. \right. \\
 \left. \left. + (30 - 96\eta - 90\eta^2)\theta^2 + (105 + 144\eta + 25\eta^2)\theta^4 \right] \right. \\
 \left. + \frac{15}{16} \gamma_4' \eta e''^2 \left[ 3 - 30\theta^2 + 35\theta^4 \right] \right\} + \varrho_0'' + \dot{n} t^2
 \end{aligned} \tag{5}$$

$$\begin{aligned}
g'' = n_0 t \left\{ -\frac{3}{2} \gamma_2' \alpha + \frac{3}{32} \gamma_2'^2 \left[ -35 + 24\eta + 25\eta^2 \right. \right. \\
\left. \left. + (90 - 192\eta - 126\eta^2)\theta^2 + (385 + 360\eta + 45\eta^2)\theta^4 \right] \right. \\
\left. + \frac{5}{16} \gamma_4' \left[ 21 - 9\eta^2 + (-270 \pm 126\eta^2)\theta^2 + (385 - 189\eta^2)\theta^4 \right] \right\} + g_0''
\end{aligned} \quad (6)$$

and

$$\begin{aligned}
h'' = n_0 t \left\{ -3\gamma_2'\theta + \frac{3}{8} \gamma_2'^2 \left[ (-5 + 12\eta + 9\eta^2)\theta + (-35 - 36\eta - 5\eta^2)\theta^3 \right] \right. \\
\left. + \frac{5}{4} \gamma_4' (5 - 3\eta^2)\theta (3 - 7\theta^2) \right\} + h_0''
\end{aligned} \quad (7)$$

The long period (dependent upon  $g''$ ) terms are computed from

$$\begin{aligned}
\delta_1 e = \frac{35}{96} \frac{\gamma_5'}{\gamma_2'} e''^2 \eta^2 \lambda \sin i'' \sin^3 g'' - \frac{1}{12} \frac{e'' \eta^2}{\gamma_2'} (3\gamma_2'^2 \beta - 10\gamma_4' \gamma) \sin^2 g'' \\
- \frac{35}{128} \frac{\gamma_5'}{\gamma_2'} e''^2 \eta^2 \lambda \sin i'' \sin g'' + \frac{1}{4} \frac{\eta^2}{\gamma_2'} \left[ \gamma_3' + \frac{5}{16} \gamma_5' (4 + 3e''^2) \delta \right] \\
\sin i'' \sin g'' + \frac{e'' \eta^2}{24\gamma_2'} [3\gamma_2'^2 \beta - 10\gamma_4' \gamma]
\end{aligned} \quad (8)$$

$$\begin{aligned}
\varrho' + g' = g'' + \varrho'' + \frac{1}{2} \left\{ \frac{1}{24\gamma_2'} \left[ -3\gamma_2'^2 \right] 2 + e''^2 - 11(2 + 3e''^2)\theta^2 \right. \\
\left. - 40(2 + 5e''^2)\theta^4 \alpha^{-1} - 400e''^2 \theta^6 \alpha^{-2} \right\} \\
+ 10\gamma_4' \left\{ 2 + e''^2 - 3(2 + 3e''^2)\theta^2 - 8(2 + 5e''^2)\theta^4 \alpha^{-1} - 80e''^2 \theta^6 \alpha^{-2} \right\}
\end{aligned}$$

$$\begin{aligned}
& + \frac{\eta^3}{\gamma_2'} \left[ \frac{\gamma_2'^2}{4} \beta - \frac{5}{6} \gamma_4' \gamma \right] \left\{ \sin 2g'' + \left\{ \frac{35}{384} \frac{\gamma_5'}{\gamma_2'} \eta^3 e'' \lambda \sin i'' \right. \right. \\
& + \frac{35}{1152} \frac{\gamma_5'}{\gamma_2'} \left[ \lambda \left\{ - e''(3 + 2e''^2) \sin i'' + \frac{e''^3 \theta^2}{\sin i''} \right\} \right. \\
& + \left. \left. 2e''^3 \theta^2 \sin i'' \right\} 5 + 32\theta^2 \alpha^{-1} + 80\theta^4 \alpha^{-2} \right] \left. \right\} \cos 3g'' \\
& + \left\{ - \frac{\gamma_3 e'' \theta^2}{4\gamma_2' \sin i''} + \frac{5}{64} \frac{\gamma_5'}{\gamma_2'} \left[ - e'' \frac{\theta^2}{\sin i''} (4 + 3e''^2) + e'' \sin i'' \right. \right. \\
& \left. \left. (26 + 9e''^2) \right] \delta - \frac{15}{32} \frac{\gamma_5'}{\gamma_2'} e'' \theta^2 \sin i'' (4 + 3e''^2) \right. \\
& \left. (3 + 16\theta^2 \alpha^{-1} + 40\theta^4 \alpha^{-2}) + \frac{1}{4} \frac{\gamma_3'}{\gamma_2'} \sin i'' \right. \\
& \left. \left( \frac{e''}{1 + \eta^3} \right) \left[ 3 - e''^2 (3 - e''^2) \right] + \frac{5}{64} \frac{\gamma_5'}{\gamma_2'} \eta^2 \delta \right. \\
& \left. \left[ \frac{e''(-32 + 81e''^4)}{4 + 3e''^2 + \eta(4 + 9e''^2)} \right] \sin i'' \right\} \cos g''
\end{aligned} \tag{9}$$

and

$$\begin{aligned}
h' = h'' + \frac{35\gamma_5' e''^3 \theta}{144\gamma_2'} & \left\{ \frac{\lambda}{2} \sin^{-1} i'' + \sin i'' \left[ 5 + 32\theta^2 \alpha^{-1} + 80\theta^4 \alpha^{-2} \right] \right\} \\
& \sin^2 g'' \cos g'' + \frac{e''^2 \theta}{12\gamma_2'} \left\{ - 3\gamma_2'^2 \left[ 11 + 80\theta^2 \alpha^{-1} + 200\theta^4 \alpha^{-2} \right] \right. \\
& + 10\gamma_4' \left[ 3 + 16\theta^2 \alpha^{-1} + 40\theta^4 \alpha^{-2} \right] \left. \right\} \sin g'' \cos g'' \\
& + \left\{ - \frac{35\gamma_5'}{576\gamma_2'} e''^3 \theta \left[ \frac{\lambda}{2} \sin^{-1} i'' + \sin i'' (5 + 32\theta^2 \alpha^{-1} + 80\theta^4 \alpha^{-2}) \right] \right. \\
& + \frac{e'' \theta}{4\gamma_2' \sin i''} \left[ \gamma_3' + \frac{5}{16} \gamma_5' (4 + 3e''^2) \delta + \frac{15}{8} \gamma_5' (4 + 3e''^2) \right. \\
& \left. \left. (3 + 16\theta^2 \alpha^{-1} + 40\theta^4 \alpha^{-2}) \sin^2 i'' \right] \right\} \cos g'' .
\end{aligned} \tag{10}$$

The short periodics (dependent upon  $E'$ ,  $f'$ ,  $\ell''$ ) are computed from:

$$a = \dot{a}''t + a'' - a'' \frac{\gamma_2}{\eta^3} (3\theta^2 - 1) + \left[ \frac{a''\gamma_2}{(1 - e'' \cos E')^3} \right] \quad (11)$$

$$[3\theta^2 - 1 + 3 \sin^2 i'' \cos(2g'' + 2f')]$$

$$\begin{aligned} e = e'' + \dot{e}''t + \delta_1 e + \frac{\eta^2 \gamma_2}{2} \left\{ \frac{3\theta^2 - 1}{\eta^6} \left[ \frac{e''}{1 + \eta^3} \left\{ 3 - e''^2 (3 - e''^2) \right\} \right. \right. \\ \left. \left. + \left\{ 3 + e'' \cos f' \cdot (3 + e'' \cos f') \right\} \cos f' \right] + \frac{3(1 - \theta^2)}{\eta^6} \right. \\ \left. \left[ e'' + \left\{ 3 + e'' \cos f' (3 + e'' \cos f') \right\} \cos f' \right] \cos (2f' + 2g'') \right\} \\ - \frac{\eta^2 \gamma_2'}{2} (1 - \theta^2) \left[ 3 \cos (2g'' + f') + \cos(2g'' + 3f') \right] \end{aligned} \quad (12)$$

$$i = i'' - \frac{e''\theta}{\eta^2 \sin i'} \delta_1 e + e'' \gamma_2' \theta \sin i'' \sin f' \sin(2f' + 2g'') \quad (13)$$

$$+ 2e'' \gamma_2' \theta \sin i'' \cdot \cos f' \cos(2f' + 2g'') + \frac{3}{2} \gamma_2' \theta \sin i'' \cos(2f' + 2g'')$$

$$\begin{aligned} g + \ell = g' + \ell' + \frac{\gamma_2'}{4} \left\{ -6\alpha(f' - \ell'' + e'' \sin f') + (3 - 5\theta^2) \right. \\ \left. \left[ 3 \sin(2f' + 2g'') + 3e'' \sin(2g'' + f') + e'' \sin(2g'' + 3f') \right] \right\} \\ + \frac{e'' \eta^2 \gamma_2'}{4(1 + \eta)} \left\{ 2(3\theta^2 - 1) (\sigma + 1) \sin f' + 3(1 - \theta^2) \right. \\ \left. \left[ (1 - \sigma) \sin (2g'' + f') + (\sigma + 1/3) \sin (2g'' + 3f') \right] \right\} \end{aligned} \quad (14)$$

$$h = h' + \left[ 2e'' \gamma_2' \theta \cos f' + \frac{3}{2} \gamma_2' \theta \right] \sin(2g'' + 2f') \quad (15)$$

$$- e'' \gamma_2' \theta \sin f' \cos(2f' + 2g'') - 3\gamma_2' \theta (f' - \ell'' + e'' \sin f')$$

and

$$\begin{aligned}
 e\delta\lambda = & \frac{1}{2} \frac{e''\eta^3}{\gamma_2'} \left\{ \frac{1}{4} \gamma_2' \beta - \frac{5}{6} \gamma_4' \gamma \right\} \sin 2g'' \\
 & - \left\{ \frac{1}{4} \frac{\gamma_3'}{\gamma_2'} \eta^3 \sin i'' + \frac{5}{64} \frac{\gamma_5'}{\gamma_2'} \eta^3 \sin i'' (4 + 9e^2) \delta \right\} . \\
 & \cos g'' + \frac{35}{384} \frac{\gamma_5'}{\gamma_2'} \eta^3 e'' \lambda \sin i'' \cos 3g'' \\
 & - \frac{1}{4} \gamma_2' \eta^3 \left\{ 2(3\theta^2 - 1) (\sigma + 1) \sin f' + 3(1 - \theta^2) (1 - \sigma) \sin (2g'' + f') \right. \\
 & \left. + (\sigma + \frac{1}{3}) \sin (2g'' + 3f') \right\}
 \end{aligned} \tag{16}$$

where

$$\sigma = \left( \frac{\eta}{1 - e'' \cos E'} \right)^2 + \left( \frac{1}{1 - e'' \cos E'} \right) \tag{17}$$

The eccentric anomaly  $E'$  is obtained from a Newton-Raphson iteration upon the Kepler equation

$$E' - e'' \sin E' = \lambda'' \tag{18}$$

and the true anomaly  $f'$  is found from

$$\sin f' = \frac{\eta \sin E'}{1 - e'' \cos E'} \tag{19}$$

$$\cos f' = \frac{\cos E' - e''}{1 - e'' \cos E'} \tag{20}$$

The final osculating values for  $a$ ,  $i$ , and  $h$  are computed from Equations 11, 13, and 15, respectively. Equations 5, 12, 14, and 16 are used to calculate final osculating values for  $\ell$ ,  $g$ , and  $e$  for the following relations:

$$A = e \cos \ell'' - e \delta \ell \sin \ell'' \quad (21)$$

$$B = e \sin \ell'' + e \delta \ell \cos \ell'' \quad (22)$$

$$\ell = \tan^{-1} (B/A) \quad (23)$$

$$g = (\ell + g) - \ell \quad (24)$$

and

$$e = (A^2 + B^2)^{1/2} \quad (25)$$

## CARTESIAN INPUT SECTION (CIS)

### FUNCTIONAL DESCRIPTION

The primary tasks performed by the CIS function are to

1. Receive from input inertial Cartesian position and velocity vectors at epoch  $t_0$ , i.e.,  $\vec{r}(t_0)$  and  $\dot{\vec{r}}(t_0)$
2. Transform the osculating inertial Cartesian components to osculating Keplerian orbital elements

The process flow of the CIS function is shown in Figure 4.

### PROCESSING EQUATIONS

The osculating inertial Cartesian vectors  $\vec{r}(t_0) = (x, y, z)$  and  $\dot{\vec{r}}(t_0) = (\dot{x}, \dot{y}, \dot{z})$  are transformed to osculating Keplerian orbital elements by using the following relationships.

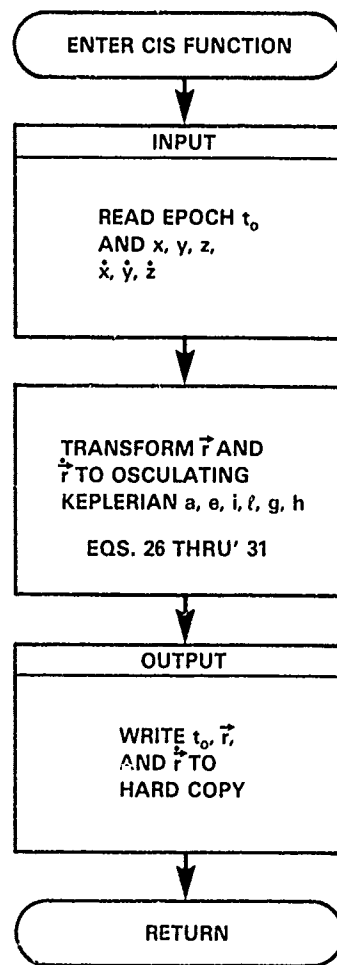


FIGURE 4. CARTESIAN INPUT SECTION PROCESS FLOW

$$a = \left[ \frac{2}{|\dot{\vec{r}}|} - \frac{|\dot{\vec{r}}|^2}{\mu} \right]^{-1} \quad (26)$$

$$e = \left\{ \frac{(\dot{\vec{r}} \cdot \dot{\vec{r}})^2}{\mu a} + \left[ \frac{|\dot{\vec{r}}|}{\mu} |\dot{\vec{r}}|^2 - 1 \right]^2 \right\}^{1/2} \quad (27)$$

$$i = \tan^{-1} \left\{ \frac{[(y\dot{z} - z\dot{y})^2 + (z\dot{x} - x\dot{z})^2]^{1/2}}{x\dot{y} - y\dot{x}} \right\} \quad (28)$$

$$\varrho = \tan^{-1} \left\{ \frac{(x\dot{y} - y\dot{x}) y}{x|x\dot{y} - y\dot{x}|} \right\} \quad (29)$$

$$h = \tan^{-1} \left\{ \frac{y\dot{z} - z\dot{y}}{x\dot{z} - z\dot{x}} \right\} \quad (30)$$

and

$$g = \tan^{-1} \left\{ \frac{z |\dot{\vec{r}} \times \dot{\vec{r}}| \left[ \frac{|\dot{\vec{r}}|}{\mu} |\dot{\vec{r}}|^2 - 1 - e^2 \right] + [x(z\dot{x} - x\dot{z}) - y(y\dot{z} - z\dot{y})] \sqrt{1 - e^2} \frac{\dot{\vec{r}} \cdot \dot{\vec{r}}}{(\mu a)^{1/2}}}{z |\dot{\vec{r}} \times \dot{\vec{r}}| \sqrt{1 - e^2} \frac{\dot{\vec{r}} \cdot \dot{\vec{r}}}{(\mu a)^{1/2}} - [x(z\dot{x} - x\dot{z}) - y(y\dot{z} - z\dot{y})] \left[ \frac{|\dot{\vec{r}}|}{\mu} |\dot{\vec{r}}|^2 - 1 - e^2 \right]} \right\} \quad (31)$$

### THE NONSINGULAR ORBITAL ELEMENT BUILDER (NEB)

The NEB function uses the osculating Keplerian elements obtained from either the BIS or CIS functions to form the osculating nonsingular element set given by Equation 1. This nonsingular element set is used to initialize the Walter mean element iterator discussed in the following section. The NEB process flow is shown in Figure 5.

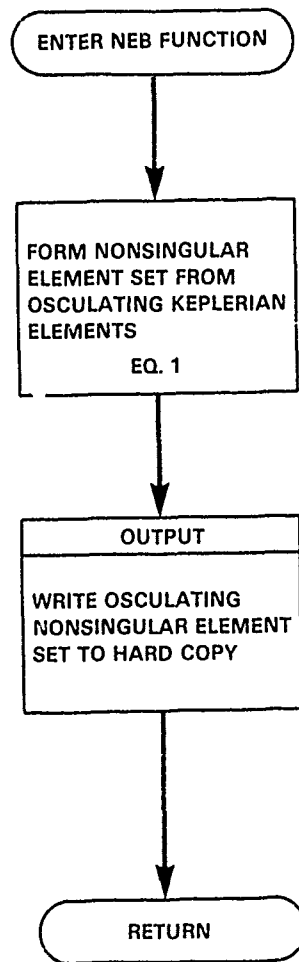


FIGURE 5. THE NEB FUNCTION PROCESS FLOW

## THE WALTER MEAN ELEMENT ITERATOR FUNCTION (WMI)

## FUNCTIONAL DESCRIPTION

The primary tasks performed by the WMI are to

1. Iteratively solve for the mean nonsingular element set associated with the osculating nonsingular element set formed by the NEB function
2. Write to hard copy a message describing the convergence status of the WMI algorithm

The mathematical computations performed by this function are relatively lengthy and quite complex. They are described in detail in the next subsection. The WMI process flow is depicted in Figure 6.

## PROCESSING EQUATIONS

The iterative technique used to find the mean nonsingular element set from the associated osculating elements is similar to that described by Walter.<sup>3</sup> This mean nonsingular element set, represented by  $\bar{\beta}_j$ , where

$$\bar{\beta}_j = \begin{cases} \bar{a} & , j = 1 \\ \bar{\lambda} & , j = 2 \\ \bar{\xi} & , j = 3 \\ \bar{\eta} & , j = 4 \\ \bar{P} & , j = 5 \\ \bar{Q} & , j = 6 \end{cases} \quad (32)$$

is obtained from the iterative process executed according to the scheme

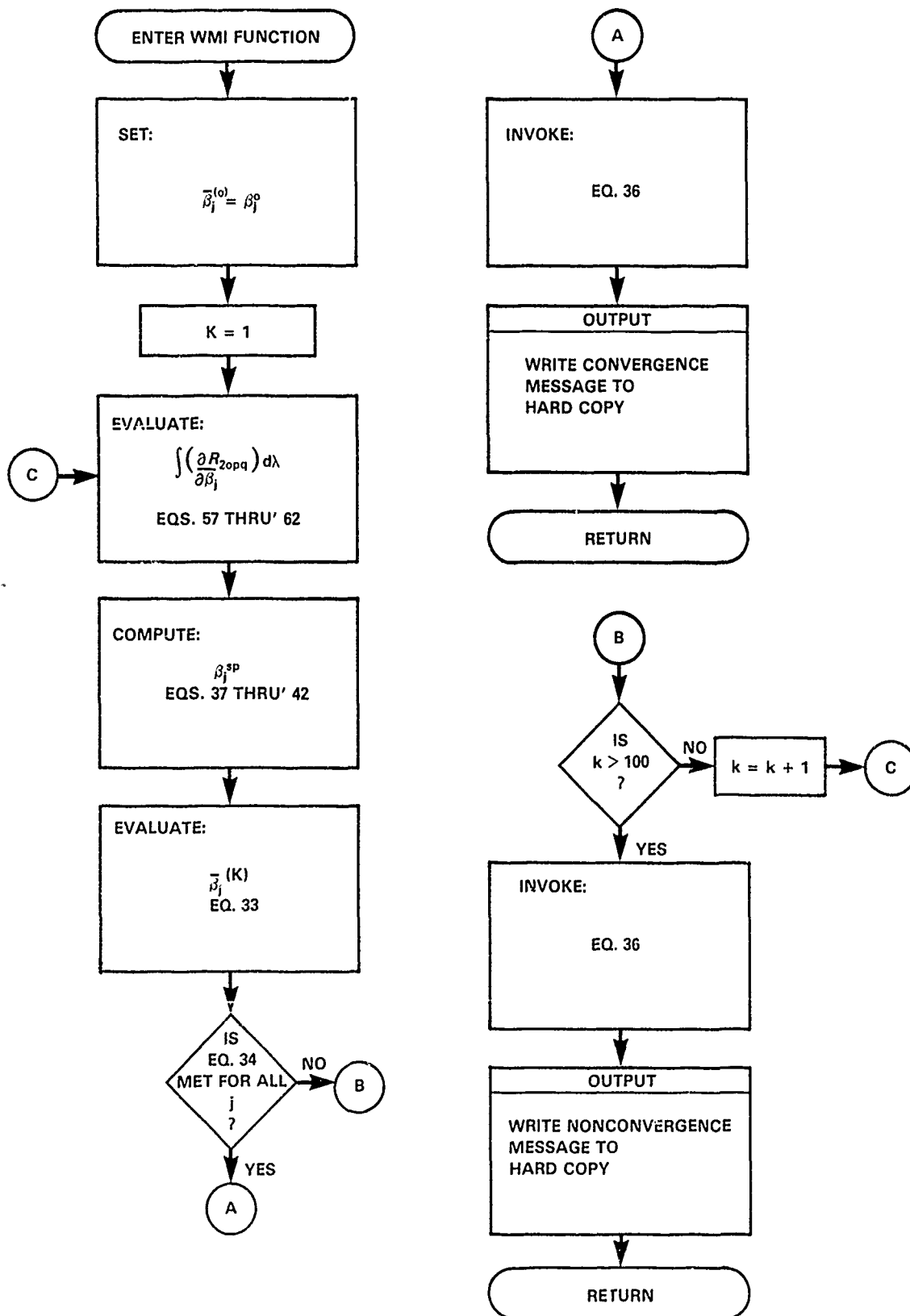


FIGURE 6. THE WMI FUNCTION PROCESS FLOW

$$\bar{\beta}_j^{(k)}(t_o) = \beta_j^o(t_o) - \beta_j^{SP} \left( \bar{\beta}_1^{(k-1)}(t_o), \dots, \bar{\beta}_6^{(k-1)}(t_o) \right), \quad (j = 1, 2, \dots, 6), \quad (33)$$

until the condition

$$\left| \bar{\beta}_j^{(k)}(t_o) - \bar{\beta}_j^{(k-1)}(t_o) \right| < \epsilon_j, \quad (j = 1, 2, \dots, 6) \quad (34)$$

is satisfied for all  $j$ . In the above expressions,  $k$  is the iteration counter;  $\beta_j^o(t_o)$  and  $\beta_j^{SP}(t_o)$  represent the osculating values and the short periodic variation of the  $j^{\text{th}}$  element at epoch  $t_o$ ; and  $\epsilon_j$  is the convergence tolerance for the  $j^{\text{th}}$  element.

To initiate this iterative process, it is assumed that

$$\bar{\beta}_j^{(o)}(t_o) = \beta_j^o(t_o), \quad (j = 1, 2, \dots, 6). \quad (35)$$

When the condition in Equation 34 is met, then

$$\bar{\beta}_j(t_o) = \bar{\beta}_j^{(k)}(t_o), \quad (j = 1, 2, \dots, 6) \quad (36)$$

As can be seen from Equation 33, short periodic variations for the nonsingular element set are required. These can be obtained by integrating the associated Lagrange planetary equations using only the  $J_2$  zonal harmonic in the gravitational disturbing function:

$$a^{SP} = J_2 \left( \frac{a_e^2}{a} \right) \left\{ \left( \frac{a}{r} \right)^3 \left[ 1 - \frac{3}{2} \sin^2 i + \frac{3}{2} \sin^2 i \cos 2(\omega + f) \right] - \left( 1 - \frac{3}{2} \sin^2 i \right) (1 - e^2)^{-3/2} \right\} \quad (37)$$

$$\begin{aligned}
\lambda^{SP} = & -\frac{2}{n^2 a} \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial a} \right) d\lambda + \frac{\gamma}{2n^2 a^2} \left[ \xi \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \xi} \right) d\lambda \right. \\
& + \eta \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \eta} \right) d\lambda \left. \right] + \frac{1}{2n^2 a^2 \gamma} \left[ P \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial P} \right) d\lambda \right. \\
& + Q \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial Q} \right) d\lambda \left. \right]
\end{aligned} \tag{38}$$

$$\begin{aligned}
\xi^{SP} = & -\frac{\gamma}{n^2 a^2 (1+\gamma)} \xi \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \lambda} \right) d\lambda - \frac{\gamma}{n^2 a^2} \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \eta} \right) d\lambda \\
& - \frac{1}{2n^2 a^2 \gamma} \eta \left[ P \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial P} \right) d\lambda + Q \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial Q} \right) d\lambda \right]
\end{aligned} \tag{39}$$

$$\begin{aligned}
\eta^{SP} = & -\frac{\gamma}{n^2 a^2 (1+\gamma)} \eta \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \lambda} \right) d\lambda + \frac{\gamma}{n^2 a^2} \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \xi} \right) d\lambda \\
& + \frac{1}{2n^2 a^2 \gamma} \xi \left[ P \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial P} \right) d\lambda + Q \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial Q} \right) d\lambda \right]
\end{aligned} \tag{40}$$

$$\begin{aligned}
P^{SP} = & -\frac{1}{2n^2 a^2 \gamma} P \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \lambda} \right) d\lambda - \frac{1}{4n^2 a^2 \gamma} \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial Q} \right) d\lambda \\
& + \frac{1}{2n^2 a^2 \gamma} P \left[ \eta \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \xi} \right) d\lambda - \xi \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \eta} \right) d\lambda \right]
\end{aligned} \tag{41}$$

and

$$\begin{aligned}
 Q^{SP} = & -\frac{1}{2n^2 a^2 \gamma} Q \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \lambda} \right) d\lambda + \frac{1}{4n^2 a^2 \gamma} \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial P} \right) d\lambda \\
 & + \frac{1}{2n^2 a^2 \gamma} Q \left[ \eta \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \xi} \right) d\lambda - \xi \sum_{pq} \int \left( \frac{\partial R_{20pq}}{\partial \eta} \right) d\lambda \right]
 \end{aligned} \tag{42}$$

In the above expressions,  $f$  is the true anomaly

$$\gamma = \sqrt{1 - e^2} \tag{43}$$

$$r = \frac{a(1 - e^2)}{1 + e \cos f} \tag{44}$$

and  $R_{20pq}$  is the  $(pq)^{th}$  contribution to the geopotential disturbing function due to the  $J_2$  zonal harmonic and is given in general by:<sup>4</sup>

$$\begin{aligned}
 R_{\ell m p q} = & \frac{\mu a_e^\ell}{a^{\ell+1}} J_{\ell m p}^{(c)} K_{\ell p q}^{(\gamma)} \left\{ \mathcal{R}_{\ell m p q} (A_{\ell m} \cos \theta_{\ell m p q} + B_{\ell m} \sin \theta_{\ell m p q}) + \right. \\
 & \left. \mathcal{H}_{\ell m p q} (A_{\ell m} \sin \theta_{\ell m p q} - B_{\ell m} \cos \theta_{\ell m p q}) \right\}
 \end{aligned} \tag{45}$$

where

$$A_{\ell m} = \begin{cases} C_{\ell m}, & \ell-m \text{ even} \\ -S_{\ell m}, & \ell-m \text{ odd} \end{cases} \tag{46}$$

$$B_{\ell m} = \begin{cases} S_{\ell m}, & \ell-m \text{ even} \\ C_{\ell m}, & \ell-m \text{ odd} \end{cases} \tag{47}$$

and

$$\theta_{\ell m p q} = (\ell - 2p + q)\lambda - m\theta \quad (48)$$

Here  $\theta$  is the Greenwich sidereal time. Note that for the case  $\ell = 2$ ,  $m = 0$

$$A_{20} = C_{20} = -J_2 \quad (49)$$

$$B_{20} = S_{20} = 0 \quad (50)$$

and

$$\theta_{20pq} = (2 - 2p + q)\lambda \quad (51)$$

The  $J_{\ell m p}$  and  $K_{\ell p q}$  functions in Equation 44 are the inclination and eccentricity functions given by

$$J_{\ell m p}^{(c)} = (-1)^k \frac{(\ell+m)!}{2^\ell p! (\ell-p)!} \sum_{j=j_1}^{j_2} (-1)^j \binom{2\ell-2p}{j} \binom{2p}{\ell-m-j} C^{2\ell-\alpha-2j} (1-C^2)^{j+(\alpha-|\alpha|)/2} \quad (52)$$

$$K_{\ell p q}(\gamma) = (-1)^{|q|} 2^\ell (1+\gamma)^{-\ell-|q|} \sum_{k=0}^{\infty} \sum_{r=0}^{|q|+k} \sum_{t=0}^k \frac{(-1)^r}{r! t!} \binom{2p-2\ell}{|q|+k-r} \binom{-2p}{k-t} \left( \frac{\ell-2p+q}{2} \right)^{r+t} \quad (53)$$

$$(1+\gamma)^{r+t-k} (1-\gamma)^k, \text{ (for } q > 0)$$

and

$$K_{\ell p q}(\gamma) = (-1)^{|q|} 2^\ell (1+\gamma)^{-\ell-|q|} \sum_{k=0}^{\infty} \sum_{r=0}^{|q|+k} \sum_{t=0}^k \frac{(-1)^t}{r! t!} \binom{-2p}{|q|+k-r} \binom{2p-2\ell}{k-t} \left( \frac{\ell-2p+q}{2} \right)^{r+t} \quad (54)$$

$$(1+\gamma)^{r+t-k} (1-\gamma)^k, \text{ (for } q < 0)$$

where

$$C = \cos\left(\frac{i}{2}\right)$$

$$k = \text{integral part of } \left\lfloor \frac{\ell-m}{2} \right\rfloor$$

$$j_1 = \max(0, -\alpha)$$

$$j_2 = \min(2\ell - 2p, \ell - m)$$

$$\alpha = m - \ell + 2p$$

The  $\mathcal{R}_{\ell m p q}$  and  $\mathcal{H}_{\ell m p q}$  functions in Equation 45 are given by

$$\mathcal{R}_{\ell m p q} = \sum_{n=0}^k \sum_{u=u_1}^{u_2} (-1)^{n+u} \delta_u \binom{|q|}{u} \binom{|\alpha|}{2n-u} \xi^{|q|-u} \eta^u p^{|\alpha|-2n+u} Q^{2n-u} \quad (55)$$

$$\mathcal{H}_{\ell m p q} = \sum_{n=0}^{k'} \sum_{u=u'_1}^{u'_2} (-1)^{n+u+1} \delta_u \binom{|q|}{u} \binom{|\alpha|}{2n+1-u} \xi^{|q|-u} \eta^u p^{|\alpha|-2n-1+u} Q^{2n+1-u} \quad (56)$$

where

$$k = \left\lfloor \frac{|q| + |\alpha|}{2} \right\rfloor, \quad k' = \left\lfloor \frac{|q| + |\alpha| - 1}{2} \right\rfloor$$

$$u_1 = \max(0, 2n - |\alpha|), \quad u_2 = \min(2n, |q|)$$

$$u'_1 = \max(0, 2n+1 - |\alpha|), \quad u'_2 = \min(2n+1, |q|)$$

$$\delta_u = 1, \text{ if } q, \alpha \text{ are both positive or negative}$$

$$\delta_u = (-1)^u, \text{ if } q \text{ or } \alpha \text{ is negative}$$

The integrals appearing in Equations 38 through 42 are given by the following expressions:

$$\int \left( \frac{\partial R_{20pq}}{\partial \lambda} \right) d\lambda = - J_2 \left( \frac{\mu a_e^2}{a^3} \right) J_{20p} K_{2pq} \left\{ R_{20pq} \cos \theta_{20pq} + \Pi_{20pq} \sin \theta_{20pq} \right\} \quad (57)$$

$$\int \left( \frac{\partial R_{20pq}}{\partial a} \right) d\lambda = \left( \frac{3}{2-2p+q} \right) J_2 \left( \frac{\mu a_e^2}{a^4} \right) J_{20p} K_{2pq} \left\{ R_{20pq} \sin \theta_{20pq} - \Pi_{20pq} \cos \theta_{20pq} \right\} \quad (58)$$

$$\int \left( \frac{\partial R_{20pq}}{\partial \xi} \right) d\lambda = - J_2 \left( \frac{\mu a_e^2}{a^3} \right) J_{20p} \left( \frac{1}{2-2p+q} \right) \left[ \left\{ \left( \frac{\partial K_{2pq}}{\partial \xi} \right) R_{20pq} + |q| K_{2pq} R_{20pq}' \right\} \right. \quad (59)$$

$$\left. \sin \theta_{20pq} - \left\{ \left( \frac{\partial K_{2pq}}{\partial \xi} \right) \Pi_{20pq} + |q| K_{2pq} \Pi_{20pq}' \right\} \cos \theta_{20pq} \right] \\ \int \left( \frac{\partial R_{20pq}}{\partial \eta} \right) d\lambda = - J_2 \left( \frac{\mu a_e^2}{a^3} \right) J_{20p} \left( \frac{1}{2-2p+q} \right) \left[ \left\{ \left( \frac{\partial K_{2pq}}{\partial \eta} \right) R_{20pq} - q K_{2pq} \Pi_{20pq}' \right\} \right. \quad (60)$$

$$\left. \sin \theta_{20pq} - \left\{ \left( \frac{\partial K_{2pq}}{\partial \eta} \right) \Pi_{20pq} + q K_{2pq} R_{20pq}' \right\} \cos \theta_{20pq} \right] \\ \int \left( \frac{\partial R_{20pq}}{\partial P} \right) d\lambda = - J_2 \left( \frac{\mu a_e^2}{a^3} \right) K_{2pq} \left( \frac{1}{2-2p+q} \right) \left[ \left\{ \left( \frac{\partial J_{20p}}{\partial P} \right) R_{20pq} \right. \right. \quad (61)$$

$$\left. + |2p-2| J_{20p} R_{2m'pq}' \right\} \sin \theta_{20pq} - \left\{ \left( \frac{\partial J_{20p}}{\partial P} \right) \Pi_{20pq} \right.$$

$$\left. + |2p-2| J_{20p} \Pi_{2m'pq}' \right\} \cos \theta_{20pq} \right]$$

and

$$\int \left( \frac{\partial R_{20pq}}{\partial Q} \right) d\lambda = - J_2 \left( \frac{\mu a_e^2}{a^3} \right) K_{2pq} \left( \frac{1}{2-2p+q} \right) \left[ \left( \frac{\partial J_{20pq}}{\partial Q} \right) R_{20pq} - (2p-2) J_{20p} \Pi_{2m'pq} \right] \quad (62)$$

$$\sin \theta_{20pq} \left[ \left( \frac{\partial J_{20p}}{\partial Q} \right) \Pi_{20pq} + (2p-2) J_{20p} R_{2m'pq} \right] \cos \theta_{20pq}$$

where use has been made of the relations

$$\left. \begin{aligned} \frac{\partial R_{\ell m p q}}{\partial \xi} &= |q| R_{\ell m p q}' \quad , & \frac{\partial \Pi_{\ell m p q}}{\partial \xi} &= |q| \Pi_{\ell m p q}' \\ \frac{\partial R_{\ell m p q}}{\partial \eta} &= -q \Pi_{\ell m p q}' \quad , & \frac{\partial \Pi_{\ell m p q}}{\partial \eta} &= q R_{\ell m p q}' \\ \frac{\partial R_{\ell m p q}}{\partial P} &= |\alpha| R_{\ell m' p q} \quad , & \frac{\partial \Pi_{\ell m p q}}{\partial P} &= |\alpha| \Pi_{\ell m' p q} \\ \frac{\partial R_{\ell m p q}}{\partial Q} &= -\alpha \Pi_{\ell m' p q} \quad , & \frac{\partial R_{\ell m p q}}{\partial Q} &= \alpha R_{\ell m' p q} \end{aligned} \right\} \quad (63)$$

and

$$q' = \begin{cases} q - 1 & (q > 0) \\ q + 1 & (q < 0) \end{cases} \quad (64)$$

$$m' = \begin{cases} -1 & (2p - 2 > 0) \\ +1 & (2p - 2 < 0) \end{cases} \quad (65)$$

The partial derivatives of the inclination and eccentricity functions are given by

$$\frac{\partial J_{\ell m p}}{\partial P} = -2P \frac{\partial J_{\ell m p}}{\partial C}, \quad (66)$$

$$\frac{\partial J_{\ell m p}}{\partial Q} = -2Q \frac{\partial J_{\ell m p}}{\partial C}, \quad (67)$$

where

$$\frac{\partial J_{\ell m p}}{\partial C} = (-1)^k \frac{(\ell+m)!}{2^\ell p! (\ell-p)!} \sum_{j=j_1}^{j_2} (-1)^j \binom{2\ell-2p}{j} \binom{2p}{\ell-m-j} \left\{ C^{2\ell-\alpha-2j-1} S^{\alpha-|\alpha|} \left[ (2\ell-|\alpha|)S^{2j} - (2j+\alpha-|\alpha|)S^{2j-2} \right] \right\} \quad (68)$$

$$S = \sin\left(\frac{i}{2}\right) \quad (69)$$

and

$$\frac{\partial K_{\ell p q}}{\partial \xi} = -\frac{\xi}{\gamma} \frac{\partial K_{\ell p q}}{\partial \gamma} \quad (70)$$

$$\frac{\partial K_{\ell p q}}{\partial \eta} = -\frac{\eta}{\gamma} \frac{\partial K_{\ell p q}}{\partial \gamma} \quad (71)$$

where

$$\begin{aligned} \frac{\partial K_{\ell p q}}{\partial \gamma} = & \frac{(-\ell-|q|)}{1+\gamma} K_{\ell p q} + (-1)^{|q|} 2^\ell (1+\gamma)^{-\ell-|q|} \sum_{k=0}^{\infty} \sum_{r=0}^{|q|+k} \sum_{t=0}^k \frac{(-1)^r}{r!t!} \\ & \left( \binom{2p-2\ell}{|q|+k-r} \binom{-2p}{k-t} \left( \frac{\ell-2p+q}{2} \right)^{r+t} (1+\gamma)^{r+t-k-1} [(r+t-k)(1-\gamma)^k - \right. \\ & \left. k(1+\gamma)(1-\gamma)^{k-1}] \right), \quad (\text{for } q > 0) \end{aligned} \quad (72)$$

$$\frac{\partial K_{\ell p q}}{\partial \gamma} = \frac{(-\ell - |q|)}{1 + \gamma} K_{\ell p q} + (-1)^{|q|} 2^{\ell} (1 + \gamma)^{-\ell - |q|} \sum_{k=0}^{\infty} \sum_{r=0}^{|q|+k} \sum_{t=0}^k \frac{(-1)^t}{r!t!} \cdot$$

$$\binom{-2p}{|q|+k-r} \binom{2p-2\ell}{k-t} \left( \frac{\ell-2p+q}{2} \right)^{r+t} (1+\gamma)^{r+t-k-1} [(r+t-k)(1-\gamma)^k -$$

$$k(1+\gamma)(1-\gamma)^{k-1}], \text{ (for } q < 0)$$

and

$$\frac{\partial K_{\ell p(2p-\ell)}}{\partial \gamma} = \frac{-2\ell+1}{\gamma} K_{\ell p(2p-\ell)} - 2\gamma^{-2\ell+2} \sum_{k=0}^{p'-1} \binom{\ell-1}{2k+|2p-\ell|} \cdot$$

$$\binom{2k+|2p-\ell|}{k} 2^{-2k-|2p-\ell|} k(1-\gamma^2)^{k-1}, \text{ (for } q = 2p - \ell \text{ and } p' = \frac{\ell - |2p-\ell|}{2} )$$

It should be mentioned that even if convergence is not achieved (Equation 34), KMEC assumes that the final mean element values are correct and continues processing with them. This is done since near convergence may have occurred and the resulting mean elements may still be usable. Messages concerning the state of nonconvergence are generated to alert the user.

## THE KEPLERIAN MEAN ELEMENT BUILDER (KEB)

### FUNCTIONAL DESCRIPTION

The KEB function decomposes the mean nonsingular element set obtained from the WMI function into a mean Keplerian element set. The KEB process flow is shown in Figure 7.

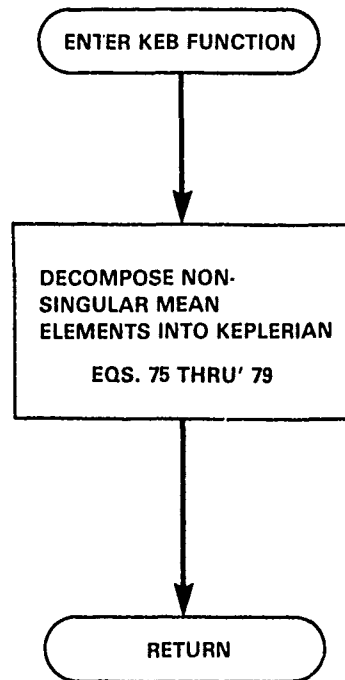


FIGURE 7. THE KEB FUNCTION PROCESS FLOW

## PROCESSING EQUATIONS

The decomposition of the mean nonsingular element set into the mean Keplerian element set is accomplished through application of the following relations:

$$\bar{e} = (\bar{\xi}^2 + \bar{\eta}^2)^{1/2} \quad (75)$$

$$\bar{h} = \tan^{-1} (\bar{Q}/\bar{P}) \quad (76)$$

$$\bar{g} = \tan^{-1} (\bar{\eta}/\bar{\xi}) - \bar{h} \quad (77)$$

$$\bar{l} = \bar{\lambda} - (\bar{g} + \bar{h}) \quad (78)$$

and

$$\bar{i} = 2 \sin^{-1} [(\bar{P}^2 + \bar{Q}^2)^{1/2}] \quad (79)$$

Of course, no decomposition of the mean semimajor axis  $\bar{a}$  is needed.

## THE GEOCEIVER FORMAT SECTION (GFS)

## FUNCTIONAL DESCRIPTION

The GFS function assembles the satellite ID, type, and mean Keplerian orbital elements; computes an earth-fixed longitude at epoch for the mean right ascension of the ascending node; and converts the epoch from modified Julian days to year, day, and minutes of day (GMT). These data are written to hard copy. The GFS process flow is illustrated in Figure 8.

## PROCESSING EQUATIONS

The time of day in minutes (GMT) is computed by using the following:

$$t_{MIN} = \frac{(t_o - 367y + (7y/4) - d + 678957.) 86400.}{60.} \quad (80)$$

where  $t_o$  is the epoch in modified Julian days,  $y$  is the year expressed as an integer, and  $d$  is the day of year. The right ascension of the Greenwich meridian  $L$  at epoch  $t_o$  is computed from

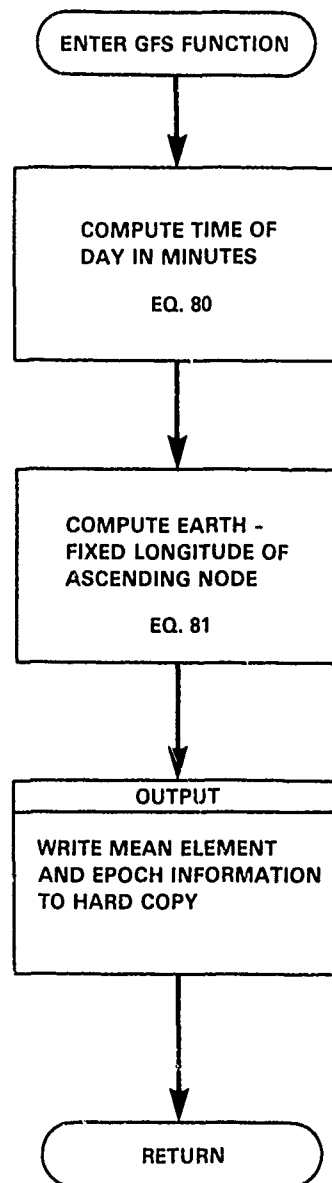


FIGURE 8. THE GFS FUNCTION PROCESS FLOW

$$L = \omega_{\oplus} (t_o - t_{ve}) \quad (81)$$

where  $\omega_{\oplus}$  is the rotation rate of the earth and  $t_{ve}$  is the time of transit of the vernal equinox expressed in modified Julian days. The earth-fixed longitude  $\Lambda$  of the ascending node of the orbit is then computed by using

$$\Lambda = h - L \quad (82)$$

### REFERENCES

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APPENDIX

COMPUTER LISTING OF PROGRAM KMEC

09.44.26 03/04/83

PROGRAM KMEC ( INPUT,OUTPUT,TAPE6=OUTPUT)

THIS PROGRAM IS THE KOZAI MEAN ELEMENT CONVERTER (KMEC) AND OPTIONALLY  
 CONVERTS BROUWER MEAN ELEMENTS OR OSCULATING CARTESIAN VECTORS TO  
 KOZAI MEAN ELEMENTS USING A NONSINGULAR ELEMENT FORMULATION. A  
 DETAILED DESCRIPTION OF KMEC CAN BE FOUND IN THE DOCUMENT ENTITLED  
 " PROGRAM KMEC = THE COMPUTATION OF KOZAI MEAN ORBITAL ELEMENTS  
 USING A NONSINGULAR FORMULATION " BY A. D. PARKS. ALL THE  
 EQUATION NUMBERS MENTIONED IN THIS PROGRAM REFER TO THOSE IN THIS  
 DOCUMENT .

## INPUT GUIDE .....

CARD 1 ---- PROCESSING OPTION

ICAOP = 0 CONVERT BROUWER MEAN ELEMENTS

ICAOP = 1 CONVERT CARTESIAN VECTORS

CARD 2---- IDENTIFIER DATA

KSA = SATELLITE NUMBER

ITYPE = SATELLITE TYPE

FT = FREQUENCY OFFSET (FPM)

TVE = VERNAL EQUINOX TRANSIT TIME (MJD)

IF ICAOP = 0 , THEN CARD 3 THROUGH CARD 7 ARE THE FIVE CARD SPASUR  
 DATA.

IF ICAOP = 1, THEN----

CARD 3---- VECTOR EPOCH

IAYR = YEAR

ADJDA = DAY OF YEAR

ADJSE = SECONDS OF DAY

CARD 4---- POSITION VECTOR COMPONENTS

RV(1) = X

RV(2) = Y

RV(3) = Z

CARD 5---- VELOCITY VECTOR COMPONENTS

RV(4) = XDOT

RV(5) = YDOT

RV(6) = ZDOT

COMMON / CON / DEGRAD, XMU, XJ2, AE

COMMON / KORBEL / BA, ES, XI, W, O, AM

COMMON / HNEL / XY

COMMON / NORBEL / A, XL, Z, XN, P, Q

DIMENSION XM(6)

DATA XMU / 398600.8 / , XJ2 / 1082.6E-06 / , AE / 6378.135 /

DATA R2 / 541.15E-06 /

DATA B0,B2,B3,B4,B5 / 398600.8, -.1755528999E+11, .26386647738E+12,

1 .1063073996E+15, .305605022E+18 /

DATA DT / 0.0 / , CN2 / 0.0 / , A1 / 0.0 / , E1 / 0.0 / , RN1/0.0 /

THIS IS THE PFS FUNCTION

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PI = 3.14159265359
DEGRAD = PI / 180.
READ *, IOAOP
READ *, KSA, IITYPE, FT, TVE

C
C
C  ENTER CIS FUNCTION

IF( IOAOP.NE.0 ) CALL ORBADJ( IOAOP,KSA,UJO, IYLD,IDAY,SEC)
IF ( IOAOP.NE.0) GO TO 40
READ 4, KSAP, IYLD,IDAY
4  FORMAT(2X,I5,40X,I1,I3)
READ 5 , UJD, ETU, H0, G0, BES, BI
5  FORMAT (8X,F14.8,5(1X,F8.4) )
READ 10, A0
10 FORMAT (1/8X,F11.5)
BXI = BI*DEGRAD
BW = G0*DEGRAD
BO = H0*DEGRAD
BAM = ETU*DEGRAD

C
C
C  SEE EQS. (2) - (3)

FN = 1. / ( 1. - ES*ES ) ** 1.5
TA = SIN( BXI )
TB = TA * TA
TE = 1. - ( 3.*TB)/2.
TAD = (( 3. * R2)/(2.*A0*A0))*TE*FN
BA = ( A0*((1. + 2.*TAD) / ( 1. - TAD )) ** 0.6666666667) * AE
SEC = 0.00
PRINT 15
15 FORMAT (1H1 )

C
C
C  ENTER BIS FUNCTION

30 CALL BRAUER (80,82,83,84,85,DT,8A,8ES,8XI,8AM,8W,30,CN2,A,ES,XI,
1  AM,H,0,A1,E1,RN1)
PRINT 34
34 FORMAT(1/,56X,*BROUWER*)
PRINT 35, KSA,UJD,A0,8A,8ES,8I,8XI,ETU,8AM,G0,8W,H0,8O
35 FORMAT (39X,*MEAN ORBITAL ELEMENTS FOR SATELLITE *,I5,*1*
1//9X,*EPOCH (JULIAN DAY MINUS 2,400,000.5) *,E22.14/9X,
2*SEMI-MAJOR AXIS*,24X,E22.14,* EARTH RADII *,E22.14,
3* KILOMETERS*/9X,*ECCENTRICITY*,26X,E22.14/9X,
4*INCLINATION*,27X,E22.14,* DEGREES*,6X,E22.14,* RADIANS*/9X,
5*MEAN ANOMALY*,26X,E22.14,* DEGREES*,6X,E22.14,* RADIANS*/9X,
6*ARGUMENT OF PERIGEE*,19X,E22.14,* DEGREES*,6X,E22.14,
7* RADIANS*/9X,*RIGHT ASCENSION OF THE ASCENDING NODE *,E22.14,
8* DEGREES*,6X,E22.14,* RADIANS*/1)
BA = A

C
C
C  ENTER NEB FUNCTION

40 CALL FORM

C
C  ENTER WHI FUNCTION

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C
CALL MEAN
XXXL = XM(2) / DEGRAD
PRINT 15
PRINT 50, KSA,UJD,XM(1),XXXL,XM(3),XM(4),XM(5),XM(6)
50 FORMAT(28X,*MEAN NONSINGULAR ORBITAL ELEMENTS FOR SATELLITE *,I5,
1* - *,//31X,*EPOCH (JULIAN DAY MINUS 2,400,000.5 ) *,E22.14/31X,
2*SEMI MAJOR AXIS*,24X,E22.14,* KILOMETERS*/31X,*LAMBDA*,32X,E22.14,
21X,
3*DEGREES*/31X,*ZETA*,34X,E22.14/31X,*ETA*,35X,E22.14/31X,*P*,37X,
4 E22.14/31X,*Q*,37X,E22.14/)

C
C
C ENTER KEB FUNCTION

CALL DCMPOS
XA = BA / AE
XXI = XI / DEGRAD
XW = W / DEGRAD
XO = O / DEGRAD
XAM = AM / DEGRAD
PRINT 60
60 FORMAT (///,57X,*KOZAI*)
PRINT 35, KSA,UJD,XA,BA,ES,XXI,XI,XAM,AM,XW,W,XO,O

C
C
C ENTER GFS FUNCTION

CALL MX15(2(KSA,ITYPE,FT,IYLD,IDAY,BA,XXI,ES,XW,XO,XAM,UJD,TVE,
1 SEC,IOAOP)
STOP
END
SUBROUTINE FORM

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C
C
C THIS IS THE NEB FUNCTION.
C OSCULATING NONSINGULAR ELEMENTS ARE FORMED FROM THE OSCULATING
C KEPLERIAN ELEMENTS. SEE EQS. (1).
C
C
C
COMMON / KORBEL / BA, ES, XI, W, O, AM
COMMON / NORBEL / A, XL, Z, XN, P, Q
PI = 3.14159265359
PI2 = 2.*PI
A = BA
XL = W + O + AM
XL = AMOD( XL,PI2 )
WB = W + Q
WB = AMOD( WB,PI2 )
Z = ES * COS( WB )
XN = ES * SIN( WB )
P = SIN(0.5*XI) * COS( O )
Q = SIN(0.5*XI)*SIN( O )
PRINT 9
9 FORMAT(1/,46X,*OSCULATING NONSINGULAR ELEMENT SET--*)
PRINT 10, A, XL, Z, XN, P, Q
10 FORMAT(1/,50X,*A =*,G16.11,*KM*,/50X,

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1      *LAMBDA=*,G16.10,*RAD*,/50X,
2      *ZETA  =*,G16.10,/50X,
3      *ETA   =*,G16.10,/50X,
4      *P     =*,G16.10,/50X,
5      *Q     =*,G16.10)

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RETURN
END
SUBROUTINE MEAN

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C

THIS IS THE MMI FUNCTION. MEAN NONSINGULAR ELEMENTS ARE OBTAINED  
USING THE WALTER ALGORITHM.

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DIMENSION XOSC(6),XM(6),TOL(6),DELX(6),XSP(6),XMN(6)
DIMENSION SUM(6)
COMMON / NORBEL / A,XL,Z,XN,P,Q
COMMON / MNEL / X1
COMMON / INTG / SUM, G, XMOT
DATA TOL / 0.01,5*0.00001 /
XOSC(1) = A
XOSC(2) = XL
XOSC(3) = Z
XOSC(4) = XN
XOSC(5) = P
XOSC(6) = Q
KOUNT = 0

```

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C

INITIALIZE PROCESS. SEE EQ. (35) .

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DO 10 I = 1,6
10 XM(I) = XOSC(I)

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C

EVALUATE SUMS OF INTEGRALS OF GEOPOTENTIAL DISTURBING FUNCTION  
PARTIAL DERIVATIVES. SEE EQS. (57)-(74),(37)-(42).

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20 CALL EVI

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LOAD SHORT PERIODIC ARRAY.

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XSP( 1 ) = XSPA ( XM )
XSP(2) = XSPL( XM, SUM, G, XMOT )
XSP(3) = XSPZ( XM, SUM, G, XMOT )
XSP(4) = XSPXN( X1, SUM, G, XMOT )
XSP(5) = XSFP( XM, SUM, G, XMOT )
XSP(6) = XSPQ( XM, SUM, G, XMOT )
KOUNT = KOUNT + 1

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C

ITERATE FOR MEAN ELEMENTS. SEE EQS. (33)-(34).

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DO 50 J = 1, 6
XMN(J) = XOSC(J) - XSP(J)
DELX(J) = ABS( XMN(J) - XM(J) )
IF ( DELX(J) .LE. TOL(J) ) GO TO 50
40 XM(J) = XMN(J)

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50 CONTINUE
   DO 50 K = 1,6
   IF(DELX(K).GT.TOL(K).AND.KOUNT.LT.100) GO TO 20
60 CONTINUE
70 IF(KOUNT.LT.100) GO TO 90
   WRITE(6,100) KOUNT
100 FORMAT(/,24X,*THE KOZAI MEAN ELEMENT CONVERSION ALGORITHM DID NOT
   1 CONVERGE IN *,I5,* ITERATIONS.*)
   PRINT 101, (TOL(I),I=1,6),(DELX(I),I=1,6)
101 FORMAT(/,5X,*THE DESIRED TOLERANCES WERE--,/,6(2X,E12.4)/5X,
   2 *THE FINAL TOLERANCES WERE--,/,6(2X,E12.4))
   RETURN
90 CONTINUE
   PRINT 91, KOUNT
91 FORMAT(/,30X,*THE KOZAI MEAN ELEMENT CONVERSION ALGORITHM CONVERGE
   10 IN *,I5,* ITERATIONS.*)
   RETURN
   END
   SUBROUTINE EVI

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C

THIS IS PART OF THE MMI FUNCTION. SUMMATIONS OVER P AND Q OF  
THE INTEGRALS OF THE GEOPOTENTIAL DISTURBING FUNCTION PARTIALS ARE  
EVALUATED. SEE EQS. (6.26) - (57)-(74),(37)-(42).

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   DIMENSION XM(6), SUM(6)
   COMMON / INTG / SUM, G, XMCT
   COMMON / CON / DEGRAD, XMU, XJ2, AE
   COMMON / MNEL / XM
   COMMON / KORBEL / AA,ES,XI,W,OM,AM
   XMCT = SQRT( XMU/ ( XM(1)**3))
   L = 2
   M = 0
   DO 10 I = 1, 6
   SUM(I) = 0.0
10 CONTINUE
   CALL DCMPDS
   AA = XM(1)
   G = 1. - ES*ES
   G = SQRT(G)
   A = XM(3)
   B = XM(4)
   C = XM(5)
   D = XM(6)
   CM = -XJ2*(XMU*AE*AE/(AA**3))
   CM1 = -CM/AA
   DO 20 I = 1, 3
   IP = I - 1
   IQT = 2*IP - 2
   IAL = M - L + 2*IP
   DO 30 J = 1, 5
   IQ = J - 3
   IF(IQ.EQ.IQT) GO TO 30
   THT = (L-2*IP+IQ)*XM(2)

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IQP = 0
IF (IQ.GT.0) IQP = IQ - 1
IF ( IQ.LT. 0 ) IQP = IQ + 1
MP = 0
IF ( IAL .LT. 0 ) MP = M + 1
IF ( IAL . GT. 0 ) MP = M - 1
CJ = XJLMP(L,M,IP,XI)
CK = XKLPQ(L,IP,IQ,G)
DKDZ = - ( XM(3)/G)*OKLPQ(L, IP, IQ, G)
DKDN = - ( XM(4)/G)*OKLPQ(L, IP, IQ, G)
DJDP = -2.*XM(5)*DJLMP(L,M,IP,XI)
DJDQ = -2.*XM(6)*DJLMP(L,M,IP,XI)
R = RLMPQ(L,M,IP,IQ,A,B,C,D)
RQP = RLMPQ(L,M,IP,IQP,A,B,C,D)
RMP = RLMPQ(L,MP,IP,IQ,A,B,C,D)
BI = BILMPQ(L,M,IP,IQ,A,B,C,D)
BIQP = BILMPQ(L,M,IP,IQP,A,B,C,D)
BIMP = BILMPQ(L,MP,IP,IQ,A,B,C,D)
CT = COS(THT)
ST = SIN(THT)
SUM(1) = SUM(1) + CH*CJ*CK*(R*CT+BI*ST)
SUM(2) = SUM(2) + (3./(2.-2.*IP+IQ))*CH*1*CJ*CK*(R*ST-BI*CT)
SUM(3) = SUM(3) + (1./(2.-2.*IP+IQ))*CH*CJ*((DKDZ*R+IABS(IQ)*CK
1 *RQP)*ST-(DKDZ*BI+IABS(IQ)*CK*BIQP)*CT)
SUM(4) = SUM(4) + (1./(2.-2.*IP+IQ))*CH*CJ*((DKDN*R-IQ*CK*BIQP)*ST
1 - (DKDN*BI +IQ*CK*RQP)*CT)
SUM(5) = SUM(5)+(1./(2.-2.*IP+IQ))*CH*CK*((DJDP*R+IABS(2*IP-2)*
1 CJ*RMP)*ST - (DJDP*BI+IABS(2*IP-2)*CJ*BIMP)*CT)
SUM(6) = SUM(6) + (1./(2.-2.*IP+IQ))*CH*CK*((DJDQ*R-(2.*IP-2)*
1 CJ*BIMP)*ST-(DJDQ*BI+(2*IP-2)*CJ*RMP)*CT)
30 CONTINUE
20 CONTINUE
RETURN
END
SUBROUTINE DCMPOS

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C  
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C

THIS IS THE KEB FUNCTION. MEAN NONSINGULAR ELEMENTS ARE  
DECOMPOSED INTO MEAN KEPLERIAN ELEMENTS. SEE EQS. (75) - (79).

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DIMENSION XM(6)
COMMON / MNEL / X4
COMMON / KORBEL / A,ES,XI,W,OM,AM
PI = 3.14159265359
PI2 = 2.*PI
ES = SQRT(XM(3)*XM(3) + XM(4)*XM(4))
OM = ARCTAN(XM(6),XM(5))
WB = ARCTAN(XM(4),XM(3))
W = WB - OM
IF ( W.LT. 0.0) W = PI2 + W
XI = 2.*ASIN(SQRT(XM(5)*XM(5)+XM(6)*XM(6)))
A = XM(1)
AM = XM(2) - WB
IF ( AM.LT. 0.0) AM = PI2 + AM

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RETURN  
END  
SUBROUTINE BRAUER(B0,B2,B3,B4,B5,DT,A2P,E2P,CI2P,CL02P,G02P,H02P,  
1CN2,A,CE,CI,CL,G,H,ADT,RND,ESD)

C  
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C

THIS IS THE BIS FUNCTION. A DRAG AUGMENTED BROUWER - LYGDANE  
THEORY IS USED TO GENERATE OSCULATING KEPLERIAN ELEMENTS FROM  
MEAN BROUWER ELEMENTS. SEE EQS. (4)-(25).

|                                                                   |          |
|-------------------------------------------------------------------|----------|
| 1004 A2P2=A2P**2                                                  | BRLY0050 |
| A2P4=A2P2**2                                                      | BRLY0060 |
| CN0=SQRT(B0/(A2P2*A2P))                                           | BRLY0070 |
| E2P2=E2P**2                                                       | BRLY0080 |
| ETA=SQRT(1.-E2P2)                                                 | BRLY0090 |
| SINEI=SIN(CI2P)                                                   | BRLY0100 |
| THETA=COS(CI2P)                                                   | BRLY0110 |
| THETA2=THETA**2                                                   | BRLY0120 |
| THETA4=THETA2**2                                                  | BRLY0130 |
| THETA6=THETA4*THETA2                                              | BRLY0140 |
| CJ2=-B2/(2.*B0*A2P2)                                              | BRLY0150 |
| ETA2=ETA**2                                                       | BRLY0160 |
| ETA3=ETA2*ETA                                                     | BRLY0170 |
| ETA4=ETA2**2                                                      | BRLY0180 |
| CJ21P=CJ2/ETA4                                                    | BRLY0190 |
| CJ31P=B3/(B0*A2P2*A2P*ETA4*ETA2)                                  | BRLY0200 |
| CJ41P=(3.*B4)/(8.*B0*A2P4*ETA4*ETA4)                              | BRLY0210 |
| CJ51P=B5/(B0*A2P4*A2P*ETA4**2*ETA2)                               | BRLY0220 |
| FUN1=3.*THETA2-1.                                                 | BRLY0230 |
| FUN2=1.-5.*THETA2                                                 | BRLY0240 |
| SINEI2=SINEI**2                                                   | BRLY0250 |
| A1=A2P*CJ2*FUN1                                                   | BRLY0260 |
| A0=-A1/ETA3                                                       | BRLY0270 |
| A2=3.*A2P*CJ2*SINEI2                                              | BRLY0280 |
| FUN5=1.-11.*THETA2-(48.*THETA4)/FUN2                              | BRLY0290 |
| FUN6=-FUN1-(8.*THETA4)/FUN2                                       | BRLY0300 |
| FUN4=THETA2/SINEI2                                                | BRLY0310 |
| FUN22=FUN2**2                                                     | BRLY0320 |
| CJ21F2=CJ21P**2                                                   | BRLY0330 |
| E01P=((E2P*ETA2)*(3.*CJ21P2*FUN5-10.*                             | BRLY0340 |
| 1CJ41P*FUN6))/(24.*CJ21P)                                         | BRLY0350 |
| E21P=-2.*E01P                                                     | BRLY0360 |
| E31P=((35.*CJ51P*E2P2*ETA2*SINEI)*(FUN2-(16.*THETA4)/FUN2))/(96.* | BRLY0370 |
| 1CJ21P)                                                           | BRLY0380 |
| E11P=-.75*E31P+((.25*ETA2*SINEI)*(CJ31P+.3125*CJ51P*(4.*3.*E2P2)* | BRLY0390 |
| 1(1.-9.*THETA2-(24.*THETA4)/FUN2))/CJ21P                          | BRLY0400 |
| CI0=-(E2P*THETA)/(ETA2*SINEI)                                     | BRLY0410 |
| CI2=CJ21P*THETA*SINEI*1.5                                         | BRLY0420 |
| CI1=E2P*CI2*.66666667                                             | BRLY0430 |
| FUN7=(-.5*ETA3*CJ21P)/E2P                                         | BRLY0440 |
| CL21P=(ETA3/CJ21P)*(1.25*CJ21P2*FUN5-.63333333*CJ41P*FUN6)        | BRLY0450 |
| CL12P=CN0*(1.+1.5*CJ21P*ETA*FUN1+.09375*CJ21P2*ETA*(-15.+16.*ETA+ | BRLY0460 |
| 125.*ETA2+(30.-96.*ETA-90.*ETA2)*THETA2+(105.+144.*ETA+25.*ETA2)* | BRLY0470 |
| 2THETA4)+.9375*CJ41P*ETA*E2P2*(3.-30.*THETA2+35.*THETA4))         | BRLY0480 |

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CL22P=.5*CN0*CN2
G21P=(1./((24.*CJ21P)))*(-3.*CJ21P2*(2.+E2P2-11.*(2.+3.*E2P2)*THETA2BRLY0500
1-40.*(2.+5.*E2P2)*THETA4/FUN2-400.*E2P2*THETA6/FUN22)+10.*CJ41P* BRLY0510
2(2. BRLY0520
3+E2P2-3.*(2.+3.*E2P2)*THETA2-8.*(2.+5.*E2P2)*THETA4/FUN2-80.*E2P2*BRLY0530
4THETA6/FUN22)) BRLY0540
G12P=CN0*(-1.5*CJ21P*FUN2+.09375*CJ21P2*(-35.+24.*ETA+25.*ETA2+ BRLY0550
1(90.-192.*ETA-126.*ETA2)*THETA2+(385.+360.*ETA+45.*ETA2)*THETA4)+ BRLY0560
2.3125*CJ41P*(21.-9.*ETA2+(-270.+126.*ETA2)*THETA2+(385.-185.*ETA2)*BRLY0570
3*THETA4)) BRLY0580
H2=1.5*CJ21P*THETA BRLY0590
H3=-2.*H2 BRLY0600
H1=.66666667*E2P*H2 BRLY0610
H31P=((35.*CJ51P*E2P2*E2P*THETA)/(144.*CJ21P))*(.5/SINEI*(FUN2-( BRLY0620
116.*THETA4)/FUN2)+SINEI*(5.+(32.*THETA2)/FUN2+80.*THETA4/FUN22)) BRLY0630
H11P=-.25* BRLY0640
1 H31P*((.25*E2P*THETA)/(CJ21P*SINEI))*(CJ31P+.3125*CJ51P* BRLY0650
2(4.+3.*E2P2)*(1.-9.*THETA2-(24.*THETA4)/FUN2)+1.875*CJ51P*SINEI2* BRLY0660
3(4.+3.*E2P2)*(3.+(16.*THETA2)/FUN2+(40.*THETA4)/FUN22)) BRLY0670
H21P=(E2P2*THETA)/(12.*CJ21P)*(-3.*CJ21P2*(11.+(80.*THETA2)/FUN2+ BRLY0680
1(200.*THETA4)/FUN22)+10.*CJ41P*(3.+(16.*THETA2)/FUN2+(40.*THETA4)/BRLY0690
2FUN22)) BRLY0700
H12P=CN0*THETA*(-3.*CJ21P+.375*CJ21P2*(-5.+12.*ETA+9.*ETA2+(-35.- BRLY0710
136.*ETA-5.*ETA2)*THETA2)+1.25*CJ41P*(5.-3.*ETA2)*(3.-7.*THETA2)) BRLY0720
AID=CJ51P/CJ21P BRLY0730
AID2=FUN2-(16.*THETA4)/FUN2 BRLY0740
C1=35./384.*AID*ETA3*E2P*SINEI*AID2 BRLY0750
AID3=THETA2/SINEI BRLY0760
AID4=THETA2*SINEI BRLY0770
E2P3=E2P2*E2P BRLY0780
C2=35./1152.*AID*((-E2P*SINEI*(3.+2.*E2P2)+E2P3*AID3)*AID2+ BRLY0790
12.*E2P3*AID4*(5.+(32.*THETA2)/FUN2+(40.*THETA4)/FUN22)) BRLY0800
C3=1.-9.*THETA2-(24.*THETA4)/FUN2 BRLY0810
AID5=CJ31P/CJ21P BRLY0820
C4=.25*AID5*(-E2P*AID3)+5./64.*AID*(-E2P*AID3*(4.+3.*E2P2)+ BRLY0830
1E2P*SINEI*(26.+9.*E2P2))*C3-15./32.*AID*E2P*AID4*(4.+3.*E2P2)* BRLY0840
2(3.+(16.*THETA2)/FUN2+(40.*THETA4)/FUN22)) BRLY0850
C5=E2P/(1.+ETA3)*(3.-E2P2*(3.-E2P2)) BRLY0860
C6=(E2P*(-32.+81.*(E2P2*E2P2)))/((4.+3.*E2P2)+ETA*(4.+9.*E2P2)) BRLY0870
C7=.25*AID5*SINEI*C5+5./64.*C3*AID*ETA2*SINEI*C6 BRLY0880
C8=-.25*AID5*ETA3*SINEI-5./64.*AID*ETA3*SINEI*(4.+9.*E2P2)*C3 BRLY0890
0510 T=DT BRLY0910
CL2P=CL12P*DT+CL22P*DT**2*CL02P + RND*DT*DT
CL2P=AMOD(CL2P,6.2831853071796)
IF(CL2P)520,530,530
520 CL2P=CL2P+6.2831853071796
530 G2P=G12P*DT+G02P BRLY0940
H2P=H12P*DT+H02P BRLY0950
SINEG=SIN(G2P) BRLY0960
COSING=COS(G2P) BRLY0970
D1E=SINEG*(SINEG*(E31P*SINEG+E21P)+E11P)+E(1P BRLY0980
H1P=((H31P*SINEG+H21P)*SINEG+H11P)*COSING+H2P BRLY0990
G1P=G2P+CL2P+.5*(CL21P+G21P)*SIN(2.*G2P)+(C1+C2)*COS(3.*G2P)
1+(C4+C7)*COSING
CL1P=CL2P BRLY1020

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U=CL2P
100 DELTAU=(U-E2P*SIN(U)-CL2P)/(1.-E2P*COS(U))
    U=U-DELTAU
    IF(ABS(DELTAU)-1.E-10)200,100,100
200 U=U-(U-E2P*SIN(U)-CL2P)/(1.-E2P*COS(U))
    E=U
    SINE1P=SIN(E)
    COSE1P=COS(E)
    G1P=G2P
    ADIVR=1./((1.-E2P*COSE1P)
    SINFP=ADIVR*ETA*SINE1P
    COSFP=ADIVR*(COSE1P-E2P)
    F1P=ARTNQ(SINFP,COSFP)
    IF(ABS(F1P-CL2P)-3.1415926535896)220,210,210
210 STOP
220 FUN3=(1.+CN2*T)**.66666667
    COSFG=COS(2.*(G1P+F1P))
    SINFG=SIN(2.*(G1P+F1P))
    ADIVR2=ADIVR**2
    ADIVR3=ADIVR**3
    CI=CI2P+CI0*D1E+CI1*SINFP*SINFG+(2.*CI1*COSFP+CI2)*COSFG
    FUN8=F1P-CL1P+E2P*SINFP
8018 H=H1P+(2.*H1*COSFP+H2)*SINFG-H1*SINFP*COSFG+H3*U=U
    KFUN=H/6.2831853071796
    FUN9=KFUN
    H=H-FUN9*6.2831853071796
    IF(H)8022,8023,8023
8022 H=H+6.2831853071796
8023 A=A2P/FUN3+AD*(A1+A2*COSFG)*ADIVR + ADT*DT
    AID6=ADIVR2*ETA2+ADIVR
    AID7=SIN(2.*G2P+F1P)
    AID8=SIN(2.*G2P+3.*F1P)
    D1=.25*CJ21P*(6.*(5.*THETA2-1.)*FUN8+(3.-5.*THETA2)*(3.*SINFG+
13.*E2P*AID7 +E2P*AID8 ))
    D2=.25*CJ21P*(2.*(3.*THETA2-1.)*(AID6+1.)*SINFP+3.*(1.-THETA2)*
1 ((-AID6+1.)*AID7+(AID6+.33333333)*AID8))
    AID9=COS(2.*G2P+F1P)
    AID10=COS(2.*G2P+3.*F1P)
    D3=-ETA2*.5*CJ21P*(1.-THETA2)*(3.*AID9+AID10)
    ETA6I=1./((ETA3*ETA3)
    D4=ETA6I*(C5 +COSFP*(3.+E2P*COSFP*(3.+E2P*COSFP)))
    D5=ETA6I*(E2P+COSFP*(3.+E2P*COSFP*(3.+E2P*COSFP)))
    D6= ETA2*CJ2*.5*((3.*THETA2-1.)*D4+3.*(1.-THETA2)*D5*COSFG)+D3
    GAL=G1P+D1+(E2P*ETA2)/(1.+ETA)*D2
    CE=(E2P-1.)*(1.+CN2*DT)**.666666666666667+1.+D1E+D6 + ESD*DT
    EDL=.5*E2P*CL21P*SIN (2.*G2P)+C8*COSING+E2P*C1*COS (3.*G2P)-
1 ETA3*D2
    AID14=SIN(CL2P)
    AID15=COS(CL2P)
    ESL=CE *AID14+EDL*AID15
    ECL=CE *AID15-EDL*AID14
    CE=SQRT (ECL*ECL+ESL*ESL)
    CL=ARTNQ(ESL,ECL)
    G=GAL-CL
    G=AMOD(G,6.2831853071796)

```

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IF(G)8024,8025,8025  
 8024 G = G + 6.2831853071796  
 8025 RETURN

8RLY156S  
 8RLY1570

END  
 SUBROUTINE CORDEL(RV,HV,Q,TAUE,U)

C  
 C  
 C  
 C  
 C  
 C  
 C

THIS IS PART OF THE CIS FUNCTION. OSCULATING CARTESIAN POSITION  
 AND VELOCITY VECTORS ARE TRANSFORMED INTO OSCULATING KEPLERIAN  
 ELEMENTS. SEE EQS. (26)-(31).

DIMENSION RV(6),HV(6)  
 X=RV(1)  
 Y=RV(2)  
 Z=RV(3)  
 XDOT=RV(4)  
 YDOT=RV(5)  
 ZDOT=RV(6)  
 R=SQRT(X\*X+Y\*Y+Z\*Z)  
 VSQ=XDOT\*XDOT+YDOT\*YDOT+ZDOT\*ZDOT  
 HX=Y\*ZDOT-Z\*YDOT  
 HY=Z\*XDOT-X\*ZDOT  
 HZ=X\*YDOT-Y\*XDOT  
 H=SQRT(HX\*HX+HY\*HY+HZ\*HZ)  
 AA=1./(2./R-VSQ/Q)  
 ESINU=(X\*XDOT+Y\*YDOT+Z\*ZDOT)/SQRT(Q\*AA)  
 ECOSU=R\*VSQ/Q-1.  
 ESQ=ESINU\*ESINU+ECOSU\*ECOSU  
 E=SQRT(ESQ)  
 ROOT=SQRT(1.-ESQ)  
 ANGLEI=ARTNQ(SQRT(HX\*HX+HY\*HY),HZ)  
 IF(ANGLEI-TAUE)1,5,5  
 1 SGNHZ=HZ/ABS(HZ)  
 ANGLEI=1.5707963257994\*(1.-SGNHZ)  
 OMEGA=0.  
 IF(E-TAUE)2,3,3  
 2 ANOMAL=ARTNQ(SGNHZ\*Y,X)  
 GO TO 9  
 3 PERIG=ARTNQ(SGNHZ\*Y,X)-ARTNQ(ROOT\*ESINU,ECOSU-ESQ)  
 IF(PERIG)4,7,7  
 4 PERIG=PERIG+6.2831853071796  
 GO TO 7  
 5 OMEGA = ARTNQ(HX,-HY)  
 IF(E-TAUE)8,6,6  
 6 PERIG=ARTNQ(Z\*H\*(ECOSU-ESQ)+(X\*HY-Y\*HX)\*ROOT\*ESINU,Z\*H\*ROOT\*  
 1ESINU-(X\*Y-Y\*HX)\*(ECOSU-ESQ))  
 7 U=ARTNQ(ESINU,ECOSU)  
 ANOMAL=U-ESINU  
 GO TO 10  
 8 ANOMAL=ARTNQ(Z\*H,Y\*HX-X\*HY)  
 9 E = 0.  
 PERIG=0.  
 10 CONTINUE  
 C TAU=-ANOMAL\*SQRT(AA\*AA\*AA/Q)/3600.

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```

TAU=ANOMAL
HV(1)=AA
HV(2)=E
HV(3)=ANGLEI
HV(4)=PERIG
HV(5)=OMEGA
HV(6)=TAU
RETURN
END
SUBROUTINE ORBADJ( IOAOP, KSA, TOA, IAYR, IOJDA, ADJSE )

```

C  
C  
C  
C  
C

THIS IS PART OF THE GIS FUNCTION.

```

COMMON / KORBEL / A01,E1,XI1,G01,H01,FL01
DIMENSION RV(6), HV(6)
IF ( IOAOP .EQ. 0 ) RETURN
READ *, IAYR, ADJDA, ADJSE
READ *, RV(1), RV(2), RV(3)
READ *, RV(4), RV(5), RV(6)
TOA = 1721044. + 367*IAYR - (7*IAYR)/4 + ADJDA + (ADJSE/86400.)
1 - 2400001.0
Q = 398600.8
TAUE = 1.E-06
CALL CORDEL ( RV,HV,Q,TAUE,U )
A01 = HV(1)
E1 = HV(2)
XI1 = HV(3)
G01 = HV(4)
H01 = HV(5)
FL01 = HV(6)
PRINT 15
15 FORMAT ( 1H1 )
25 PRINT 20, IOAOP
20 RNAT(42X, 'PROCESSING OPTION *,I4,* SELECTED*,/')
PRINT 10, IAYR, ADJDA, ADJSE
10 FORMAT (39X, 'THE POST ORBIT ADJUST EPOCH IS- YEAR*,I5,* DAY *,
1 F6.1,* SEC *,G16.10)
IOJDA = ADJDA
PRINT 11
11 FORMAT (/,39X, 'INPUT POST ORBIT ADJUST CARTESIAN VECTORS ARE --*)
PRINT 12, ( RV(I), I=1,6 )
12 FORMAT(/,39X, 'X =*,G16.10,* KM*/39X,
1 *Y =*,G16.10,* KM*/39X,
2 *Z =*,G16.10,* KM*/39X,
3 *XDOT=*,G16.10,* KM PER SEC*/39X,
4 *YDOT=*,G16.10,* KM PER SEC*/39X,
5 *ZDOT=*,G16.10,* KM PER SEC*)
RETURN
END
SUBROUTINE MX1502(KSA, ITYPE, FT, IY, ID, BA, XXI, ES, XM, XO, XAM, UJD, TVE,
1 SEC, IOJOP)

```

C  
C

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```

C      THIS IS THE GFS FUNCTION.  THE FOLLOWING SHOULD BE NOTED=
C      (1) THE INPUT VALUE FOR TVE SHOULD BE PERIODICALLY UPDATED
C          FROM THE NAUTICAL ALMANAC.
C      (2) THE YEAR COMPUTATION IAYR ASSUMES THAT ALL VECTOR EPOCHS
C          ARE BETWEEN THE YEARS 1980 AND 1989 .  EPOCH YEARS OTHER
C          THAN THESE WILL NECESSITATE CHANGING 1980 IN THE IAYR
C          EXPRESSION TO THE APPROPRIATE DECADE.
C
C      DATA WE / 7.292115856E-05 /
C      PRINT 10
10  FORMAT ( 1H1 )
C      PI = 3.14159265359
C      PI2 = 2.*PI
C      DEG = PI / 180.
C      IF ( IOAOP .EQ. 0 ) GO TO 15
C      IAYR = IY
C      TM = SEC / 60.
C      GO TO 18
C
C      COMPUTE MINUTES OF DAY (GMT) .  SEE EQ. (80).
C
15  IAYR = 1980 + IY
C      TS = UJD - 1721044. - 367*IAYR + (7*IAYR)/4 - ID + 2400001.0
C      TM = (TS * 86400.) / 60.
C
C      COMPUTE EARTH FIXED LONGITUDE OF THE ASCENDING NODE.  SEE EQ. (81)
C
18  RAG = WE * ( UJD - TVE ) * 86400.
C      RAG = AMOD ( RAG, PI2)
C      RAG = RAG / DEG
C      XOL = XO - RAG
C      IF ( XOL.LT. 0.00 ) XOL = 360. + XOL
C      PRINT 20
20  FORMAT(////,37X,*GEODETIC SATELLITE ORBIT PARAMETERS FOR THE MX 15
C      102-DS GEOCEIVER*)
C      PRINT 30, KSA,ITYPE,IAYR,ID,TH,XAM,XW,ES,BA,XOL,XXI,FT
30  FORMAT(//,57X,*SATELLITE IDENTIFICATION*,6X,I5/57X,
C      1  *SATELLITE TYPE*,16X,I5/57X,
C      2  *ELEMENT SET EPOCH (GMT)*,/76X,
C      3      *YEAR*,7X,I5/77X,
C      4      *DAY*,7X,I5/76X,
C      5      * MIN*,5X,F7.2///44X,
C      6  *MEAN ANOMALY*,20X,E22.14,* DEG*/44X,
C      7  *ARGUMENT OF PERIGEE*,13X,E22.14,* DEG*/44X,
C      8  *ECCENTRICITY*,20X,E22.14/44X,
C      9  *SEMI-MAJOR AXIS*,17X,E22.14,* KM*/44X,
C      A  *LONGITUDE OF ASCENDING NODE*,5X,E22.14,* DES*/44X,
C      B  *INCLINATION*,21X,E22.14,* DEG*/44X,
C      C  *TRANSMISSION FREQUENCY*,10X,E22.14,* PPM*/)
C      RETURN
C      END
C      FUNCTION XSPA(XM)

```

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```

C      COMPUTE THE SHORT PERIODIC VARIATION OF THE SEMI-MAJOR AXIS.
C      SEE EQ. (37).
C
      DIMENSION XM(6)
      COMMON / KORBEL / A, ES, XI, W, OM, AM
      COMMON / CON / DEGRAD, XMU, XJ2, AE
      F = AM + 2.*ES*SIN(AM)+(5./4.)*ES*ES*SIN(2.*AM)+(13./12.)*ES*ES*
1     SIN(3.*AM)
      P = XM(1) * ( 1. - ES*ES )
      RR= P / ( 1. + ES*COS(F) )
      XSPL = XJ2*(AE*AE/XM(1))*(((XM(1)/RR)**3.)*(1.-1.5*SIN(XI)*
1     SIN(XI)+1.5*SIN(XI)*SIN(XI)*COS(2.*W+2.*F))-(1.-1.5*SIN(XI)*
2     SIN(XI))*(1.-ES*ES)**(-1.5))
      RETURN
      END
      FUNCTION XSPL ( XM, SUM, G, XMOT )

C      COMPUTE THE SHORT PERIODIC VARIATION OF LAMBDA.  SEE EQ. (38).
C
      DIMENSION SUM(6), XM(6)
      XSPL = - (2./(XMOT*XMOT*XM(1)))*SUM(2)+(G/(2.*XMOT*XMOT*XM(1)
1     *XM(1)))*XM(3)*SUM(3)+XM(4)*SUM(4)+(1./(2.*XMOT*XMOT*XM(1)*
2     XM(1)*G))*XM(5)*SUM(5)+XM(6)*SUM(6)
      RETURN
      END
      FUNCTION XSPZ ( XM, SUM, G, XMOT )

C      COMPUTE THE SHORT PERIODIC VARIATION OF XI.  SEE EQ. (39).
C
      DIMENSION SUM(6), XM(6)
      XSPZ = -(G/(XMOT*XMOT*XM(1)*XM(1)*(1.+G)))*XM(3)*SUM(1)-(G/(XMOT*
1     XMOT*XM(1)*XM(1))*SUM(4)-(1./(2.*XMOT*XMOT*XM(1)*XM(1)*G))*
2     XM(4)*(XM(5)*SUM(5)+XM(6)*SUM(6))
      RETURN
      END
      FUNCTION XSPXN ( XM, SUM, G, XMOT )

C      COMPUTE THE SHORT PERIODIC VARIATION OF ETA.  SEE EQ. (40).
C
      DIMENSION SUM(6), XM(6)
      XSPXN=-(G/(XMOT*XMOT*XM(1)*XM(1)*(1.+G)))*XM(4)*SUM(1)+(G/(XMOT*
1     XMOT*
2     XM(1)*XM(1))*SUM(3)+(1./(2.*XMOT*XMOT*XM(1)*XM(1)*G))*XM(3)
2     *(XM(5)*SUM(5)+XM(6)*SUM(6))
      RETURN
      END
      FUNCTION XSPP ( XM, SUM, G, XMOT )

```

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```

C
C
C
C
      COMPUTE THE SHORT PERIODIC VARIATION OF F.  SEE EQ.  (41).

      DIMENSION SUM(6), XM(6)
      C1 = 1./(2.*XMOT*XMOT*XM(1)*XM(1)*G)
      XSPQ = -C1*XM(5)*SUM(1) - 9.5*C1*SUM(6) + C1*XM(5)*( XM(4)*SUM(3)-
1  XM(3)*SUM(4))
      RETURN
      END
      FUNCTION XSPQ  ( XM, SUM, G, XMOT )

C
C
C
C
      COMPUTE THE SHORT PERIODIC VARIATION OF Q.  SEE EQ.  (42).

      DIMENSION SUM(6), XM(6)
      C1 = 1./(2.*XMOT*XMOT*XM(1)*XM(1)*G)
      XSPQ = -C1*XM(6)*SUM(1) + 0.5*C1*SUM(5) + C1*XM(6)*( XM(4)*SUM(3)-
1  XM(3)*SUM(4))
      RETURN
      END
      FUNCTION XJLMP(L,M,P,I)

C
C
C
C
      EVALUATE THE INCLINATION FUNCTION J.  SEE EQ.  (52).

      REAL JLMP, I
      INTEGER F, ALPHA, AALPHA, P2
      C = COS(0.5*I)
      S = SIN(0.5*I)
      ALPHA = M-L+2*P
      MALPHA = -ALPHA
      AALPHA = IABS(ALPHA)
      L1 = L-M
      K = 0.5*L1
      DO 5 JJJ = 1,10,2
        IF (L1.EQ.JJJ) K = K + 1
5  CONTINUE
      L2 = 2*L-2*P
      P2 = 2*P
      J1 = 0
      J2 = L1
      L3 = L+M
      L4 = L-P
      FMULT = (FACT(L3)/(FACT(P)*FACT(L4)*(2.**L)))*((-1.)**K)
      IF (MALPHA.GT.J1) J1 = MALPHA
      IF (L2.LT.J2) J2=L2
      IF (J1.GT.J2) PRINT 18, J1, J2
10  FORMAT(1H0,20X,*PROGRAM TERMINATED IN FUNCTION  JLMP---J1 GT J2--
      A      J1 AND J2 =*,2I5)
      IF (J1.GT.J2) STOP
      JLMP = 0.00
      J1 = J1+1

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```

J2 = J2+1
DO 20 JJ = J1, J2
  J = JJ-1
  L5 = L1-J
  F = ((-1.)**J)*BINOM(L2, J)*BINOM(P2, L5)*FMULT
  JLMP = JLMP+F*(C**(2*L-ALPHA-2*J))*(S**(ALPHA-AALPHA+2*J))
20 CONTINUE
XJLMP = JLMP
RETURN
END
FUNCTION DJLMP(L, M, P, I)

```

C  
C  
C  
C  
C  
C

EVALUATE THE DERIVATIVE OF THE INCLINATION FUNCTION WITH RESPECT  
TO  $\cos(0.5*I)$ . SEE EQ. (68).

```

REAL I
INTEGER P, ALPHA, AALPHA, P2
C = COS(0.5*I)
S = SIN(0.5*I)
ALPHA = M-L+2*P
HALPHA = -ALPHA
AALPHA = IABS(ALPHA)
L1 = L-M
K = 0.5*L1
DO 5 JJJ = 1, 10, 2
  IF (L1.EQ.JJJ) K = K + 1
5 CONTINUE
L2 = 2*L-2*P
P2 = 2*P
J1 = 0
J2 = L1
L3 = L+M
L4 = L-P
FMULT = (FACT(L3)/(FACT(P)*FACT(L4)*(2.**L)))*((-1.)**K)
IF (HALPHA.GT.J1) J1 = HALPHA
IF (L2.LT.J2) J2=L2
IF (J1.GT.J2) PRINT 10, J1, J2
10 FORMAT(1H0,20X,'PROGRAM TERMINATED IN FUNCTION DJLMP---J1 GT J2--
A   J1 AND J2 =',2I5)
IF (J1.GT.J2) STOP
DLMP = 0.0
J1 = J1+1
J2 = J2+1
DO 20 JJ = J1, J2
  J = JJ-1
  L5 = L1-J
  F = ((-1.)**J)*BINOM(L2, J)*BINOM(P2, L5)*FMULT
  DLMP = DLMP + F*(C**(2*L-ALPHA-2*J-1))*(S**(ALPHA-AALPHA))*
A   ((2*L-AALPHA)*(S**(2*J))-(2*J+ALPHA-AALPHA)*(S**(2*J-2)))
20 CONTINUE
DJLMP = DLMP
RETURN
END

```

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C  
C  
C  
C  
C

FUNCTION XKLPQ(L,P,Q,GAMA)

EVALUATE THE ECCENTRICITY FUNCTION K. SEE EQS. (53)-(54).

```

INTEGER P,Q,R,T,AQ,RR,TT,PP,PU,PUP
ITST = 2*P - L
XLPQ = 0.00
IF (Q.EQ.ITST) GO TO 100
L1 = -2*P
L2 = 2*P - 2*L
F1 = 0.5*(L - 2*P + Q)
AQ = IABS(Q)
F2 = 1. + GAMA
F3 = 1. - GAMA
F4 = ((-1.)**AQ)*(2.**L)*(F2**(-L-AQ))
DO 10 KK = 1, 3
  K = KK-1
  KU = AQ + KK
  DO 20 RR = 1,KU
    R = RR - 1
    DO 30 TT = 1,KK
      T = TT-1
      L3 = AQ + K - R
      L4 = K - T
      IF(Q.GE.0) BIFAC=((-1.)**R)*BINOM(L2,L3)*BINOM(L1,L4)
    A
      IF(Q.LT.0) BIFAC=((-1.)**T)*BINOM(L1,L3)*BINOM(L2,L4)
    TERM=F4*(BIFAC/(FACT(R)*FACT(T)))*(F1**(R+T))*(F2**(R+T-K))*(F3**K)
    A
      XLPQ = XLPQ + TERM
      TST = ABS(TERM)
      TST1 = 0.61 * XLPQ
      TST1 = ABS ( TST1)
      IF((TST.LT.TST1).AND.(TST.NE.0.00)) GO TO 40
30    CONTINUE
20    CONTINUE
10    CONTINUE
PRINT 11, L,P,Q,GAMA,TST,TST1
11 FORMAT (1H0,3X,'XKLPQ DID NOT CONVERGE- L P Q GAMA TST TST1=',3I5,
A      3F15.7,/)
GO TO 40
100 IPP = IABS(ITST)
L6 = L-1
PP = (L - IPP)/2
PU = PP - 1
IF (PU.LT.0) GO TO 40
PUP = PU + 1
DO 50 KK = 1,PUP
  K = KK - 1
  L7 = 2*K + IPP
  XLPQ = XLPQ + (GAMA**(1.-2.*L))*BINOM(L6,L7)*BINOM(L7,K)*(2.**(-L
A      7))*((1.-GAMA*GAMA)**K)
50 CONTINUE

```

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```

40 XKLPQ = XLPQ
RETURN
END
FUNCTION DKLPQ(L,P,Q,GAMA)

```

C  
C  
C  
C  
C

EVALUATE THE DERIVATIVE OF THE ECCENTRICITY FUNCTION K WITH RESPECT  
TO GAMMA ( = SQRT(1. - E\*E) ) . SEE EQS. (72)-(74).

```

INTEGER P,Q,K,T,AQ,RR,TT,PP,PU,PUP
ITST = 2*P - L
XLPQ = 0.00
X1 = XKLPQ(L,P,Q,GAMA)
IF (Q.EQ.ITST) GO TO 100
L1 = -2*P
L2 = 2*P - 2*L
F1 = 0.5*(L - 2*P + Q)
AQ = IABS(Q)
F2 = 1. + GAMA
F3 = 1. - GAMA
F4 = ((-1.)**AQ)*(2.**L)*(F2**(-L-AQ))
X1 = ((-L-AQ)/F2)*X1
DO 10 KK = 1, 3
  K = KK-1
  KU = AQ + KK
  DO 20 RR = 1, KU
    R = RR - 1
    DO 30 TT = 1, KK
      T = TT-1
      L3 = AQ + K - R
      L4 = K - T
      IF(Q.GE.0) BIFAC=((-1.)**R)*BINOM(L2,L3)*BINOM(L1,L4)
      IF(Q.LT.0) BIFAC=((-1.)**T)*BINOM(L1,L3)*BINOM(L2,L4)
      TERM=F4*(BIFAC/(FACT(R)*FACT(T)))*(F1** (R+T))*(F2** (R+T-K-1))*
      ((F3**K)*(R+T-K)-K*F2*(F3** (K-1)))
      XLPQ = XLPQ + TERM
      TST = ABS(TERM)
      TST1 = 0.01 * XLPQ
      YST1 = ABS ( TST1)
      IF((TST.LT.TST1).AND.(TST.NE.0.00)) GO TO 40
30    CONTINUE
20    CONTINUE
10    CONTINUE
PRINT 11, L,P,Q,GAMA,TST,TST1
11 FORMAT (1H0,3X,'XKLPQ DID NOT CONVERGE- L P Q GAMA TST TST1=',3I5,
A      3F15.7,/)
GO TO 40
100 IPP = IABS(ITST)
X1 = ((-2*L+1)/GAMA) * X1
L6 = L-1
PP = (L - IPP)/2
PU = PP - 1
IF (PU.LT.0) GO TO 40

```

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```

PUP = PU + 1
DO 50 KK = 1,PUP
  K = KK - 1
  L7 = 2*K + IPP
  XLPQ = XLPQ - 2.* (GAMA** (2.-2.*L)) * BINOM(L6,L7) * BINOM(L7,K) *
    (2.** (-L7)) * K * ((1.-GAMA*GAMA)** (K-1))
A
50 CONTINUE
40 DKLPQ = XLPQ + X1
RETURN
END
FUNCTION RLMPQ(L,M,P,Q,A,B,C,D)

C
C
C
C
EVALUATE THE R FUNCTION. SEE EQ. (55).

INTEGER ALPHA,AALPHA,P,Q,AQ,U1,U2,U,UU1,UU2,UU
ALPHA = M - L + 2*P
AALPHA = IABS(ALPHA)
AQ = IABS(Q)
R = 0.00
K = 0.5*(AQ + AALPHA)
KK = K + 1
DEL = 1.0
IPRD = Q*ALPHA
DO 10 NN = 1, KK
  N = NN - 1
  L1 = 2*N - AALPHA
  L2 = 2*N
  U1 = 0
  IF (L1.GT. 0) U1 = L1
  U2 = L2
  IF (AQ .LT. L2) U2 = AQ
  UU1 = U1 + 1
  UU2 = U2 + 1
  IF(U1.GT.U2) PRINT 11, U1, U2
11 FORMAT (1H0,3X,'LOWER BOUND GY UPPER IN SUM OVER U IN FUNCTION RLM
APQ = U1 U2 =*,2I6,/)
  IF (U1.GT.U2) STOP
  DO 20 UU = UU1, UU2
    U = UU - 1
    IF (IPRD.LT.0) DEL = (-1.0)**U
    N1 = 2*N - U
    R = R + ((-1.0)** (N+U)) * DEL * BINOM(AQ,U) * BINOM(AALPHA,N1) *
      (A** (AQ-U)) * (B**U) * (C** (AALPHA-2*N+U)) * (D** (2*N-U))
A
  CONTINUE
20
10 RLMPQ = R
RETURN
END
FUNCTION BILMPQ(L,M,P,Q,A,B,C,D)

```

C  
C  
C  
C

EVALUATE THE I FUNCTION. SEE EQ. (56).

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C

```

INTEGER ALPHA,AALPHA,P,Q,AQ,U1,U2,U,UU1,UU2,UU
ALPHA = M - L + 2*P
AALPHA = IABS(ALPHA)
AQ = IABS(Q)
R = 0.00
K = 0.5*(AQ + AALPHA - 1)
KK = K + 1
DEL = 1.0
IPRD = Q*ALPHA
DO 10 NN = 1, KK
  N = NN - 1
  L1 = 2*N - AALPHA + 1
  L2 = 2*N + 1
  U1 = 0
  IF (L1.GT. 0) U1 = L1
  U2 = L2
  IF ( AQ .LT. L2) U2 = AQ
  UU1 = U1 + 1
  UU2 = U2 + 1
  IF ( U1 .GT. U2) R = 0.0
  IF ( U1 .GT. U2) GO TO 30
  DO 20 UU = UU1, UU2
    U = UU - 1
    IF (IPRD.LT.0) DEL = (-1.0)**U
  IF ((IPRD.EQ.0).AND.(Q.LT.0)) DEL = (-1.0)**U
  IF ((IPRD.EQ.0).AND.(ALPHA.LT.0)) DEL = (-1.0)**U
  N1 = 2*N - U + 1
  R = R + ((-1.0)**(N+U+1))*DEL*BINOM(AQ,U)*GINOM(AALPHA,N1)*
    (A** (AQ-U)) * (B**U) * (C** (AALPHA-2*N+U-1)) * (D** (2*N-U+1))
20 CONTINUE
10 CONTINUE
30 CONTINUE
BILMPQ = R
RETURN
END
FUNCTION BINOM(M,N)

```

C  
C  
C  
C  
C

EVALUATE A BINOMIAL EXPANSION COEFFICIENT.

```

IF (N.LT.0) BINOM=0.00
IF (N.LT.0) RETURN
IF (N.EQ.0) BINOM = 1.00
IF (N.EQ.0) RETURN
J = N
IF (M.LT.0) M = N-M-1
L = M - N
IF ( L .LT. 0 ) BINOM = 0.00
IF ( L .LT. 0 ) RETURN
BINOM = FACT(M) / ( FACT(N)*FACT(L) )
IF (J.LT.0) BINOM = ((-1.0)**N) * BINOM
M = J
RETURN

```

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```
END  
FUNCTION FACT(K)
```

```
C  
C  
C  
C  
C
```

```
EVALUATE A FACTORIAL.
```

```
IF (K.LT.0) PRINT 10,K  
10 FORMAT(1H0,20X,*PROGRAM HAS TERMINATED DUE TO FACTORIAL OF A NEGAT  
AIVE INTEGER---K=*,I5)  
IF (K.LT.0) STOP  
FACT = 1.0  
IF (K.EQ.0) RETURN  
DO 20 I = 1,K  
    FACT = FACT * I  
20 CONTINUE  
RETURN  
END  
6/7/8/9
```

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