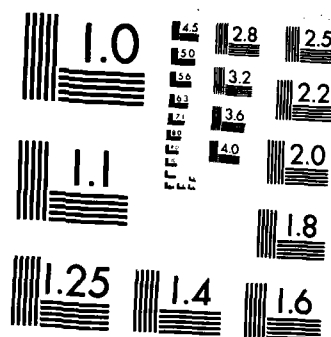


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THE WEAPONS SUPPORT SYSTEM DATA SMOOTHING AND
SPECTRUM ANALYSIS PROGRAM DOCUMENTATION

James V. White

The Analytic Sciences Corporation
One Jacob Way
Reading, Massachusetts 01867

1 October 1981

Scientific Report No. 3

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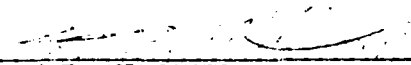
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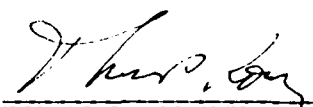
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CONTRACTOR REPORTS

This technical report has been reviewed and is approved for publication.

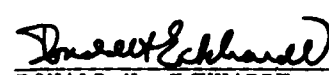


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PREFACE

Documentation for the Weapons Support System (WSS), developed for the Defense Mapping Agency by the Analytic Sciences Corporation, consists of a User's Guide (TR-1946-1) that gives an overview of the WSS hardware and software, as well as detailed operating instructions for the major application areas listed below; and four volumes of detailed program documentation:

- Gravity Data Evaluation Software (TR-1946-2)
- Data Smoothing and Spectrum Analysis Programs (TR-1946-3)
- MULTISENSOR Simulation Software (TR-1946-4)
- GEOFAST Software Documentation (TR-1946-5)

This volume, Data Smoothing and Spectrum Analysis Program Documentation, documents the programs developed for these application areas, at a level suitable for use by programmers who will be adapting or modifying these programs. It is intended for use in conjunction with the Weapons Support System User's Guide and relevant operating system (VAX/VMS) documentation.

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1. PROGRAM OVERVIEW

1.1 INTRODUCTION

This report documents the four programs, GETBELL, AR, FFT, and PLOTFFT, which make up the data smoothing and spectrum analysis software supplied with the Weapons Support System. The programs are designed to run under the VAX/VMS operating system. Information regarding file structures and file-naming conventions is provided in the VAX/VMS System Reference Manuals. The mathematical techniques referred to in this report are discussed in Ref. 1.

1.2 GENERAL SOFTWARE OVERVIEW

The programs GETDATA, AR, FFT, and PLOTFFT are intended for interactive analyses of two-channel time series data. The tasks of selecting data from master data files, smoothing and resampling the selected data, removing linear trends from the data, and plotting the data on the Lexidata graphics terminal or Printronix printer/plotter are all handled by program GETDATA. (The programs AR and PLOTFFT also have plotting capabilities.) Once selected data have been saved in data files by program GETDATA, they may be analyzed by running program AR, which does autoregressive modeling and power-spectrum analyses. Alternatively, the data may be analyzed by running program FFT, which smooths and resamples the master data file, divides it into separate batches and removes linear trends, windows each batch to reduce spectral leakage, and then computes power-spectrum estimates based on periodograms. The results of running program FFT are spectrum files that can be plotted by running program PLOTFFT.

In summary, GETDATA provides a graphical window through which master data sets may be viewed. GETDATA also creates new files of smoothed and resampled data that may be autoregressively analyzed by program AR. Programs FFT and PLOTFFT, in contrast, can be run as a complementary pair for operating directly on master data sets to compute and plot periodogram spectrum estimates.

1.3 DOCUMENT OVERVIEW

Section 2 of this report is a self-contained user's guide for the data smoothing and spectrum analysis programs. The user's guide explains how to run each program, how to interpret the prompting messages displayed on the VT100 monitor, and how to produce graphical plots on the Lexidata graphics terminal and hardcopy plots on the Printronix printer/plotter.

Section 3 contains detailed documentation for programmers who are maintaining the software. It describes the program organization, the data structures used, and the calling sequence of upper-level subroutines. It also lists the names and functions of all subroutines in the programs.

2. USER'S GUIDE TO THE SMOOTHING AND SPECTRUM ANALYSIS PROGRAMS

2.1 INTRODUCTION

There are four programs for the smoothing and spectrum analysis of time series data:

- Data smoothing and selection program GETDATA
- Autoregressive spectrum analysis and plotting program AR
- Periodogram spectrum analysis program FFT
- Periodogram spectrum plotting program PLOTFFT.

These programs are currently set up for analyzing two-channel time series (such as test data for gravity gradiometers that have two output channels). The program GETDATA is used to view selected portions of a master data set on the interactive graphics terminal. In addition, this program allows the user to prepare special data sets for later analysis via the program AR; in particular, the user may use GETADATA to smooth selected portions of a master data set, to resample the smoothed data at a selected rate, to view the results on the graphics terminal, and to save the smoothed and resampled data in a new file on disk.

After using the program GETDATA to plot data for analysis, the user has two choices for further analysis of these data. The first choice is to create a file of selected data by using program GETDATA and then to do spectrum analysis of these data by running program AR. This choice uses both the

GETDATA and AR programs and yields autoregressive spectrum estimates. The second choice is to create a file of periodogram spectrum estimates by running program FFT and then to plot the spectrum estimates by running program PLOTFFT. This second choice uses the FFT and PLOTFFT programs and yields smoothed periodogram spectrum estimates.

The program AR computes an optimal autoregressive model for the process that generated the data and uses this model to plot the following quantities:

- Data being analyzed
- Sum-squared prediction errors of the AR model
- Autospectra for the two data channels
- Spectral coherence between the two data channels.

The program FFT smooths selected batches of data from a master data set, windows the batches to reduce spectral leakage, computes a periodogram spectrum estimate, and saves the results in a disk file for later use by the periodogram plotting program PLOTFFT.

The program PLOTFFT uses the data file created by FFT to plot the following quantities:

- Autospectra for the two data channels
- Spectral coherence between the two data channels.

2.2 DATA FORMATS

The programs are currently set up for two-channel data sets. This means that there should be two data values at each sample time in the master data set. With GETDATA, a maximum of 5000 two-channel samples can be selected for viewing from the master data set at one time.

The master data set should be a file with the following structure:

First Record = ASCII header describing the data set

(length equal to or less than 80 characters)

Example: BELL AF GGI S/N2 UMB ANGLE RUN 06/20/80

Other Records = pairs of data points in BINARY
(2A4) format

2nd Record = 1st datum pair for channel 1 and
channel 2

3rd Record = 2nd datum pair for channel 1 and
channel 2

4th Record = 3rd datum pair for channel 1 and
channel 2

5th Record = 4th datum pair for channel 1 and
channel 2

*

*

*

The name of the master data set should preferably end in the suffix ".DAT"; e.g., "BELL1.DAT" is a valid data file name. (There are certain reserved system names that cannot be used as file names. You may wish to consult your systems manager for information about restrictions on data file names.)

2.3 DATA SELECTION (PROGRAM GETDATA)

To select data from a master data set, type RUN GETDATA. You will be asked for the name of the file containing the master data set. There are now three courses of action available to you:

1. Enter <RETURN> to choose the current default file
2. Type the name of a file using standard VAX format
3. Type QUIT to terminate the program. You may use this command whenever the program asks you a question.

You will now be asked for the number of smoothed data samples you wish to generate from the master data set. You may either enter <RETURN> to select the current default value displayed on the VT100 monitor, or you may type in a number of your choice followed by <RETURN>.

The next prompt asks for the location of the first datum you wish to select from the master data set. Typing 1 chooses the first datum in the master data set. You have the choice of either entering <RETURN> to select the current default value for the location, or typing a number of your choice.

The next prompt asks for the decimation factor. This number controls the degree to which the data are to be smoothed and resampled. For example, a decimation factor of 1 results in no smoothing and no resampling, while a value of 10 produces a selected set of data in which each sample is the average of 10 samples in the master data set. Therefore, in this case the sampling rate in the smoothed data set is one tenth the rate in the master data. You have the choice

The program continues by displaying a header that becomes the first part of the new data file. The first line of this header is a copy of the header in the master data set. The remaining lines list the decimation factor, the name of the file to be created, the number of data samples in the new file, the sampling rate of the smoothed data, the location of the first datum taken from the master data file in creating the current data set, and information about any linear trends (ramps) that might have been removed from the resampled data (the slopes of linear trends are expressed in data units per sample). You are now given the opportunity of typing a single line of additional information about the smoothed and resampled data. The maximum length of this line is 72 characters. Avoid using the dollar sign "\$" because this symbol will cause all others following it to be ignored during certain printing operations. The valid characters are letters of the alphabet, integers, comma, period, and the following special symbols:

+ - * / () =

You may use the <DELETE> key to correct typing errors before pressing <RETURN>. After <RETURN> is pressed, the program asks for a plot header. This is a short string of characters (length equal to or less than 22 characters) that will be printed automatically at the top of each plot. After you type the plot header, enter <RETURN>.

The program continues by creating the new data file on disk and then displaying

DO YOU WANT ANOTHER REGION OF DATA?

The program continues with the prompt

ENTER FRACTIONAL BANDWIDTH

The fractional bandwidth is a real number greater than or equal to zero that determines the frequency resolution of the spectrum estimates after the averaged periodogram is smoothed with a running mean. The width of the running mean is a fraction of its center frequency, and this fraction is the fractional bandwidth. This method of smoothing the periodogram yields constant-percent-bandwidth resolution, which is appropriate for the logarithmic frequency scales that are normally used for plotting spectra.

The program continues by processing the specified batches of data. During this processing, the means are subtracted from each windowed batch and then periodograms for each batch of data are computed with a mixed-radix Fast Fourier Transform (FFT) algorithm. These periodograms are averaged to produce a mean periodogram; the standard deviations of the periodograms are also computed if more than one batch of data is available.

The program continues by asking for the name of the file in which you wish to save the mean periodogram and standard deviations:

PLEASE ENTER THE NAME OF THE OUTPUT FILE

You may enter <RETURN> to choose the default file name, or type a file name of your choice. Examples of valid file names are BELL1 or BELL1.DAT. (The program will append the suffix ".DAT" automatically if you omit it.)

Entering <RETURN> selects the default option. Typing Y or N selects YES or NO options. Selecting option YES causes the program to subtract a zero-mean least-squares straight line from each channel of data. These lines are computed for each batch of data.

The program continues by displaying

ENTER KAISER-WINDOW ALPHA PARAMETER

You may enter <RETURN> to select the default value, or type a real number greater than or equal to zero. The Kaiser window is a bell-shaped weighting function. The program multiplies each batch of data by this weighting function to reduce spectral leakage in the periodograms. A value of zero for the alpha parameter yields a rectangular window (constant weighting). The larger the alpha parameter, the greater the attenuation of spectral leakage. However, larger values of alpha also yield lower frequency resolution in the periodogram. A value of 5.4414 has given satisfactory results in the analysis of gravity gradiometer self-noise data. The Kaiser window is defined in Ref. 1.

The program continues by asking for the number of seconds between adjacent samples in the master data set

ENTER SAMPLING INTERVAL

You may enter <RETURN>, which chooses the default sampling interval displayed on the VT100 monitor, or type a positive real number of your choice. If you do not know the correct sampling interval, you may type 1. The program will then continue to execute, but the spectral plots will be incorrectly labeled.

You may enter <RETURN> to choose the default number of batches, or type an integer of your choice. The program will attempt to segment the master data file into as many nonoverlapping adjacent batches as you specify. If there are insufficient data, the program forms as many batches as it can from the available data. These batches will contain the number of records you requested for each batch.

The program continues by displaying

ENTER STARTING RECORD

You may enter <RETURN> to select the default record, or type an integer. For example, if you type 1000, the program will skip the first 999 pairs of data in the master data file and start the first batch with record 1000. (Plots of the master data set can be generated interactively by running program GETDATA before you run program FFT.)

The program continues by displaying

ENTER DECIMATION FACTOR

You may respond by entering <RETURN> to choose the default decimation factor, or typing an integer. The value 1 results in no decimation (no smoothing and resampling). A value greater than unity causes the program to smooth the master data with a running mean and then to resample these smoothed data. The decimation factor is equal to the number of data points in the running mean, and it is also equal to the factor by which the sampling interval in the master data set is increased by the resampling process.

The program continues by displaying

DO YOU WANT A RAMP REMOVED FROM EACH BATCH?

Whenever the program asks you a question, you may terminate the program by typing

QUIT

The program starts by prompting for the name of the master data file that you wish to use:

ENTER INPUT FILE NAME

You may respond by entering <RETURN> to select the default file name displayed on the VT100 monitor, or you may type a file name of your choice.

The program responds by displaying the prompt

ENTER NUMBER OF RECORDS IN A BATCH

You may enter <RETURN>, which chooses the default value displayed on the VT100 monitor, or type a positive integer. This integer may be any integral power of 2 in the range from 2 to 4096, or it may be a highly composite number, such as 500, 1000, etc. If you choose a number that is not highly composite, such as 499, then the program will fail when the periodogram computation is attempted. (This requirement for composite batch sizes is inherent in the FFT algorithm.) The number of records in a batch is the number of consecutive data pairs from channels 1 and 2 of the master data file that are used to compute each periodogram. The larger this number, the greater the number of frequency bins in the periodogram.

The program continues by displaying

ENTER NUMBER OF BATCHES

These choices are self-explanatory except for choice F, which allows you to recompute the power spectra and coherence with a bandwidth and frequency resolution of your choice. The F option allows you to specify a minimum frequency limit, a maximum frequency limit, and the number of different frequencies at which the spectra and coherence are to be evaluated between these limits. The P, C, and F options cause the program to ask for minimum and maximum frequency limits of your choice. You may invoke autoscaling for these limits as follows: type 1 for both minimum and maximum frequency limits -- the program will respond by autoscaling the frequency scale. This technique of typing 1 for both minimum and maximum limits also invokes autoscaling when the program asks for y-axis limits in the P and C options.

The original time series data are not available for replotting once the power spectra have been computed.

2.5 PERIODOGRAM SPECTRUM ANALYSIS AND PLOTTING (PROGRAMS FFT AND PLOTFFT)

Any master data file having the format described in section 2.2 can be analyzed by running program FFT. This program estimates the power spectrum of the data by computing an averaged matrix periodogram from selected batches of windowed data in the master data file. Program FFT automatically saves the periodogram and its standard deviations in a disk file. After program FFT has created the periodogram file, you may plot the auto spectra and squared coherence on the Lexidata graphics terminal by running program PLOTFFT.

To run program FFT, type

RUN FFT

It is usually advisable to choose Y; later in the program you have the choice of overriding the autoscaling mode. (For illustrative purposes it is assumed that autoscaling is chosen.)

The program continues by computing the complex spectral density matrix for the best AR model. This matrix is computed at 300 logarithmically-spaced frequencies spanning the range from 0.3 times the reciprocal of the data length to the folding frequency (half the sampling frequency of the data being analyzed).

The program then continues by plotting the autospectra for both data channels on the Lexidata graphics terminal. The program displays the usual three-choice menu to give you the opportunity to print copies of the plots on the Printronix printer/plotter.

Entering <RETURN> or typing CL clears the screen. The program then continues by plotting the spectral coherence between the two channels of data, and the usual menu is displayed on the VT100 monitor. You may then choose to print the coherence function on the Printronix printer/plotter.

You may now replot the autospectra or the coherence with linear or logarithmic scales of your choice. This is done by entering <RETURN> or typing CL, causing a new menu to appear on the VT100 monitor:

- P - PLOT POWER SPECTRUM
- C - PLOT COHERENCE
- F - CHANGE THE FREQUENCY RESOLUTION
- Q - QUIT (terminates the program)

You have the choice of entering <RETURN>, which chooses the default value for model order, or you may type a number in the range of 1 through 40. The larger the order selected, the longer the computations take for fitting the models to the data. A maximum model order of 20 is reasonable for many data sets.

Once the maximum AR model order is selected, the program subtracts the means from both channels of data and then fits a family of AR models to them. The Akaike Information Criterion (AIC) is computed for each model and displayed on the VT100 monitor. The model order for which the AIC attains its algebraically smallest value is selected by the program. This is the model order that is best supported by the data for the purposes of modeling the process that generated the data. The program continues by plotting the sum-squared one-step-ahead prediction errors of the selected model for each data channel. If these plots are nearly straight lines, then the data have uniform statistical properties. On the other hand, if there are sudden large deviations from a straight line in these plots, then the locations of these deviations show places in the data set where there is exceptional behavior that is atypical of the rest of the data. These plots are useful for identifying outliers in nonhomogeneous data sets.

The menu described previously is now displayed; you have the option of continuing the program, printing the color graphics plot on the Printronix printer/plotter, or ending the program. When you continue the program by entering <RETURN> or typing CL, the following prompt appears on the VT100 monitor:

DO YOU WANT AUTO SCALING OF PLOTS?

You may respond by typing either Y or N.

terminal. A caption is plotted at the bottom of the display. The first line of the caption is the header from the master data file that was used in creating the data set currently being analyzed. The second and third lines contain the following information

- Decimation Factor (used in smoothing and resampling of data)
- Data Set (name of data file currently being analyzed)
- Number of Records (number of data samples currently being analyzed)
- Ramps (tells whether or not ramps were removed from the data and lists their slopes in data units per sample).

The fourth line in the caption contains any information that you previously typed into the header when the current data set was created by program GETDATA. This four-line caption will appear at the bottom of each plot generated by program AR.

After the program is finished plotting the data, it displays a menu on the VT100 monitor. The three options in this menu have the following meanings

<cr>/CLear - clears the Lexidata screen and continues the program

Print - prints a copy of the Lexidata plot on the Printronix printer/plotter

Quit - terminates the program

After you enter <RETURN> or type CL, the program displays the prompt

MAXIMUM AUTOREGRESSIVE MODEL ORDER (<41)

GETDATA are valid. In typing the name of the desired file, the suffix ".DAT" may be omitted; e.g., BELL1 is a valid file name. (There are certain reserved system names that can not be used as file names. You may wish to consult your systems manager for information about restrictions on data file names. Examples of invalid file names for this program are QUIT and EXIT.) Entering <RETURN> selects the default filename.

After reading the data from the selected file, the program displays the prompt

DO YOU WANT RAMPS REMOVED?

(You always have the option of terminating the program when a yes-no type questions like this one is asked; just type QUIT <RETURN>.)

To continue the program, you may respond by entering <RETURN>, which selects the default choice displayed on the VT100 monitor, or you may type Y or N. If you respond with Y, then a straight line is fitted to each channel of data by least squares and subtracted; the straight line is constrained to have an average value of zero so that the arithmetic means of the data in each channel are unchanged. The slopes of the straight lines are listed on the VT100 monitor and also listed in the captions of all graphs. The units of the slopes are data units per sample.

The program next displays the prompt

DO YOU WANT A PLOT OF THE DATA?

You may respond by typing Y or N. If you respond with Y, then both channels of data will be plotted on the Lexidata graphics

of each plot. After you type the plot header, enter <RETURN>. The program then creates the new file of smoothed and resampled data.

The next prompt asks if you want to start all over again and select more data from a master data file. You may respond with Y or N (or YES, NO). Typing Y causes the program to start over again. Typing N causes the program to terminate.

2.4 AUTOREGRESSIVE SPECTRUM ANALYSIS (PROGRAM AR)

Any data file created by GETDATA may be analyzed by running program AR, which plots the data, subtracts the mean from each channel of data and fits a family of autoregressive (AR) models to the residuals, selects the best model via the Akaike information criterion, and uses this model to generate prediction error, power spectrum, and coherence plots. The program AR is interactive and prompts the user at each step. In the following discussion, it is assumed for illustrative purposes that a data file named BELL1.DAT has previously been created by program GETDATA.

It is recommended that you start by typing DIR *.DAT <RETURN>. This causes a list of the data files in the current library to be displayed on the VT100 monitor for easy reference. For the present example, BELL1.DAT would be among the file names listed on the monitor. (If you do not want a list of the data files displayed on the VT100 monitor, then this step should be omitted.)

To run program AR, type RUN AR. The program AR will then respond with a prompt asking for the name of the data file that is to be analyzed. Data files created by program

automatically by the program to save your resampled data set for later analysis by the program AR. If you do not plan to use AR, then there is no need to use this file-creating option. The rest of this paragraph describes how to proceed should you wish to save your resampled data set. Enter <RETURN>, to select the default file name displayed on the VT100 terminal, or type a file name of your own choice. Examples of valid file names are BELL1 or BELL1.DAT. (The program will append the suffix ".DAT" automatically if you omit it.) The program then displays a header that becomes the first part of the new data file. The first line of this header is a copy of the header in the master data set. The remaining lines list the decimation factor, the name of the file to be created, the number of data samples in the new file, the sampling rate of the smoothed data, the location of the first datum taken from the master data file in creating the current data set, and information about any linear trends (ramps) that might have been removed from the resampled data (the slopes of linear trends are expressed in data units per sample). You are now given the opportunity of typing a single line of additional information about the smoothed and resampled data. The maximum length of this line is 72 characters. Avoid using the dollar sign "\$" because this symbol will cause all others following it to be ignored during certain printing operations. The valid characters are letters of the alphabet, integers, comma, period, and the following special symbols

+ - * / () =

You may use the <DELETE> key to correct typing errors before pressing <RETURN>. After <RETURN> is pressed, the program asks for a plot header. This is a short string of characters (length $\leq$ 22 characters) that will be printed automatically at the top

The program next generates plots of the smoothed and resampled data on the Lexidata graphics terminal, provided that the Lexidata terminal is available. The plot is for channel 1.

A menu is now displayed on the VT100 monitor; you may select any of the three possibilities listed there by typing the command you select. As a shortcut, you may type only the capitalized portion of the command. In this menu, <cr> is an abbreviation for <RETURN>. The commands are used for the following purposes:

<cr>/CLear - to clear the Lexidata screen and continue to the plotting of the data in channel 2

Print - to print a copy of the graph displayed on the Lexidata screen

Quit - to terminate the program

The abbreviations CL, P, and Q are valid abbreviations for CLEAR, PRINT, and QUIT.

After the data for channel 2 have been plotted on the Lexidata terminal, the menu is again displayed. This time when <cr>/CLear is selected, the program responds by asking whether you wish to save the data in a disk file. You may respond by typing YES, NO, or QUIT. The abbreviations Y, N, and Q are valid.

Typing Q will terminate the program. Typing N will cause the program to ask if you wish to continue by selecting some more data from the master data set (typing NO or N in response to this question will terminate the program). Typing Y will cause the program to ask for the name you want to use for the output file. This is the file that will be created

of either entering <RETURN> to select the default value for the decimation factor, or you may type a number of your choice.

The next prompt asks you for the sampling interval. This is the number of seconds between adjacent samples in the master data set. If you do not know the correct value for this quantity, you may type 1; the program will then continue to run and you will be able to view the data, but the header information that is generated for the selected data may be incorrect. You have the choice of either entering <RETURN> to select the current default value of the sampling interval, or you may type a number of your own choice.

The next prompt asks for the name of the dimensional units for the data being analyzed. For example, the correct response to this prompt for gravity gradiometer data is EOTVOS. As usual, you may simply enter <RETURN> to select the current default value for data units displayed on the VT100 monitor.

The program now responds by reading the master data file, listing the ASCII header from the master file, and printing the number of smoothed and resampled data points that have been produced during the reading operation. The number of resampled data points may be smaller than the number you requested; this happens when there are not enough data in the master data file to satisfy your request.

The next prompt asks whether you want linear trends removed from each channel of the processed data set. Responding with Y (or YES) to this prompt causes a best-fit least-squares straight line of zero mean value to be subtracted from each channel of resampled data. You may respond by entering <RETURN> for the current default setting, or you may type either N (or NO) to defeat the straight-line subtraction.

This allows you either to terminate the program by typing N, or to start the program over again by typing Y. You may alternatively enter <RETURN> to select the default option displayed on the VT100 monitor.

To plot the periodogram spectra on the Lexidata graphics terminal, you should run program PLOTFFT by typing

RUN PLOTFFT

Whenever the program asks you a question, you may terminate the program by typing

QUIT

Program PLOTFFT starts by displaying

ENTER INPUT FILE NAME

Any file created by program FFT is valid. You may either enter <RETURN> to select the default file name, or you may type a file name of your choice.

The program continues by displaying a four or five line header that describes the data set selected for analysis. The first line is the header from the master data set from which the current data set was derived. The second line gives the decimation factor and the name of the current data set. The third line lists the number of batches, number of samples per batch, the sampling interval, and the location (NSTART) of the first datum taken from the master data file in creating the current data set. The fourth line lists the Kaiser-window alpha parameter and the fractional bandwidth. The last line may be blank or may contain a descriptive message about the data.

The program continues by plotting the auto spectra on the Lexidata graphics terminal. (The plotting scales are auto-scaled. The solid lines are the estimated spectra; the dotted lines are standard errors of these estimates.) A menu is displayed on the VT100 terminal, which gives you the option of typing

<return> to clear the Lexidata screen and continue

P<return> to produce a copy of the graphics on the printer/plotter

Q<return> to terminate the the program

The program continues by plotting the estimated spectral coherence between the two data channels and the standard errors of the estimate (dotted lines). The menu is displayed to give you the opportunity of selecting a hard copy of the graphics.

The program continues by displaying a new menu that allows you to change the plotting parameters

P - PLOT POWER SPECTRUM

C - PLOT COHERENCE

F - CHANGE THE FREQUENCY RESOLUTION

Q - QUIT (terminates the program)

These choices are self-explanatory except for choice F, which allows you to replot the power spectra and coherence with a different amount of smoothing applied to the periodogram. Typing F results in the prompt

ENTER FRACTIONAL BANDWIDTH

You may enter <RETURN> to select the default bandwidth, or type a positive fraction of your choice. The smaller the fractional bandwidth, the greater the frequency resolution of estimated spectra. However, smaller bandwidths also increase the standard errors of the estimated spectra. Therefore, some experimentation is needed to choose a fractional bandwidth that yields the best tradeoff between resolution and stability of the estimated spectra. Typical values of the fractional bandwidth lie in the range of 0.1 to 0.3.

If you type either P to replot the power spectrum, or C to replot the coherence, then the program asks for the type of scales for the X (horizontal) and Y (vertical) axes of the plots. The choices include all possible combinations of linear and logarithmic scales.

The program continues by asking you for the minimum and maximum frequency limits. You may enter <RETURN> to select the default value, or type a number of your choice. To invoke autoscaling, type 1 for both the minimum and maximum frequency limits. Autoscaling for the y-axis scales may also be invoked by typing 1 for both minimum and maximum limits.

The program continues by plotting the auto spectra or coherence on the Lexidata graphics terminal. After the plotting is finished, a menu appears on the VT100 monitor to give you the choice of printing a copy of the graphics on the plotter, continuing, or quitting. Typing Q terminates the program. Typing P causes the graphics to be printed on the plotter. Entering <RETURN> or typing CL clears the screen and continues the program by returning to the menu of plotting options.

2.6 ERROR AND DIAGNOSTIC MESSAGES

The smoothing and spectrum analysis programs have many built-in error traps to flag erroneous data supplied by the user. When you are prompted to select an option from a menu, the program checks the validity of any selected option and reprompts you if an invalid option is requested. The program also checks for consistency in the plotting limits and reprompts if some of the values are invalid, such as a zero value to be plotted on a logarithmic scale.

The following message will appear on the screen if the program cannot open and read a data file you specified:

ERROR IN OPENING DATA FILE

The two most likely causes of this error are that you misspelled the file name, or that the file you specified is not compatible with the program you are running. For example, programs GETDATA and FFT can read master data files, while program AR is designed to read data files created by GETDATA, and program PLOTFFT is designed to read data files created by program FFT.

Program AR is intended for analyzing bivariate (two-channel) time series. The program will fail if the time series is degenerate (i.e., one of the channels is identically zero or if both channels contain identical data).

3.

DETAILED PROGRAM DESCRIPTIONS

3.1 GETDATA PROGRAM

GETDATA performs the data selection phase for the autoregressive analysis program (AR). GETDATA may also be used to view and plot data from a master data file, (before processing the data, for example, with program FFT). A block flow chart of the subroutine structure is given in Fig. 3.1-1. Some of the subroutines are common to both phases (GETDATA and AR). Section 3.7 gives a brief programmer-oriented description of each subroutine, including function and input/output variables.

The GETDATA program consists of a main program and a library containing the subroutines, named ARLIB.OLB. The main program is a driver routine that reads default file names and parameter input from a datastream file, and sets default parameters for the default data set. The name of the datastream file is GETDATA.DEF and can be modified to fit the new data sets. The following is an example of the default values for a BELL gradiometer:

NRECS	= 1000	!number of records in output data set
NDEC	= 1	!decimation factor
NSTART	= 1	!starting record number
SINT	= 2.0	!sampling interval in seconds
UNITS	= 'EOTVOS'	!data units
CSLOPE	= 'NO'	!slope removal flag
NMAX	= 5000	!maximum number of output records
FNAME	= 'BELLUMB.DAT'	!input master data file name
ONAME	= 'TEMPAR'	!output selected data set name
	;	!end of input stream

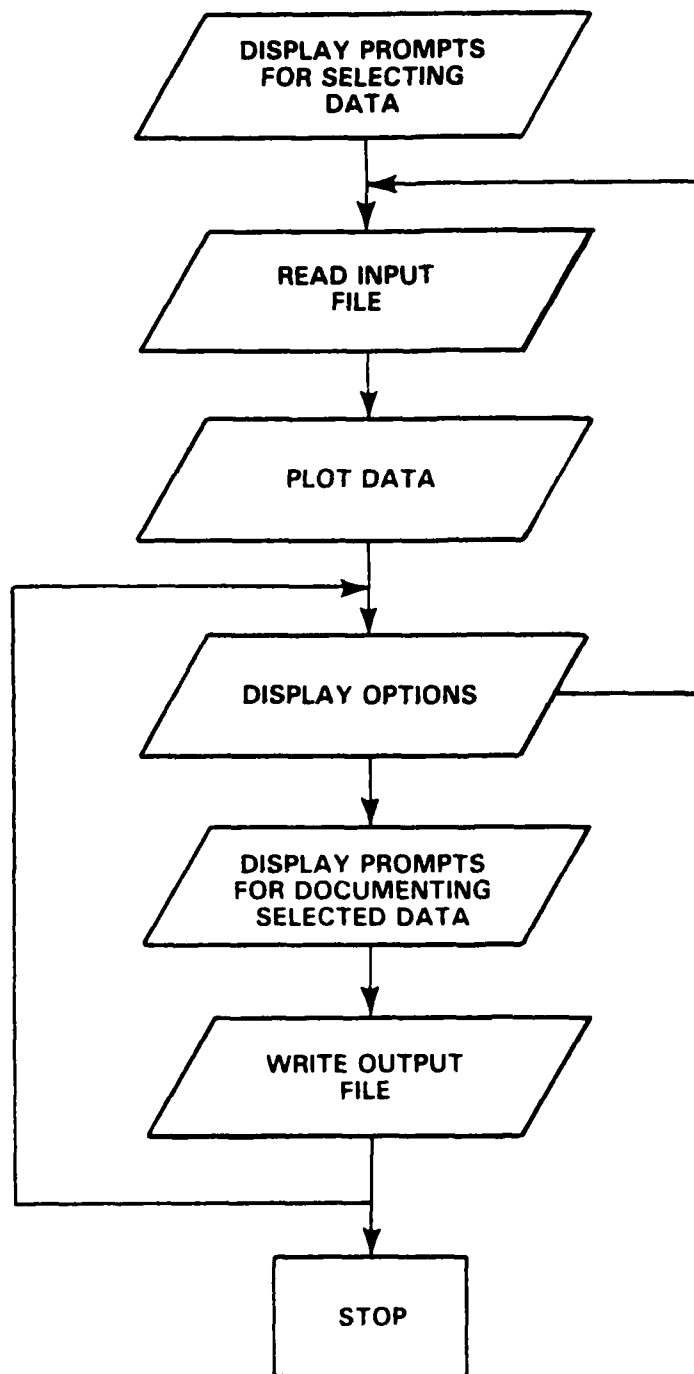


Figure 3.1-1 Block Flow Chart of GETDATA

The maximum size of an output data set is fixed at 5000 points. To increase this size it is necessary to increase the array dimensions to the desired size in the main program and then recompile and link the program. To recompile and link GETDATA the user enters the following commands:

```
FORTRAN GETDATA      !compiles main program
@LG                  !links program and creates new
                     !load module
```

If changes in any of the subroutines are necessary then the following command sequence can be used to create a new version of GETDATA from the updated source:

```
@CL 'program name'  !compiles and loads into ARLIB
@LG                  !links program and creates new
                     !load module
```

Basic program information in a standard format is given below.

Name: GETDATA

Function: This main program creates and plots data files from master data files

Common Blocks: *NONE*

Subprograms called: NMREAD, INPCHR, INPIN4, INPRL4, PRTINF, JREMOV, ANOTAT, EZXY, DSCLS

3.2 AR PROGRAM

AR performs autoregressive modeling and plotting functions for autoregressive spectrum estimation. A block flow chart of the subroutine structure is given in Fig. 3.2-1. Some

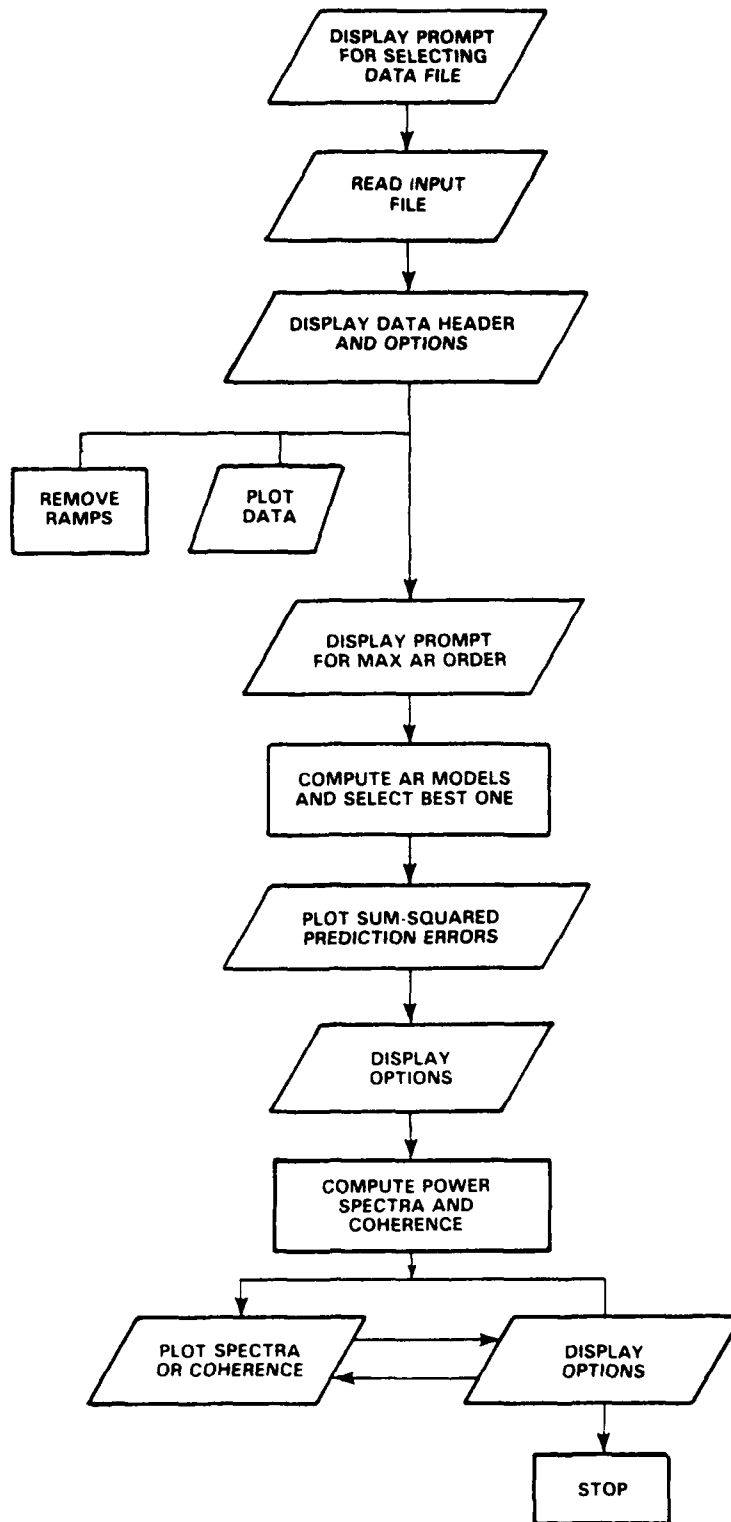


Figure 3.2-1 Block Flow Chart of AR

of the subroutines are common to both GETDATA and AR. The mathematical algorithms involved are discussed in Ref. 1. Section 3.7 gives a brief programmer-oriented description of each subroutine, including function and input/output variables.

The AR program consists of a main program and a library containing the subroutines, named ARLIB.OLB. The main program is a driver routine that reads default file names and parameter input from a datastream file, and sets default parameters for the default data set. The name of the datastream file is AR.DEF and can be modified to fit the current data set. The following is an example of the default values for a BELL gradiometer:

NBINS	= 300	!number of frequency bins
MAXORDER	= 20	!maximum autoregressive order used
NDIM	= 2	!number of data channels
IPLT	= 'YES'	!raw data plot flag
AUTOSCAL	= 'YES'	!automatic plot scaling flag
RAMP	= 'NO'	!ramp removal flag
;		!end of input stream

The maximum size of the input data set is fixed at 5000 points. To increase this size, it is necessary to increase the array dimensions to the desired amount in both the main program and JPLOT; then recompile and link the programs. To recompile and link the programs, the user enters the following commands:

FORTRAN AR	!compiles main program
@CL JPLOT	!compiles and loads JPLOT into ARLIB
@LG	!links program and creates new load module

If changes in any additional subroutines are found to be necessary then the following command sequence can be used to create a new version of AR using the updated source:

```
@CL 'program name'  !compiles and loads into ARLIB
@LG                 !links program and creates new
                   load module
```

Basic program information in a standard format is given below.

Name: AR

Function: Autoregressive modeling of time series and spectrum estimation

Common Blocks: WKSPAC

Subprograms called: NMREAD, INPCHR, PRTINF, JREAD, AGGETP, J8REM, JPLOT1, INPIN4, JEQUAT, MMAXEN, JSUMRD, CPOWER, INPRL4, DSCLS

3.3 FFT PROGRAM

FFT performs the data selection and periodogram computations for use by the periodogram spectrum plotting program (PLOTFFT). A block flow chart of the subroutine structure is given in Fig. 3.3-1. Some of the subroutines are common to both FFT and PLOTFFT. A discussion of the mathematical algorithms involved is presented in Ref. 1. Section 3.8 gives a brief programmer-oriented description of each subroutine, including function and input/output variables.

The FFT program consists of a main program and a library containing the subroutines named PERIODLIB.OLB. The main program is a driver routine that reads default file names and parameter input from a datastream file, and sets default parameters for the default data set. The name of the datastream file is FFT.DEF and can be modified to fit the current data set. The following is an example of the default values for a BELL gradiometer:

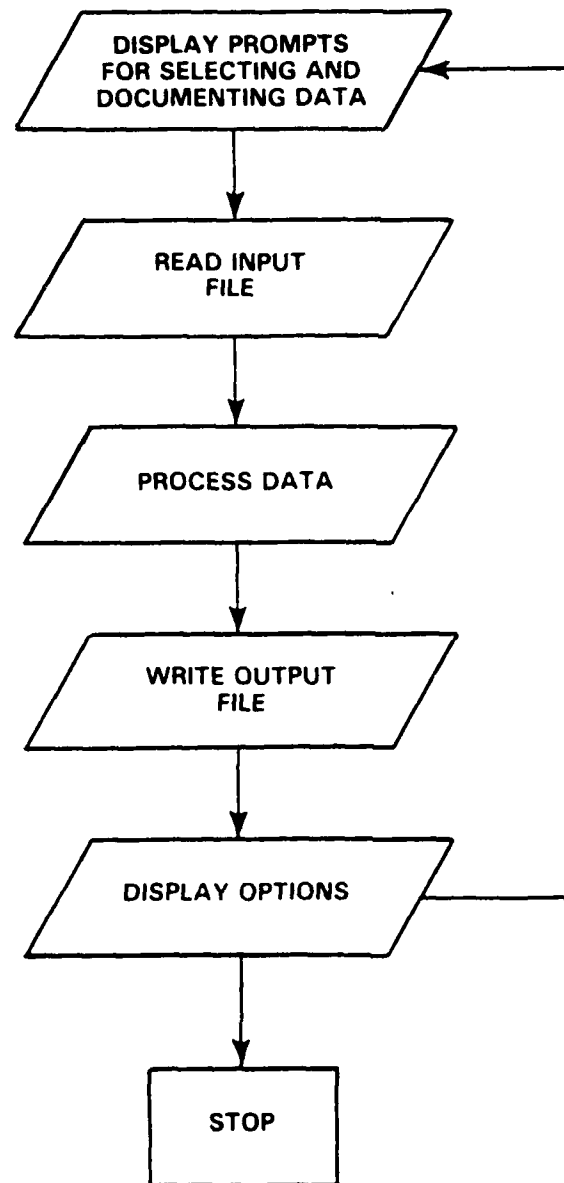


Figure 3.3-1 Block Flow Chart of FFT

```

CSLOPE = NO           !slope removal flag
NSIZE  = 1024         !size of individual batches
NBATCH = 20           !number of batches
NSTART = 1            !starting record number in master data file
FRACB  = 0.1          !fractional bandwidth
ALPHA  = 5.4414       !Kaiser data-window parameter
SINT   = 2.E0         !sampling interval in seconds
NDEC   = 1            !decimation factor
IN FILE = 'BELLUMB.DAT' !input file name
OUT FILE = 'TEMPFILE'  !output file name
;                    !end of input stream

```

The maximum size of a batch is fixed at 4096 points. To increase this size it is necessary to increase the array dimensions to the desired size in the main program and then recompile and link the program. To recompile and link FFT, the user enters the following commands:

```

FORTRAN FFT          !compiles main program
@FFT                 !links program and creates new
                     !load module

```

If changes in any of the subroutines are necessary then the following command sequence may be used to create a new version of FFT from the updated source:

```

@CL 'program name'   !compiles and loads into arlib
@FFT                 !links program and creates new
                     !load module

```

Basic program information in a standard format is given below.

Name: FFT

Function: Estimates power spectra and coherence from two-channel time series via periodogram analysis

Common Blocks: TRIGS, NNTEST

Subprograms called: NMREAD, INPCHR, INPIN4, PRTINF, INPRL4, RDATA,
SMEAN, RSLOPE, KWINDOW, FFTKW, REALVEC, FMEAN,
STDEV

3.4 PLOTFFT PROGRAM

PLOTFFT plots the power spectra and coherence data produced by the FFT phase of the power spectrum analysis. A block flow chart of the subroutine structure is given in Fig. 3.4-1. Some of the subroutines are common to both phases (FFT and PLOTFFT). A discussion of the mathematical algorithms involved is presented in Ref. 1. Section 3.8 gives a brief programmer-oriented description of each subroutine, including function and input/output variables.

The maximum size of the input data set is fixed at 4096 points. To increase this size it is necessary to increase the array dimensions to the desired size in the main program PLOTFFT and subroutines SMFFTS, SMPLOTS, and PPLOT; then recompile and link the programs. To recompile and link PLOTFFT one enters the following commands:

FORTRAN PLOTFFT	!compiles main program
@CL SMFFTS	!compiles and loads SMFFTS into PERIODLIB
@CL SMPLOTS	!compiles and loads SMPLOTS into PERIODLIB
@CL PPLOT	!compiles and loads PPLOT into PERIODLIB
@PLOTFFT	!links with main program and creates new load module

If changes in any additional subroutines are found to be necessary then the following command sequence can be used to create a new version of PLOTFFT from the updated source:

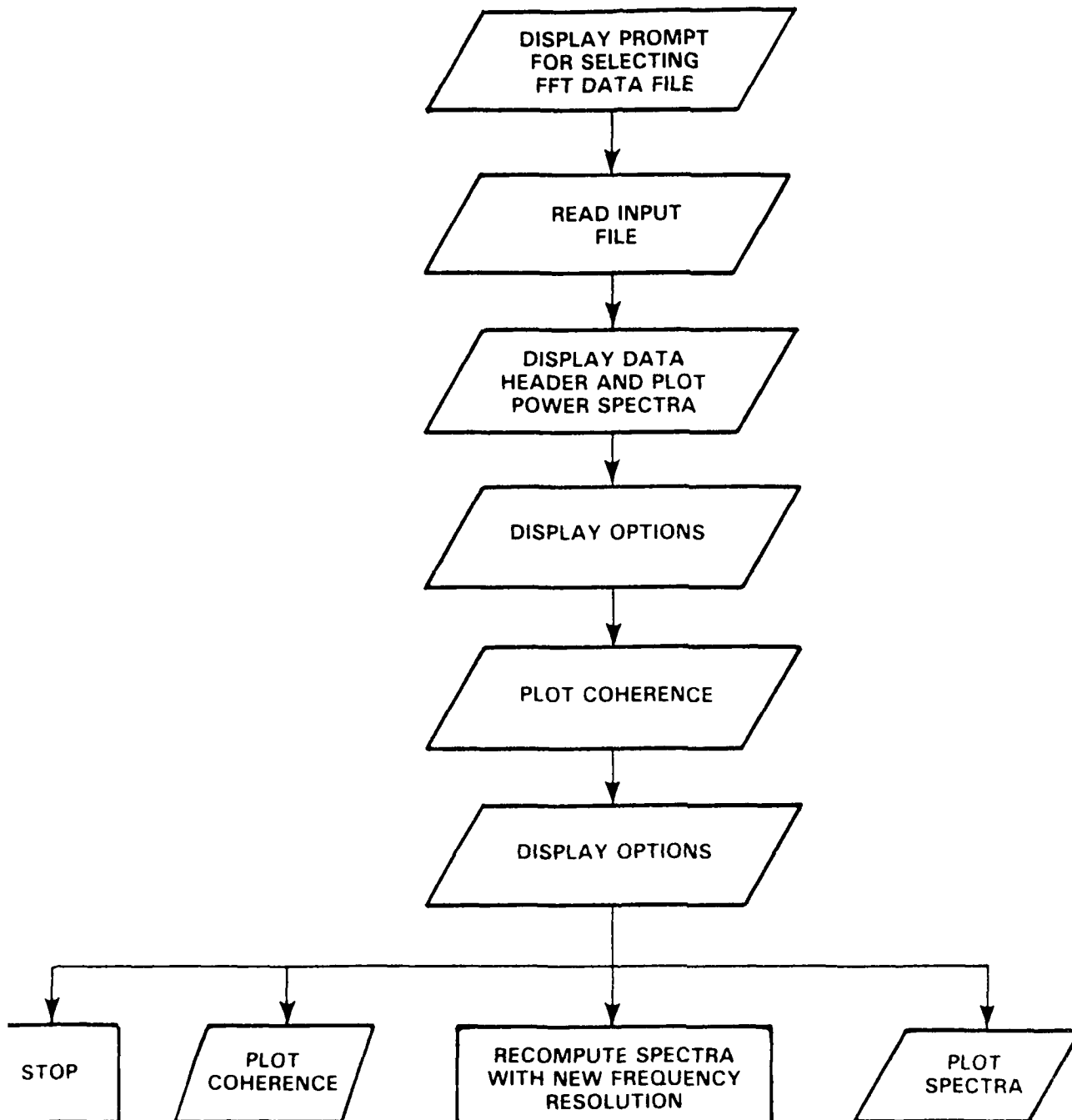


Figure 3.4-1 Block Flow Chart of PLOTFFT

3.7.16 Subroutine MMAXEN

CALLING SEQUENCE:

MMAXEN(N,ND,M,X,IA,V,A,B,PF,PB,Y,U,W,C,D,E,NE,AICVEC)

FUNCTION: Multivariate maximum entropy stepwise estimation of autoregression prediction matrices and one step prediction covariance matrix with order selection by Akaike's information criterion

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
N	I*4	-	Input	Sample size
ND	I*4	-	Input	Dimension of multivariate time series
M	I*4	-	Input	Maximum order autoregression searched for best fit
X	R*8	(*,*)	Input	Matrix of time series data
Y	R*8	(*,*)	Input	Work matrix with same dimension as X used for residuals from backward predictions. Residuals from forward prediction are stored in X. Original data are destroyed.
W	R*8	(*,*)	Input	Work matrix (ND,ND)
C	R*8	(*,*)	Input	Work matrix (ND,ND)
D	R*8	(*,*)	Input	Work matrix (ND,ND)
E	R*8	(*,*)	Input	Work matrix (NE,NE)
NE	I*4	-	Input	First dimension of matrix E in calling program (NE=ND*ND)
IA	I*4	-	Output	Order selected by Akaike's information criterion
V	R*8	(*,*)	Output	Estimated one step forward prediction covariance matrix

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.15 Subroutine JSUMRD

CALLING SEQUENCE:

JSUMRD(YS,XS,IA,X,A,N,HEADER,IHD,HDINFO,RDLAB)

FUNCTION: This routine calculates the sum-squared residuals

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
YS	R*4	(0:*,2)	Input	Plot work array
XS	R*4	(5000)	Input	Plot work array
IA	I*4	-	Input	Order selected by Akaike's information criterion
X	R*8	(*,2)	Input	Original data arrays
A	R*8	(2,2,40)	Input	Estimated partial forward auto-regression matrices
N	I*4	-	Input	Number of samples
HEADER	CHAR	-	Input	Plot header line
IHD	I*4	-	Input	Number of plot identification lines
HDINFO	CHAR	(5)	Input	Plot identification lines
RDLAB	CHAR	-	Input	Ordinate label for plots

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: DISPLA,AGSETP,ANOTAT,EZXY,PWRY,FRAME

COMMON BLOCKS:

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	R*8	(*,2)	Output	Contains data which were read in
N	I*4	-	Output	Number of data points in X
IFLAG	I*4	-	Output	= 0 - Successful read = 2 - Data set did not contain enough points

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.14 Subroutine JREMOV

CALLING SEQUENCE:

JREMOV(X,NSIZE,HDINFO,SLOPE)

FUNCTION: Removes the slope from both channels found in X

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	R*4	(5000,2)	Input	Raw data to have slopes removed
NSIZE	I*4	-	Input	Number of points in raw data
HDINFO	CHAR	(5)	Output	Slope information will be added to HDINFO
SLOPE	R*4	(2)	Output	Slope values returned to main program
X	R*4	(5000,2)	Output	Return data with slopes removed

3.7.12 Subroutine JPLOTL

CALLING SEQUENCE:

JPLOTL(X,Y,N,HEADER,HDINFO,IHD,RDLAB)

FUNCTION: Generates plots of the selected data

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	R*8	(*,2)	Input	Input double precision data arrays (N,2)
Y	R*4	(*,2)	Input	Input temporary single precision data arrays (N,2)
N	I*4	-	Input	Size of data arrays to be plotted
HEADER	CHAR	-	Input	Header associated with input data
HDINFO	CHAR	(5)	Input	Additional information carried along with the input data
IHD	I*4	-	Input	Number of HDINFO lines
RDLAB	CHAR	-	Input	Label along the ordinate axis

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: DISPLA,ANOTAT,AGSETP,EZY,PWRY,FRAME

3.7.13 Subroutine JREAD

CALLING SEQUENCE:

JREAD(X,N,IFLAG)

FUNCTION: Routine to read input data to the AR modeling program

3.7.11 Subroutine JPLOT

CALLING SEQUENCE:

JPLOT(S,NIN,F,IA,HEADER,IHD,HDINFO,ENDFLG,PSLABEL,AUTOFG)

FUNCTION: This routine plots the power spectrum for each channel of data and the coherence between channels

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
S	C*16	(2,2,*)	Input	Power spectral density matrix for each frequency bin
NIN	I*4	-	Input	Number of frequency bins
F	R*8	(*)	Input	Array of frequencies at which power spectral densities are calculated
IA	I*4	-	Input	Order selected by Akaike's information criterion
HEADER	CHAR	-	Input	Plot frame header
IHD	I*4	-	Input	Number of plot description lines
HDINFO	CHAR	(5)	Input	Plot description lines
PSLABEL	CHAR	-	Input	Ordinate axis label for power spectral density plot
AUTOFG	I*4	-	Input	= 1 - Use automatic scaling = 0 - Prompt for axis limits
ENDFLG	I*4	-	OUTPUT	Set by user to terminate the program

COMMON BLOCKS: WKSPAC

SUBPROGRAMS CALLED: AGSETP, INPRL4, DISPLA, ANOTAT, AGSETF, EZXY, PWRV, FRAME, INPIN4

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	R*8	(*,2)	Input	Contains original input data
NSIZE	I*4	-	Input	Number of points in input data set
X	R*8	(*,2)	Output	Raw data with slopes removed
HDINFO	CHAR	(5)	Output	Header lines with slope information added to be carried along with the data file

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.10 Subroutine JEQUAT

CALLING SEQUENCE:

JEQUAT(X,XX,NSIZE)

FUNCTION: This routine equates matrix XX to matrix X

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	R*8	(*,2)	Input	Input matrix(NSIZE,2)
NSIZE	I*4	-	Input	Size of matrix X and matrix XX
XX	R*8	(*,2)	Output	Equals matrix X upon return to calling program

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.8 Subroutine INVERT

CALLING SEQUENCE:

INVERT(B,N,ND)

FUNCTION: Inverts N by N matrix B in place; ND is first dimension of B in calling program

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
B	R*8	(*,*)	Input	Matrix to be inverted (N,N)
N	I*4	-	Input	Dimension of matrix B
ND	I*4	-	Input	First dimension of B in calling program
B	R*8	(*,*)	Output	Resultant inverted matrix

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.9 Subroutine J8REM

CALLING SEQUENCE:

J8REM(X,NSIZE,HDINFO)

FUNCTION: This routine removes slopes from both channels of input data X

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
IA	I*4	-	Input	Order selected by Akaike's information criterion
A	R*8	(2,2,40)	Input	Estimated partial forward autoregression matrices
SAMPT	R*8	-	Input	Sampling interval
AA	C*16	(2,2)	Input	Work array
V	R*8	(2,2)	Input	Estimated one step forward prediction covariance matrix
F	R*8	(*)	Input	Array of frequencies at which power spectral densities are to be calculated
NIN	I*4	-	Input	Number of frequency bins
S	C*16	(2,2,*)	Output	Power spectral density matrix for each frequency bin

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: CMZERO, CMI2X2, CMMULT, CMCTR, CMSAL

3.7.7 Subroutine DET

CALLING SEQUENCE:

DET(B,A,N,D)

FUNCTION: Evaluates determinant of an N by N matrix B

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
B	R*8	(*,*)	Input	N by N input matrix
A	R*8	(*,*)	Input	N by N work matrix
N	I*4	-	Input	Dimension of B and A
D	R*8	-	Output	Determinant of matrix B

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.5 Subroutine CMZERO

CALLIN SEQUENCE:

CMZERO (A,N1,N2)

FUNCTION: This routine zeroes out a complex matrix

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
A	C*16	(*,*)	Input	Complex matrix to be zeroed out (N1,N2)
N1	I*4	-	Input	First dimension of A
N2	I*4	-	Input	Second dimension of A

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.6 Subroutine CPOWER

CALLING SEQUENCE:

CPOWER(IA,A,SAMPT,AA,S,V,F,NIN)

FUNCTION: This routine calculates the power spectrum density matrix for each frequency bin requested

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	C*16	(*,*)	Input	First input complex matrix (N1,N2)
Y	C*16	(*,*)	Input	Second input complex matrix (N2,N3)
N1	I*4	-	Input	First dimension of X and Z
N2	I*4	-	Input	Second dimension of X, first dimension of Y
N3	I*4	-	Input	Second dimension of Y and Z
Z	C*16	(*,*)	Output	Output complex matrix (N1,N3)

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.4 Subroutine CMSCAL

CALLING SEQUENCE:

CMSCAL(X,Y,Z,N1,N2)

FUNCTION: This routine multiplies a complex matrix (Y) by a complex scalar (X) and returns the result in a complex matrix (Z)

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	C*16	-	Input	Input complex scalar
Y	C*16	(*,*)	Input	Input complex matrix (N1,N2)
N1	I*4	-	Input	First dimension of Y and Z
N2	I*4	-	Input	Second dimension of Y and Z
Z	C*16	(*,*)	Output	Output complex matrix

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.2 Subroutine CMI2X2

CALLING SEQUENCE:

CMI2X2(A,AINV)

FUNCTION: Inverts a complex 2 by 2 matrix

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
A	C*16	(2,2)	Input	Complex 2 by 2 matrix to be inverted
AINV	C*16	(2,2)	Output	Resultant 2 by 2 complex inverted matrix

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.7.3 Subroutine CMMULT

CALLING SEQUENCE:

CMMULT(X,Y,Z,N1,N2,N3)

FUNCTION: This program multiplies two complex matrices and
returns the result in a third matrix (Z)

REALVEC	Computes components of complex matrix periodogram spectrum for i-th batch of data
RSLOPE	Removes zero-mean least-squares linear trend from data
SMEAN	Subtracts mean from data array
SMFFTS	Reads, smoothes, and plots power spectra and coherence
SMPLOTS	Computes the smoothed estimates plus and minus standard errors
STDEV	Computes standard deviation of batch of periodograms at each frequency
STNDE	Computes standard errors from standard deviations
WEIGHT	Computes data-weighting window
XIO	Calculates IO(X)

3.7 GETDATA AND AR DETAILED SUBROUTINE DESCRIPTIONS

3.7.1 Subroutine CMCTR

CALLING SEQUENCE:

CMCTR(X,Y,N1,N2)

FUNCTION: Takes the complex conjugate transpose of X and places the result in Y

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	C*16	(*,*)	Input	Input complex array
N1	I*4	-	Input	First dimension of X, second dimension of Y
N2	I*4	-	Input	Second dimension of X, first dimension of Y
Y	C*16	(*,*)	Output	The complex conjugate transpose of X

JPLOT	Plots the power spectrum for each channel of data and the coherence between channels
JPLOT1	Generates plots of the selected data
JREAD	Reads input data to the AR modeling program
JREMOV	Removes the slope from both channels found in X
JSUMRD	Calculates the sum-squared residuals
MMAXEN	Multivariate maximum entropy stepwise estimation of autoregression prediction matrices and one step prediction covariance matrix with order selection by Akaike's information criterion
PRTINF	Prints out an informational line

3.6 FFT AND PLOTFFT SUBROUTINE DESCRIPTIONS

AVED	Computes arithmetic mean of spectral quantities SD(F1).....SD(F2)
AVEE	Computes RMS of standard errors SE(F1)...SE(F2)
FFTKW	Computes FFT of time series λ using mixed-radix algorithm
FFTMR	Compute FFT by mixed-radix algorithm
FMEAN	Computes mean 2 by 2 complex periodogram
KWINDO	Multiplies two vectors of data by Kaiser window
LOGAV	Smooths a spectrum using a constant-percent-bandwidth running mean
LXPRMT	Prompts the user for different plot options at the end of each frame
MSKPS	Computes coherence and cross-spectrum magnitude and phase
PLOT	Plots the power spectrum and coherence
PRTINF	Prints an informational line of text
RDATA	Reads and decimates raw data one batch at a time

```

@CL 'program name'    !compiles and loads into
                      PERIODLIB
@PLOTFFT              !links with main program and
                      creates new load module

```

Basic program information in a standard format is given below.

Name: PLOTFFT

Function: Plots power spectra and coherence from data files created by the program FFT

Common Blocks: *NONE*

Subprograms called: INPCHR, SMFFTS

3.5 GETDATA AND AR SUBROUTINE DESCRIPTIONS

CMCTR	Takes the complex conjugate transpose of X and places the result in Y
CM12X2	Inverts a complex 2 by 2 matrix
CMMULT	Multiplies two complex matrices and returns the result in a third matrix (Z)
CMSCAL	Multiplies a complex matrix (Y) by a complex scalar (X) and returns the result in a complex matrix (Z)
CMZERO	Zeroes out a complex matrix
CPOWER	Calculates the power spectral density matrix for each frequency bin requested
DET	Evaluates determinant of an N by N matrix B
INVERT	Inverts N by N matrix B in place; ND is first dimension of B in calling program
J8REM	Removes slopes from both channels of input data X
JEQUAT	Equates matrix XX to matrix X

A	R*8	(*,*,*)	Output	Estimated partial forward auto regression matrices
B	R*8	(*,*,*)	Output	Estimated partial backward autoregression matrices
PF	R*8	(*,*,*)	Output	Residual sum of squares and cross product matrices for each order fit. Last dimension of PF must be M+1 in calling program since order zero is stored in first position.
PB	R*8	(*,*,*)	Output	Same for backward prediction
U	R*8	(*,*)	Output	Estimated one step backward prediction covariance matrix
AICVEC	R*8	(*)	Output	Vector containing AICs for each model searched

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: DET, INVERT

3.7.17 Subroutine PRTINF

CALLING SEQUENCE:

PRTINF(LINE)

FUNCTION: Prints out a informational line

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
LINE	CHAR	-	Input	Line to be printed

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8 FFT AND PLOTFFT DETAILED SUBROUTINE DESCRIPTIONS

3.8.1 Subroutine AVED

CALLING SEQUENCE:

CALL AVED(F1,F2,SD,SMD,COUNT)

FUNCTION: Computes arithmetic mean of spectral quantities

SD(F1).....SD(F2)

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
F1	I*4	-	Input	Minimum frequency bin number
F2	I*4	-	Input	Maximum frequency bin number
SD	R*4	(*)	Input	Array of spectral quantities
COUNT	I*4	-	Input	Index value
SMD	R*4	(*)	Output	Arithmetic mean

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.2 Subroutine AVEE

CALLING SEQUENCE:

CALL AVEE(F1,F2,SE,SME,COUNT)

FUNCTION: Computes RMS of standard errors SE(F1)...SE(F2)

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
F1	I*4	-	Input	Minimum frequency bin number
F2	I*4	-	Input	Maximum frequency bin number
SE	R*4	(*)	Input	Array of standard errors
COUNT	I*4	-	Input	Index value
SME	R*4	(*)	Output	RMS of standard errors

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.3 Subroutine FFTKW

CALLING SEQUENCE:

CALL FFTKW(X,NFREQ,NSIZE,RS,IS,IM)

FUNCTION: Computes FFT of time series X using mixed-radix algorithm

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	R*8	(*)	Input	Array of time series data
NFREQ	I*4	-	Input	Number of frequency bins
NSIZE	I*4	-	Input	Number of data in time series
RS	R*8	(*)	Output	Array of real parts of FFT of X
IS	R*8	(*)	Output	Array of imaginary parts of FFT of X
IM	R*8	(*)	Output	Dummy variable array

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: FFTMR

3.8.4 Subroutine FFTMR

CALLING SEQUENCE:

CALL FFTMR(A,B,NTOT,N,NSPAN,ISN)

FUNCTION: Compute FFT by mixed-radix algorithm

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
A	R*8	(1)	Input	Real data array
B	R*8	(1)	Input	Imaginary data array
NTOT	I*4	-	Input	Number of complex data
N	I*4	-	Input	Dimension of current variable
NSPAN	I*4	-	Input	Algorithm parameter
ISN	I*4	-	Input	Sign of complex exponential argument
A	R*8	(1)	Output	Real data array
B	R*8	(1)	Output	Imaginary data array

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.5 Subroutine FMEAN

CALLING SEQUENCE:

CALL FMEAN(NFREQ,RS12,MRS12,IS12,MIS12,S11,MS11,
S22,MS22,A12,P12,K12,NBATCH,NSIZE,STEMP)

FUNCTION: Computes mean 2 by 2 complex periodogram

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NFREQ	I*4	-	Input	Number of frequency bins
RS12	R*8	(*,*)	Input	Matrix of real cross-spectra
IS12	R*8	(*,*)	Input	Matrix of imaginary cross-spectra
S11	R*8	(*,*)	Input	Matrix of channel 1 autospectra
S22	R*8	(*,*)	Input	Matrix of channel 2 autospectra
NBATCH	I*4	-	Input	Data batch number
NSIZE	I*4	-	Input	Number of data per batch
MRS12	R*8	(*)	Output	Array of mean real cross-spectrum
MIS12	R*8	(*)	Output	Array of mean imaginary cross-spectrum
MS11	R*8	(*)	Output	Array of mean channel 1 auto-spectrum
MS22	R*8	(*)	Output	Array of mean channel 2 auto-spectrum
A12	R*8	(*)	Output	Array of magnitude of mean complex cross-spectrum
P12	R*8	(*)	Output	Array of phase of mean complex cross-spectrum
K12	R*8	(*)	Output	Array of spectral coherence
STEMP	R*4	(*)	Output	Dummy array variable

COMMON BLOCKS: NNTEST, TRIGS

SUBPROGRAMS CALLED: *NONE*

3.8.6 Subroutine KWINDO

CALLING SEQUENCE:

CALL KWINDO(N,X1,X2,WT,ALPH)

FUNCTION: Multiplies two vectors of data by Kaiser window

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
N	I*4	-	Input	Number of data in each vector of data
X1	R*8	(*)	Input	First vector of data (array)
X2	R*8	(*)	Input	Second vector of data (array)
ALPH	R*8	-	Input	Kaiser window alpha parameter
WT	R*8	(*)	Output	Kaiser window (array)

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: WEIGHT

3.8.7 Subroutine LOGAV

CALLING SEQUENCE:

CALL LOGAV(NFREQ,FRACB,SD,SE,NCFREQ,CENTF,SMD,SME,NBATCH)

FUNCTION: Smoothes a spectrum using a constant-percent-bandwidth running mean

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NFREQ	I*4	-	Input	Number of frequency bins

FRACB	R*4	-	Input	Fractional band-width parameter
SD	R*4	(*)	Input	Estimated spectrum (array)
SE	R*4	(*)	Input	Standard errors (array)
NCFREQ	I*4	-	Output	Number of center frequencies
CENTF	R*4	(*)	Output	Center frequency bin numbers (array)
SMD	R*4	(*)	Output	Smoothed spectrum (array)
SME	R*4	(*)	Output	Smoothed standard errors (array)
NBATCH	I*4	-	Output	Number of data batches

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: AVED

3.8.8 Subroutine LXRMT

CALLING SEQUENCE:

CALL LXRMT(ICHG)

FUNCTION: Prompts the user for different plot options at the end of each frame

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
ICHG	I*4	-	Output	Contains the code for the desired option

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.9 Subroutine MSKPS

CALLING SEQUENCE:

CALL MSKPS(NFREQ,MRS12,MIS12,MS12,K12,PS12,MS11,MS22)

FUNCTION: Computes coherence and cross-spectrum magnitude and phase

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NFREQ	I*4	-	Input	Number of frequency bins
MRS12	R*8	(*)	Input	Mean real cross-spectrum array
MIS12	R*8	(*)	Input	Mean imaginary cross-spectrum array
MS11	R*8	(*)	Input	Mean autospectrum array for channel 1
MS22	R*8	(*)	Input	Mean autospectrum array for channel 2
MS12	R*8	(*)	Output	Magnitude of mean cross-spectrum array
K12	R*8	(*)	Output	Coherence array
PS12	R*8	(*)	Output	Phase array

COMMON BLOCKS: NNTEST

SUBPROGRAMS CALLED: *NONE*

3.8.10 Subroutine PLOT

CALLING SEQUENCE:

CALL PLOT(NFREQ,NCFREQ,FREQ,T1,T2,XMIN,XMAX,YMIN,YMAX,
LTYPE,NBATCH,IFIRST,NSTRT,HEADER,HDINFO,IHD,ENDFLG,NPLOTS)

FUNCTION: This routine plots the power spectrum and coherence

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NFREQ	I*4	-	Input	Number of frequency bins
NCFREQ	I*4	-	Input	Number of center frequencies
FREQ	R*4	(*)	Input	Array of frequencies
T1	R*4	(4097,3)	Input	Power spectrum or coherence array to be plotted
T2	R*4	(4097,3)	Input	Channel 2 power spectrum to be plotted
XMIN	R*4	-	Input	Default X-axis minimum
XMAX	R*4	-	Input	Default X-axis maximum
YMIN	R*4	-	Input	Default Y-axis minimum
YMAX	R*4	-	Input	Default Y-axis maximum
LTYPE	I*4	-	Input	Default X and Y axis type
NBATCH	I*4	-	Input	Number of batches of data
IFIRST	I*4	-	Input	Automatic scaling flag
NSTRT	I*4	-	Input	Starting data point to be plotted
HEADER	CHAR	-	Input	Plot header line
HDINFO	CHAR	(5)	Input	Plot descriptive information to be plotted at the bottom of the frame
IHD	I*4	-	Input	Number of lines of HDINFO
NPLOTS	I*4	-	Input	Flag for power spectrum or coherence plots
XMIN	R*4	-	Output	Current default X-axis minimum
XMAX	R*4	-	Output	Current X-axis maximum
YMIN	R*4	-	Output	Current Y-axis minimum

YMAX	R*4	-	Output	Current Y-axis maximum
LTYPE	I*4	-	Output	Current X and Y axis types
IFIRST	I*4	-	Output	Automatic scaling flag
ENDFLG	R*4	-	Output	Program termination flag

COMMON BLOCKS: WSPACE

SUBPROGRAMS CALLED: AGSETP, INPRL4, DISPLA, ANOTAT, AGSETF, AGSETI,
EZMXY, PWRY, FRAME

3.8.11 Subroutine PRTINF

CALLING SEQUENCE:

CALL PRTINF(LINE)

FUNCTION: Prints an informational line of text

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
LINE	CHAR	-	Input	Line to be printed

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.12 Subroutine RDATA

CALLING SEQUENCE:

CALL RDATA(X1, X2, NSIZE, NDEC, IERR, ENDFLG)

FUNCTION: Reads and decimates raw data one batch at a time

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NSIZE	I*4	-	Input	Size of batch to be read
NDEC	I*4	-	Input	Decimation factor
X1	R*8	(*)	Output	Channel 1 decimated data
X2	R*8	(*)	Output	Channel 2 decimated data
IERR	I*4	-	Output	Error during read flag
ENDFLG	L*1	-	Output	End of file flag

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.13 Subroutine REALVEC

CALLING SEQUENCE:

```
CALL REALVEC(NBATCH,RS1,RS2,IS1,IS2,S11,S22,RS12,IS12,
             NFREQ,I,NSIZE)
```

FUNCTION: Computes components of complex matrix periodogram spectrum for i-th batch of data

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NBATCH	I*4	-	Input	Number of batches
RS1	R*8	(*)	Input	Real part of channel 1 FFT array
RS2	R*8	(*)	Input	Real part of channel 2 FFT array
IS1	R*8	(*)	Input	Imaginary part of channel 1 FFT array
IS2	R*8	(*)	Input	Imaginary part of channel 2 FFT array
NFREQ	I*4	-	Input	Number of frequency bins
I	I*4	-	Input	Batch number

NSIZE	I*4	-	Input	Number of data per batch
S11	R*8	(*,*)	Output	Channel 1 autospectrum matrix
S22	R*8	(*,*)	Output	Channel 2 autospectrum matrix
RS12	R*8	(*,*)	Output	Real part of cross-spectrum matrix
IS12	R*8	(*,*)	Output	Imaginary part of cross-spectrum matrix

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.14 Subroutine RSLOPE

CALLING SEQUENCE:

CALL RSLOPE(NSIZE,X,SLOPE)

FUNCTION: Removes zero-mean least-squares linear trend from data

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NSIZE	I*4	-	Input	Number of data
X	R*8	(*)	Input	Data array
SLOPE	R*4	-	Output	Slope of linear trend
X	R*8	(*)	Output	Data array

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.15 Subroutine SMEAN

CALLING SEQUENCE:

CALL SMEAN(NSIZE,X)

FUNCTION: Subtracts mean from data array

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NSIZE	I*4	-	Input	Number of data
X	R*8	(*)	Input	Data array
X	R*8	(*)	Output	Data array with mean subtracted

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.16 Subroutine SMFFTS

CALLING SEQUENCE:

CALL SMFFTS(NFREQ,NBATCH,FSAMP,CENTF,FREQ,FRACB,
S1,SDS1,SES1,SMS1,SMES1,
S2,SDS2,SES2,SMS2,SMES2,
A12,SDA12,SEA12,SMA12,SMEA12,
P12,SDP12,SEP12,SMP12,SMEP12,
K12,SDK12,SEK12,SMK12,SMEK12,
TS1,TS2,TA12,TP12,TK12,HEADER,HDINFO,IHD)

FUNCTION: Reads, smoothes, and plots power spectra and
coherence

RGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NFREQ	I*4	-	Input	Number of frequency bins
NBATCH	I*4	-	Input	Number of batches of data
FSAMP	R*4	-	Input	Sampling frequency of time series data
CENTF	R*4	(*)	Input	Array of frequency bin numbers of center frequencies
FREQ	R*4	(*)	Input	Array of frequencies
FRACB	R*4	-	Input	Fractional bandwidth parameter
S1	R*4	(*)	Input	Autospectrum array for channel 1
SDS1	R*4	(*)	Input	Standard deviation of auto spectrum for channel 1 (array)
S2	R*4	(*)	Input	Autospectrum array for channel 2
SDS2	R*4	(*)	Input	Standard deviation of auto spectrum for channel 2 (array)
A12	R*4	(*)	Input	Magnitude of cross-spectrum (array)
SDA12	R*4	(*)	Input	Standard deviation of cross-spectrum (array)
P12	R*4	(*)	Input	Magnitude of phase (array)
SDP12	R*4	(*)	Input	Standard deviation of phase (array)
K12	R*4	(*)	Input	Magnitude of coherence (array)
SDK12	R*4	(*)	Input	Standard deviation of coherence (array)
TS1	R*4	(4097,3)	Input	Channel 1 overlay of SMS1, plus and minus standard error
TS2	R*4	(4097,3)	Input	Channel 2 overlay of SMS2, plus and minus standard error

TA12	R*4	(4097,3)	Input	Magnitude overlay of SMA12, plus and minus standard error
TP12	R*4	(4097,3)	Input	Phase overlay of SMP12, plus and minus standard error
TK12	R*4	(4097,3)	Input	Coherence overlay of SMK12, plus and minus standard error
HEADER	CHAR	-	Input	Plot header line
HDINFO	CHAR	(5)	Input	Plot data identification lines written at the bottom of all plots
IHD	I*4	-	Input	Number of HDINFO lines
SES1	R*4	(*)	Output	Standard errors of autospectrum for channel 1 (array)
SMS1	R*4	(*)	Output	Smoothed autospectrum array for channel 1 (array)
SMES1	R*4	(*)	Output	Standard errors of smoothed autospectrum for channel 1 (array)
SES2	R*4	(*)	Output	Standard errors of autospectrum for channel 2 (array)
SMS2	R*4	(*)	Output	Smoothed autospectrum array for channel 2 (array)
SMES2	R*4	(*)	Output	Standard errors of smoothed autospectrum for channel 2 (array)
SEA12	R*4	(*)	Output	Standard errors of cross-spectrum (array)
SMA12	R*4	(*)	Output	Smoothed magnitude of cross-spectrum (array)
SMEA12	R*4	(*)	Output	Standard errors of smoothed magnitude of cross-spectrum (array)

SEP12	R*4	(*)	Output	Standard errors of cross-spectrum (array)
SMP12	R*4	(*)	Output	Smoothed cross-spectrum phase (array)
SMEP12	R*4	(*)	Output	Standard errors of smoothed cross-spectrum phase (array)
SEK12	R*4	(*)	Output	Standard errors of coherence (array)
SMK12	R*4	(*)	Output	Smoothed coherence (array)
SMEK12	R*4	(*)	Output	Standard errors of smoothed coherence (array)

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: STNDE, LOGAV, SMPLOTS, INPRL4

3.8.17 Subroutine SMPLOTS

CALLING SEQUENCE:

```
CALL SMPLOTS(NFREQ, NCFREQ, NBATCH, CENTF, FSAMP, FREQ,
             SMES1, SMES2, SMEA12, SMEP12, SMEK12,
             TS1, TS2, TA12, TP12, TK12, HEADER, HDINFO, IHD, ENDFLG)
```

FUNCTION: Computes the smoothed estimates plus and minus standard errors

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NFREQ	I*4	-	Input	Number of frequency bins
NCFREQ	I*4	-	Input	Number of center frequencies
NBATCH	I*4	-	Input	Number of batches of data

CENTF	R*4	(*)	Input	Array of frequency bin numbers of center frequencies
FSAMP	R*4	-	Input	Sampling frequency of time series data
FREQ	R*4	(*)	Input	Array of frequencies
SMES1	R*4	(*)	Input	Standard errors of smoothed autospectrum for channel 1 (array)
SMES2	R*4	(*)	Input	Standard errors of smoothed autospectrum for channel 2 (array)
SMEA12	R*4	(*)	Input	Standard errors of smoothed magnitude of cross-spectrum (array)
SMEP12	R*4	(*)	Input	Standard errors of smoothed cross-spectrum phase (array)
SMEK12	R*4	(*)	Input	Standard errors of smoothed coherence (array)
TS1	R*4	(4097,3)	Input	Channel 1 overlay of SMS1, plus and minus standard error
TS2	R*4	(4097,3)	Input	Channel 2 overlay of SMS2, plus and minus standard error
TA12	R*4	(4097,3)	INPUT	Magnitude overlay of SMA12, plus and minus standard error
TP12	R*4	(4097,3)	INPUT	Phase overlay of SMP12, plus and minus standard error
TK12	R*4	(4097,3)	INPUT	Coherence overlay of SMK12, plus and minus standard error
HEADER	CHAR	-	INPUT	Plot header line
HDINFO	CHAR	(5)	INPUT	Plot data identification lines written at the bottom of all plots

IHD	I*4	-	INPUT	Number of HDINFO lines
ENDFLG	I*4	-	OUTPUT	Program termination flag

COMMON BLOCKS: WSPACE

SUBPROGRAMS CALLED: AGGETP,PLOT,INPIN4

3.8.18 Subroutine STDEV

CALLING SEQUENCE:

```
CALL STDEV(NFREQ,RS12,MRS12,IS12,MIS12,S11,MS11,S22,MS22,
          A12,SDS11,SDS22,SDA12,SDP12,SDK12,K12,NBATCH,SDIS12,
          SDRS12,STEMP)
```

FUNCTION: Computes standard deviation of batch of
periodograms at each frequency

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NFREQ	I*4	-	Input	Number of frequency bins
RS12	R*8	(*,*)	Input	Real parts of cross-spectra (matrix)
MRS12	R*8	(*)	Input	Mean of real cross-spectra (array)
IS12	R*8	(*,*)	Input	Imaginary parts of cross-spectra (array)
MIS12	R*8	(*)	Input	Mean imaginary cross-spectra (matrix)
S11	R*8	(*,*)	Input	Autospectra for channel 1 (matrix)
MS11	R*8	(*)	Input	Mean autospectrum for channel 1 (array)

S22	R*8	(*,*)	Input	Autospectra for channel 2 (matrix)
MS22	R*8	(*)	Input	Mean autospectra for channel 2 (array)
A12	R*8	(*)	Input	Magnitude of cross-spectrum (array)
K12	R*8	(*)	Input	Coherence (array)
NBATCH	I*4	-	Input	Number of data batches
STEMP	R*4	(*)	Input	Real * 4 disk storage temporary array
SDS11	R*8	(*)	Output	Standard deviations of channel 1 autospectra (array)
SDS22	R*8	(*)	Output	Standard deviations of channel 2 autospectra (array)
SDA12	R*8	(*)	Output	Standard deviations of cross-spectrum magnitude (array)
SDP12	R*8	(*)	Output	Standard deviation of cross-spectrum phase (array)
SDK12	R*8	(*)	Output	Standard deviations of coherence (array)
SDIS12	R*8	(*)	Output	Standard deviations of imaginary cross-spectra (array)
SDRS12	R*8	(*)	Output	Standard deviations of real cross-spectra (array)

COMMON BLOCKS: TRIGS, NNTEST

SUBPROGRAMS CALLED: *NONE*

3.8.19 Subroutine STNDE

CALLING SEQUENCE:

CALL STNDE(NFREQ, SD, SE, NBATCH)

FUNCTION: Computes standard errors from standard deviations

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
NFREQ	I*4	-	Input	Number of frequency bins
SD	R*4	(*)	Input	Array of standard deviations
NBATCH	I*4	-	Input	Number of batches of data
SE	R*4	(*)	Output	Array of standard errors

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.20 Subroutine WEIGHT

CALLING SEQUENCE:

CALL WEIGHT(WT,N,UF,SF,IOPT,ALPH)

FUNCTION: Computes data-weighting window

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
N	I*4	-	Input	Number of coefficients
IOPT	I*4	-	Input	Window selection parameter
ALPH	R*8	-	Input	Window parameter
WT	R*8	(1)	Output	Array of weighting coefficients
UF	R*8	-	Output	Sum of squares of coefficients
SF	R*8	-	Output	Scale factor

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

3.8.21 Function XIO

CALLING SEQUENCE:

RX = XIO(X)

FUNCTION: Calculates $IO(X) = 1 + \sum_{K=1}^{\infty} (((X/2)**K/FAC(K))**2)$
where FAC is factorial

ARGUMENTS:

NAME	TYPE	DIMENSION	IN/OUT	DESCRIPTION
X	R*8	-	Input	Function value to be evaluated
XIO	R*8	-	Output	Output of function

COMMON BLOCKS: *NONE*

SUBPROGRAMS CALLED: *NONE*

REFERENCES

1. White, J.V., "Error Models for Gravity Gradiometers in Airborne Surveys," Technical Report AFGL-TR-80-0220, Air Force Geophysics Laboratory, Hanscom AFB, January 1980. ADA097745.

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