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ON THE SPATIAL EIGENVALUE APPROACH
TO HIGH REYNOLDS NUMBER FLOW IN
A ROTATING AND NUTATING CYLINDER

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1. Introduction and formulation

It is well-known that there can be a significant difference in behavior in flight between liquid-filled and solid-filled projectiles. The difference is caused by the motion of the liquid inside the spinning projectile. This motion causes forces to act on the projectile which can ultimately cause the flight of the projectile to be prematurely terminated by instability. The initial motion of the projectile necessarily causes the fluid motion in the cylinder to be time-dependent; later it can be assumed that the flow is steady. It is the latter situation which is investigated here.

Our concern then is with the motion of an incompressible viscous fluid of kinematic viscosity ν in a cylinder of length $2Aa$ and radius a . The cylinder is spinning at a constant rate about its axis, which, in turn, is nutating at a constant rate and yaw angle about an axis directed along its trajectory. Throughout this calculation a will be used as an appropriate length scale while the magnitude of a typical velocity is taken to be $(\Omega + \bar{\tau} \cos K_0)a$, where Ω is the magnitude of the angular velocity of the projectile as observed in the aeroballistic reference frame, $\bar{\tau}$ is the dimensional coning frequency, and K_0 is the coning angle. (In the aeroballistic frame, one cartesian coordinate lies along the cylinder axis and a second coordinate lies in the plane of the cylinder axis and trajectory; see, e.g., Reference 1). K_0 is assumed small throughout this work, so that it is possible to determine the properties of the forced motion using a simple perturbation expansion in powers of K_0 . We nondimensionalize the coning frequency as follows:

$$\tau = \frac{\bar{\tau}}{(\Omega + \bar{\tau} \cos K_0)} \quad (1.1)$$

and we define the Reynolds number by

$$Re = \frac{(\Omega + \bar{\tau} \cos K_0)a^2}{\nu} \quad (1.2)$$

By use of the assumption already introduced, $K_0 \ll 1$, $\cos K_0$ is replaced by unity in the above and in the following. The typical time, length, velocity, and pressure scales used to nondimensionalize the flow variables are $(\Omega + \bar{\tau})^{-1}$, a , $(\Omega + \bar{\tau})a$, and $\rho a^2(\Omega + \bar{\tau})^2$, where ρ is the density of the fluid.

In a previous paper, Hall et al.^[2], hereafter referred to as HSG, we used the spatial eigenvalue approach to determine the forced motion in the projectile at finite values of the Reynolds number. Here we aim to extend the results of HSG to the high Reynolds number limit and show that in this case there is a significant simplification of the problem. We now indicate how the appropriate partial differential system to be solved can be derived

from the Navier-Stokes equations of motion. Our derivation is of course only brief so the reader is referred to HSG and the references therein for a more complete account.

First the Navier-Stokes equations are written down with respect to an inertial cylindrical polar co-ordinate system (r, θ, x) with corresponding velocity components (u, v, w) and pressure p . The coning motion of the projectile means that the boundary conditions must be applied at a moving surface; this is the main drawback in the use of the inertial reference frame but for small angles a Taylor series expansion of the different flow quantities about the mean position of the cylinder produces boundary conditions at a fixed surface. Thus it is assumed that the motion of the cylinder is proportional to $\exp i\{ft - \theta\}$ where t denotes nondimensional time and f is, in general, the nondimensional complex frequency of the projectile motion. We shall take f to be real so that

$$f = \tau. \quad (1.3)$$

Since K_0 , the coning angle, has been assumed small we expect only small departures from solid body rotation and the velocity components are therefore $(-K_0 u^*, r - K_0 v^*, -K_0 w^*)$ and the pressure is $\frac{1}{2}r^2 - K_0 p^*$. Here a star denotes a perturbation quantity.

It is assumed in our calculation that the motion of the cylinder is specified; we do not allow for a feedback from the fluid to the projectile. Following HSG it is convenient to define complex velocities and pressures, denoted by underlining and given by

$$(u^*, v^*, w^*, p^*) = \text{Real}\{(\underline{u}, \underline{v}, \underline{w}, \underline{p}) \exp i\{ft - \theta\}\}. \quad (1.4)$$

The resulting nondimensional Navier-Stokes equations are

$$\begin{aligned} i(f-1)\underline{u} - 2\underline{v} &= -\underline{p}_r + Re^{-1}[\nabla^2 \underline{u} - \frac{2}{r^2}\underline{u} + \frac{2i}{r^2}\underline{v}], \\ i(f-1)\underline{v} + 2\underline{u} &= (i/r)\underline{p} + Re^{-1}[\nabla^2 \underline{v} - \frac{2}{r^2}\underline{v} - \frac{2i}{r^2}\underline{u}], \\ i(f-1)\underline{w} &= -\underline{p}_x + Re^{-1}[\nabla^2 \underline{w} - \frac{1}{r^2}\underline{w}], \\ (r\underline{u})_r - i\underline{v} + r\underline{w}_x &= 0, \end{aligned}$$

where $\nabla^2 = \partial_r^2 + \frac{1}{r}\partial_r + \partial_x^2$ and subscripts denote partial derivatives.

We now choose to seek separable solutions of these equations which take the form

$$(\underline{u}, \underline{v}, \underline{w}, \underline{p}) = (U(r) \sin Kx, V(r) \sin Kx, W(r) \cos Kx, P(r) \sin Kx), \quad (1.5)$$

and in this case it follows that (U, V, W, P) satisfies

$$\begin{aligned} [Re^{-1}(\nabla_1 - r^{-2}) - iM]U + 2[1 + (i/r^2 Re)]V - P_r &= 0 \\ [Re^{-1}(\nabla_1 - r^{-2}) - iM]V - 2[1 + (i/r^2 Re)]U + \frac{iP}{r} &= 0 \\ [Re^{-1}\nabla_1 - iM]W - KP &= 0 \\ (rU)_r - iV - KrW &= 0. \end{aligned} \quad (1.6a, b, c, d)$$

where $M = f - 1$

and $\nabla_1 \equiv d_{rr} + \frac{1}{r}d_r - [\frac{1}{r^2} + K^2]$.

These equations must be solved subject to an inhomogeneous condition at $r = 1$ and an appropriate condition at $r = 0$. We obtain

$$U - iV = W = P = 0, \quad r = 0 \quad (1.7)$$

at the centre of the cylinder. For a discussion of the origin of (1.7) the reader is referred to Batchelor and Gill^[3]. The ordinary differential system specified by (1.6), (1.7) and the no-slip condition at the cylinder constitutes an eigenvalue problem

$$K = K(Re, f). \quad (1.8)$$

At finite values of Re there are no real solutions of that eigenvalue problem so that solid body rotation is stable to the first non-axisymmetric Taylor vortex mode of instability. At infinite values of Re solid body rotation is, according to Rayleigh's criterion, neutrally stable and this is the origin of the real values of K which that eigenvalue problem has in the limit $Re \rightarrow \infty$. We restrict our attention to the case when the pivot point for the cylinder is midway along its axis. From Gerber et al.^[4] we know that the forced problem at small amplitudes then leads to the following conditions for $\underline{u}$, $\underline{v}$ and $\underline{w}$ at the walls:

$$\begin{aligned} \underline{u} &= -i(1 - f)x, \\ \underline{v} &= -(1 - f)x, \\ \underline{w} &= i(1 - f)r. \end{aligned} \quad (1.9a, b, c)$$

Following HSG we transfer the inhomogeneous boundary conditions to the endwall by writing

$$\begin{aligned} \underline{u} &= -i[1 - f]x + u(r, x), \\ \underline{v} &= -[1 - f]x + v(r, x), \\ \underline{w} &= i[1 - f]r + \phi(r) + w(r, x), \\ \underline{p} &= -[1 - f^2]rx + p(r, x), \\ \phi(r) &= 2if \left[r - \frac{J_1(\lambda r)}{J_1(\lambda)} \right]. \end{aligned} \quad (1.10a, b, c, d, e)$$

Here J_1 denotes the order 1 Bessel function of the first type while λ is defined by

$$\lambda = (1 + i)\sqrt{(1 - f)Re/2}. \quad (1.11)$$

Since we will be interested in the limit of high Reynolds numbers we can drop the Bessel function terms in (1.10) away from $r = 1$. We note in passing here that λ , and hence u etc. have a branch cut at $f = 1$; we will not investigate such high frequencies here since the ballistic applications of our work typically correspond to values of f less than .5. In a separate study eigenvalues and the pressure coefficient c_p were calculated for $1 < f \leq 1.5$. The results were to be compared with experimental data for $f > 1$. In terms of u, v, w and p it follows from above that the appropriate boundary conditions are

$$\begin{aligned} u - iv = w = p = 0, \quad r = 0 \\ u = v = w = 0, \quad r = 1 \\ u = v = 0, w = -\phi(r) \quad x = \pm A. \end{aligned} \quad (1.12a, b, c)$$

Furthermore, we note that u, v, w, p satisfy (1.5) and (1.6) and following the spatial eigenvalue approach of HSG we therefore write

$$\begin{aligned} u = \sum_{n=1}^{\infty} \frac{\alpha_n \sin k_n x}{\sin k_n A} u_n(r), \quad v = \sum_{n=1}^{\infty} \frac{\alpha_n \sin k_n x}{\sin k_n A} v_n(r), \\ w = \sum_{n=1}^{\infty} \frac{\alpha_n \cos k_n x}{\sin k_n A} w_n(r), \quad p = \sum_{n=1}^{\infty} \frac{\alpha_n \sin k_n x}{\sin k_n A} p_n(r). \end{aligned} \quad (1.13a, b, c, d)$$

Here $\{k_n\}$ is the infinite sequence of eigenvalues of (1.6), (1.7) and the coefficients $\{\alpha_n\}$ are to be determined. In HSG these constants were found by collocation at the endwalls; here we shall make the further limit $Re \rightarrow \infty$ and show how asymptotic expressions for these coefficients can be found.

2. The spatial eigenvalue approach in the limit $Re \rightarrow \infty$.

On the basis of our experience with the eigenvalue problem for $\{k_n\}$ at finite values of Re we know that the eigenvalues split into groups of three which can be ordered by, for example, connecting them to the eigenvalues of a related Taylor vortex problem. We anticipate this split and therefore consider the three sets of eigenvalues $\{k_{j,n}\}$, for $j = 1, 2, 3$ and $n = 1, 2, 3, \dots$. The distribution of eigenvalues is illustrated in Figure 1 for the case $Re = 1000$ and $f = .1$; this is taken from HSG wherein the labeling of the eigenvalues is discussed. (Note that the eigenvalue $k_{3,1}$ was labeled on the basis of an analysis pertinent to the limit $Re \rightarrow 0$. However, in the limit $Re \rightarrow \infty$, it lies in the set of values determined

by Stewartson's analysis^[5].) In Figure 2 we show the development of the spectrum as Re increases, for fixed f and $j = 1$. We note that for sufficiently small n the eigenvalues approach the real axis while for large n the eigenvalues approach the imaginary axis.

On the other hand numerical calculations in HSG showed that the $j = 2, 3$ branches had $|k_{jn}| \sim Re^{\frac{1}{2}}$ for $Re \gg 1$. We shall now describe the precise details of the asymptotic structure of $\{k_{jn}\}$ for $j = 1, 2, 3$ in the limit $Re \rightarrow \infty$ and hence put ourselves into a position to find asymptotic forms for $\{\alpha_n\}$. Splitting of the eigenvalues into triplets leads naturally to the definitions of branches 1, 2 and 3 in the complex k plane for $j = 1, 2, 3$ respectively. The corresponding amplitude functions are denoted by ϵ_n, γ_n and β_n , replacing the α_n .

First we derive the asymptotic structure of the $j = 1$ modes; the analysis is, of course, essentially inviscid and based on that given by Stewartson^[5]. We seek an asymptotic solution for the n^{th} eigenvalue k_{1n} by writing

$$k_{1n} = k_{1n}^0 + k_{1n}^1 Re^{-\frac{1}{2}} + k_{1n}^2 Re^{-1} + \dots, \quad (2.1)$$

while

$$u_{1n} = u_{1n}^0 + u_{1n}^1 Re^{-\frac{1}{2}} + u_{1n}^2 Re^{-1} + \dots, \quad (2.2)$$

together with similar expansions for v_{1n}, w_{1n} and p_{1n} . If we define σ by

$$\sigma = i\{f - 1\}$$

then substitution of (2.1), (2.2) into (1.6) with K replaced by k_{1n} yields a system of equations for $u_{1n}^0, v_{1n}^0, w_{1n}^0, p_{1n}^0$ which has the solution

$$\begin{aligned} \{\sigma^2 + 4\}u_{1n}^0 &= i \left\{ (1-f)\mu_{1n}^0 J_1'(\mu_{1n}^0 r) + \frac{2J_1(\mu_{1n}^0 r)}{r} \right\}, \\ \{\sigma^2 + 4\}v_{1n}^0 &= \left\{ (1-f)\frac{J_1(\mu_{1n}^0 r)}{r} + 2\mu_{1n}^0 J_1'(\mu_{1n}^0 r) \right\}, \\ w_{1n}^0 &= \frac{-k_{1n}^0}{\sigma} J_1(\mu_{1n}^0 r), \\ p_{1n}^0 &= J_1(\mu_{1n}^0 r). \end{aligned} \quad (2.3a, b, c, d)$$

$$(\mu_{1n}^0)^2 = (k_{1n}^0)^2 \frac{(3-f)(1+f)}{(1-f)^2}. \quad (2.4)$$

The radial velocity component above satisfies the sidewall condition if

$$(1-f)\mu_{1n}^0 J_1'(\mu_{1n}^0) + 2J_1(\mu_{1n}^0) = 0 \quad (2.5)$$

which has an infinite number of eigenvalues on the positive real axis. [Information on this spectrum can be found in 'Conduction of Heat in Solids' by H.S. Carslaw and J.C. Jaeger, Oxford University Press, 1947.] The eigenvalues with $j = 1$ become progressively closer to the real axis in the limit $Re \rightarrow \infty$. The Stewartson eigenvalues correspond to a choice of frequency such that $k_{1n}^0 A = \frac{\pi}{2} \{1 + N\}$ for N an even non-negative integer. Thus, at the resonant frequencies of a cylinder of given length, the $j = 1$ eigenvalues approach Stewartson's eigenvalues. Since we have not retained viscous effects in the perturbation equations, we can, of course, only make the radial velocity component vanish at $r = 1$. The other velocity components must therefore be reduced to zero within a boundary layer of thickness $Re^{-\frac{1}{2}}$ near $r = 1$. The determination of the velocity field within this layer will enable us to determine the order $Re^{-\frac{1}{2}}$ term in (2.1), we shall give some of the detail necessary to determine it.

First we note that the perturbation equations for $r = 0(1)$ proceed in powers of Re so that if we define μ_{1n} by

$$\mu_{1n} = \mu_{1n}^0 + \mu_{1n}^1 Re^{-\frac{1}{2}} + \dots, \quad (2.6)$$

then correct to order $Re^{-\frac{1}{2}}$, (2.3) are valid if $u_{1n}^0, v_{1n}^0, w_{1n}^0, p_{1n}^0$ and μ_{1n}^0 are replaced by $u_{1n}^0 + u_{1n}^1 Re^{-\frac{1}{2}}$, etc. and then

$$\mu_{1n}^0 \mu_{1n}^1 = k_{1n}^0 k_{1n}^1 \frac{(3-f)(1+f)}{(1-f)^2}. \quad (2.7)$$

It follows that when $r \rightarrow 1$

$$(3-f)(1+f)u_{1n}^1 \rightarrow -\mu_{1n}^1 i \left\{ \mu_{1n}^0 J_1(\mu_{1n}^0)(1-f) \left[1 + \frac{(3-f)(1+f)}{(\mu_{1n}^0)^2(1-f)^2} \right] \right\} + C\zeta, \quad (2.8)$$

where $\zeta = Re^{\frac{1}{2}}(1-r)$, and C is a constant to be determined later. In the boundary layer at $r = 1$ we write

$$Re^{\frac{1}{2}}u_{1n} = \tilde{u}_{1n}^0 + \tilde{u}_{1n}^1 Re^{-\frac{1}{2}} + \dots$$

together with similar expansions for v_{1n}, w_{1n} , and p_{1n} . It can be shown that the zeroth order system obtained by substituting these expansions into (1.6) written in terms of ζ yields

$$\tilde{p}_{1n}^0 = \text{constant}, \quad \tilde{u}_{1n}^0 = \left\{ \frac{ip_{1n}^0}{1-f} \right\} \left\{ 1 + (k_{1n}^0)^2 \right\} \left\{ \frac{1}{\sqrt{\sigma}} e^{-\sqrt{\sigma}\zeta} - \frac{1}{\sqrt{\sigma}} + \zeta \right\}$$

when solved subject to $\tilde{u}_{1n}^0 = \tilde{v}_{1n}^0 = \tilde{w}_{1n}^0 = 0$, $\zeta = 0$. Thus matching the pressure and radial velocity component when the inviscid and viscous regions meet give

$$\tilde{p}_{1n}^0 = J_1(\mu_{1n}^0), \quad \frac{iJ_1(\mu_{1n}^0)[1 + (k_{1n}^0)^2]}{(1-f)\sigma^{\frac{1}{2}}} = \frac{i\mu_{1n}^0 J_1(\mu_{1n}^0)(1-f)\mu_{1n}^1}{(3-f)(1+f)} \left\{ 1 + \frac{(3-f)(1+f)}{(\mu_{1n}^0)^2(1-f)^2} \right\},$$

and an expression for C the constant in (2.8). Using (2.7) and the eigenrelation (2.5) the second of the above equations can be manipulated to give

$$k_{1n} = k_{1n}^0 \left\{ 1 + \frac{1+i}{\sqrt{1-f}\sqrt{2Re}} + \dots \right\} \quad (2.9)$$

where k_{1n}^0 is determined by (2.4) and (2.5).

The viscous modes have an x -dependence on a $Re^{-\frac{1}{2}}$ lengthscale. For convenience we drop the '2n' and '3n' notation for the moment and seek a viscous eigenvalue of (1.6) and (1.7) by writing

$$k = \sqrt{Re} k_0 + \frac{1}{\sqrt{Re}} k_1 + \dots \quad (2.10)$$

while u, v, w , and p expand as

$$\begin{aligned} u &= u_0 + \frac{1}{\sqrt{Re}} u_1 + \dots, \\ v &= v_0 + \frac{1}{Re^{\frac{1}{2}}} v_1 + \dots, \\ w &= \frac{w_0}{\sqrt{Re}} + \frac{w_1}{Re} + \dots, \\ p &= \frac{p_0}{Re} + \frac{p_1}{Re^{\frac{3}{2}}} + \dots. \end{aligned} \quad (2.11a, b, c, d)$$

The zeroth order approximation to (1.6) then yields

$$\begin{aligned} \{-k_0^2 - if + i\}u_0 + 2v_0 &= 0, \\ \{-k_0^2 - if + i\}v_0 - 2u_0 &= 0, \\ w_0 &= \frac{du_0}{dr} / k_0. \end{aligned} \quad (2.12a, b, c)$$

The equations (2.12a,b) are consistent if

$$\begin{aligned} k_0^2 &= (3-f)i \\ \text{or } k_0^2 &= -(1+f)i \end{aligned} \quad (2.13a, b)$$

which correspond to the $j = 3$ and $j = 2$ modes respectively. At higher order u_0 , and hence v_0, w_0 and p_0 , are determined as solvability conditions. We find that (2.13a,b) respectively lead to

$$u_0 = J_2(s_n r), \quad u_0 = J_0(t_n r) \quad (2.14a, b)$$

where

$$J_2(s_n) = J_0(t_n) = 0.$$

A more detailed derivation of (2.14) is found in HSG. Thus (2.13a,b) correspond to two countably infinite sets of eigenvalues and $\{s_n\}, \{t_n\}$ determine the order $Re^{-\frac{1}{2}}$ correction terms in the expansion of k . In fact the appropriate equations are found to be

$$k_1 = \frac{-(1-i)(4-f)s_n^2}{2\sqrt{2(3-f)^3}} \quad (2.15)$$

and

$$k_1 = \frac{-(2+f)t_n^2(1+i)}{2\sqrt{2}(1+f)^{3/2}}. \quad (2.16)$$

Henceforth we refer to the eigenvalues corresponding to (2.13a,b) using the indices $j = 3$ and $j = 2$ respectively. The structure outlined above breaks down when n is formally $O(Re)$ in which case the eigenfunction has both axial and radial variations on the $Re^{-\frac{1}{2}}$ lengthscale. We can now modify the spatial eigenvalue expansions (1.13a,b,c,d) using the asymptotic structure developed above; we therefore write:

$$\begin{aligned} u &= \sum_1^\infty \left[\frac{\epsilon_n i}{\{3-f\}\{1+f\}} \left\{ \frac{(f+1)J_1(\mu_{1n}^0 r)}{r} + (1-f)\mu_{1n}^0 J_0(\mu_{1n}^0 r) \right\} \frac{\sin k_{1n}^0 x}{\sin k_{1n}^0 A} \right. \\ &\quad \left. + \gamma_n J_0(t_n r) \frac{\sin \lambda_2 x}{\sin \lambda_2 A} + \beta_n J_2(s_n r) \frac{\sin \lambda_1 x}{\sin \lambda_1 A} \right] + \dots, \\ v &= \sum_1^\infty \left[\frac{\epsilon_n}{\{3-f\}\{1+f\}} \left\{ -(1+f) \frac{J_1(\mu_{1n}^0 r)}{r} + 2\mu_{1n}^0 J_0(\mu_{1n}^0 r) \right\} \frac{\sin k_{1n}^0 x}{\sin k_{1n}^0 A} \right. \\ &\quad \left. - i\gamma_n J_0(t_n r) \frac{\sin \lambda_2 x}{\sin \lambda_2 A} + i\beta_n J_2(s_n r) \frac{\sin \lambda_1 x}{\sin \lambda_1 A} \right] + \dots, \\ w &= \sum_1^\infty \frac{\epsilon_n i k_{1n}^0}{(f-1)} J_1(\mu_{1n}^0 r) \frac{\cos k_{1n}^0 x}{\sin k_{1n}^0 A} + \dots, \\ p &= \sum_1^\infty \epsilon_n J_1(\mu_{1n}^0 r) \frac{\sin k_{1n}^0 x}{\sin k_{1n}^0 A} + \dots. \end{aligned} \quad (2.17a, b, c, d)$$

where

$$\lambda_1 = (1+i)\sqrt{\frac{3-f}{2}}\sqrt{Re}, \quad \lambda_2 = (1-i)\sqrt{\frac{1+f}{2}}\sqrt{Re}.$$

We now apply the endwall condition (1.12c) on the axial velocity component w to give

$$-2if r = \sum_1^{\infty} \frac{\epsilon_n i k_{1n}^0}{(f-1)} J_1(\mu_{1n}^0 r) \cot k_{1n}^0 A$$

so that after multiplying both sides of the equation by $r J_1(\mu_{1n}^0 r)$, integrating from $r = 0$ to $r = 1$ and using the well-known properties of Bessel functions we obtain

$$\epsilon_n = \frac{4f(3-f) \tan k_{1n}^0 A}{k_{1n}^0 (\mu_{1n}^0)^2 \left\{ 1 + \frac{1}{(k_{1n}^0)^2} \right\} J_1(\mu_{1n}^0)} \quad (2.18)$$

while the endwall conditions on the radial and circumferential velocity components yield equations which can be solved for $\{\beta_n\}$ and $\{\gamma_n\}$ in terms of $\{\epsilon_n\}$. It is at this stage that the differences between the finite and infinite Reynolds number have emerged. Thus the amplitude functions $\{\epsilon_n\}$, $\{\beta_n\}$ and $\{\gamma_n\}$ in the limit Re approaches ∞ can be found explicitly using orthogonality properties of Bessel functions. At finite values of Re HSG found it necessary to use either collocation or a least squares approach in order to find the corresponding amplitudes. At higher order in the procedure that we have begun above, the sidewall boundary layer comes into play at the endwalls; at that stage the higher order amplitude coefficients must be found by collocation or the least squares method.

It follows from (2.17d) and (2.18) that p on the endwall, $x = A$, can be expressed as the infinite series

$$p = 4f(3-f) \sum_1^{\infty} \frac{k_{1n}^0 \tan k_{1n}^0 A J_1(\mu_{1n}^0 r)}{(\mu_{1n}^0)^2 J_1(\mu_{1n}^0) [1 + (k_{1n}^0)^2]}, \quad (2.19)$$

which is correct up to order Re^0 .

The pressure coefficient

We shall now use the high Reynolds number form of the spatial eigenvalue approach to find the pressure exerted by the fluid on the container. Since almost all of the available experimental results for pressure correspond to measurements made at the endwall we shall concentrate on that situation. More precisely we use our theory to find the amplitude of the pressure measured by a pressure transducer located at the endwall. More detail of the derivation of this pressure coefficient can be found in Gerber et al.^[4]

First we note that p^+ , the pressure calculated from the Navier-Stokes equations is

$$p^+ = \frac{1}{2} r^2 - K_0 p^*,$$

in the case $K_0 \rightarrow 0$. We must now relate our calculations in an inertial frame to the frame $\tilde{r}, \tilde{\theta}, \tilde{x}$ used in HSG where the $\tilde{x}$ axis coincides with the cylinder axis. A routine calculation based on a Taylor series expansion shows that, for example

$$\frac{1}{2}r^2 = \frac{1}{2}\tilde{r}^2 - K_0\tilde{r}\tilde{x}\cos\{ft - \tilde{\theta}\} + O(K_0^2).$$

We now let Δp be the disturbance pressure not including the contribution from solid body rotation. At points fixed on the surface of the cylinder

$$\Delta p = \frac{1}{2}r^2 - K_0 p^*(\tilde{r}, \tilde{\theta}, \tilde{x}, t) - \frac{1}{2}\tilde{r}^2$$

and since $p^*(\tilde{r}, \tilde{\theta}, \tilde{x}, t) - p^*(r, \theta, x, t) = O(K_0)$ it follows that

$$\Delta p = -K_0[-\underline{p}_i \sin(ft - \theta) + (\underline{p}_r + rx) \cos(ft - \theta)] + O(K_0^2).$$

Here subscripts r and i denote the real and imaginary parts of a complex quantity. Now since

$$\underline{p}_r + rx = p_r + f^2 rx, \quad \underline{p}_i = p_i$$

the pressure coefficient $c_p = \frac{|\Delta p|}{K_0}$ may be written as

$$c_p = \sqrt{(\underline{p}_r + rx)^2 + p_i^2}. \quad (2.20)$$

It then follows from (2.19) and (2.20) that c_p calculated using the spatial eigenvalue approach in the limit $Re \rightarrow \infty$ is given by

$$c_p = \left| \frac{4f(1-f)^2}{1+f} \sum_1^\infty \frac{\tan k_{1n}^0 A J_1(\mu_{1n}^0 r)}{k_{1n}^0 [1 + (k_{1n}^0)^2] J_1(\mu_{1n}^0)} + f^2 r A \right| + O(Re^{-\frac{1}{2}}). \quad (2.21)$$

The above equation is not uniformly valid since at the Stewartson eigenfrequencies the unperturbed state can support neutral eigenfunctions. However, it is clear that c_p will become unbounded if k_{1n}^0 corresponds to one of the Stewartson eigenvalues in which case $\tan k_{1n}^0 A = \infty$. In fact our expansion procedure is readily modified to take account of this possibility. Suppose then that f is an eigenfrequency $\bar{f}$ of the inviscid problem. Without any loss of generality we can suppose that $k_{11}^0 A = \pi/2$ so that ϵ_1 is formally infinite. In order to remove this singularity we note from (2.18) that if higher order terms are retained then ϵ_1 is in fact $O(Re^{\frac{1}{2}})$ when $f = \bar{f}$. Since $k_{1n}^0 = k_{1n}^0(f)$ and we wish to obtain a correct zeroth order approximation to c_p for all values of f we must therefore

modify (2.17) for $f - \bar{f} \sim 0(Re^{-\frac{1}{2}})$. Thus (2.21) is the correct zeroth order approximation to c_p for $f - \bar{f} \gg 0(Re^{-\frac{1}{2}})$; we refer to this situation as the nonresonant case.

It follows from our calculations above for the nonresonant case by taking the limit $f \rightarrow \bar{f}$ that $\epsilon_1, \{\beta_n\}, \{\gamma_n\}$ for all n becomes formally $0(Re^{+\frac{1}{2}})$ when $f - \bar{f} = 0(Re^{-\frac{1}{2}})$.

We therefore make use of the new scalings inferred above and therefore replace (2.17) by

$$\begin{aligned}
u &= \frac{i\epsilon_1\sqrt{Re}}{(3-f)(1+f)} \left\{ \frac{(1+f)}{r} J_1(\mu_{11}^0 r) + (1-f)\mu_{11}^0 J_0(\mu_{11}^0 r) + \dots \right\} \frac{\sin k_{11}x}{\sin k_{11}A} \\
&+ \sum_2^\infty \frac{\epsilon_n i}{(3-f)(1+f)} \left\{ \frac{1+f}{r} J_1(\mu_{1n}^0 r) + (1-f)\mu_{1n}^0 J_0(\mu_{1n}^0 r) + \dots \right\} \frac{\sin k_{1n}^0 x}{\sin k_{1n}^0 A} \\
&+ \sqrt{Re} \sum_1^\infty \gamma_n J_0(t_n r) \frac{\sin \lambda_2 x}{\sin \lambda_2 A} + \dots \\
&+ \sqrt{Re} \sum_1^\infty \beta_n J_2(s_n r) \frac{\sin \lambda_1 x}{\sin \lambda_1 A} + \dots \\
v &= \frac{\epsilon_1\sqrt{Re}}{(3-f)(1+f)} \left\{ -(1+f) \frac{J_1(\mu_{11}^0 r)}{r} + 2\mu_{11}^0 J_0(\mu_{11}^0 r) + \dots \right\} \frac{\sin k_{11}x}{\sin k_{11}A} \\
&+ \sum_2^\infty \frac{\epsilon_n}{(3-f)(1+f)} \left\{ \frac{-(f+1)J_1(\mu_{1n}^0 r)}{r} + 2\mu_{1n}^0 J_0(\mu_{1n}^0 r) + \dots \right\} \frac{\sin k_{1n}^0 x}{\sin k_{1n}^0 A} \\
&+ \sqrt{Re} \sum_1^\infty i\beta_n J_2(s_n r) \frac{\sin \lambda_1 x}{\sin \lambda_1 A} + \dots \\
&- \sqrt{Re} \sum_1^\infty i\gamma_n J_0(t_n r) \frac{\sin \lambda_2 x}{\sin \lambda_2 A} + \dots \\
w &= \frac{ik_{11}^0}{(f-1)} \epsilon_1 \sqrt{Re} [J_1(\mu_{11}^0 r) + \dots] \frac{\cos k_{11}x}{\sin k_{11}A} \\
&+ \sum_2^\infty \frac{ik_{1n}^0}{(f-1)} \epsilon_n [J_1(\mu_{1n}^0 r) + \dots] \frac{\cos k_{1n}^0 x}{\sin k_{1n}^0 A} \\
&+ \sum_1^\infty \frac{s_n \beta_n J_2'(s_n r)(1-i)}{\sqrt{3-f}\sqrt{2}} \frac{\cos \lambda_1 x}{\sin \lambda_1 A} + \dots \\
&+ \sum_1^\infty \frac{-\gamma_n t_n J_1(t_n r)(1+i)}{\sqrt{1+f}\sqrt{2}} \frac{\cos \lambda_2 x}{\sin \lambda_2 A} + \dots \\
p &= \epsilon_1 \sqrt{Re} J_1(\mu_{11}^0 r) \frac{\sin k_{11}x}{\sin k_{11}A} + 0(1),
\end{aligned}$$

(2.22a, b, c, d)

We note that in the above expansion f formally differs from $\bar{f}$ by $O(Re^{-\frac{1}{2}})$. If we now apply the endwall conditions on the radial and circumferential velocity components we find, after using the well-known orthogonality properties of Bessel functions, that

$$\begin{aligned}\gamma_n J_1^2(t_n) &= \frac{-\mu_{11}^0 \epsilon_1 i}{(1+f)} \int_0^1 r J_0(\mu_{11}^0 r) J_0(t_n r) dr, \\ \beta_n [J_2'(s_n)]^2 &= \frac{-i \epsilon_1 \mu_{11}^0}{(3-f)} \int_0^1 r J_2(\mu_{11}^0 r) J_2(s_n r) dr.\end{aligned}\quad (2.23a, b).$$

The endwall condition (1.12c) on the axial velocity component then gives

$$\begin{aligned}\frac{ik_{11} \epsilon_1 \sqrt{Re}}{(f-1)} J_1(\mu_{11}^0 r) \cot k_{11} A &+ \sum_2^\infty \frac{\epsilon_n i k_{1n}^0}{(f-1)} J_1(\mu_{1n}^0 r) \cot k_{1n}^0 A \\ &+ \sum_1^\infty \frac{s_n \beta_n J_2'(s_n r) (1-i) \cot \lambda_1 A}{\sqrt{2} \sqrt{3-f}} \\ &+ \sum_1^\infty \frac{-\gamma_n t_n J_1(t_n r) (1+i) \cot \lambda_2 A}{\sqrt{2} \sqrt{1+f}} = -2if r.\end{aligned}$$

We now multiply the above equation by $r J_1(\mu_{11}^0 r)$ and integrate from $r = 0$ to $r = 1$ and replace β_n and γ_n using (2.23). After some manipulations we can show that

$$\begin{aligned}\epsilon_1 \left[\frac{\sqrt{Re}}{2} \cot k_{11} A \left(1 + \frac{1}{(k_{11})^2} \right) J_0^2(\mu_{11}^0) \left(\frac{f-1}{f+1} \right)^2 \frac{(\mu_{11}^0)^2 i k_{11}}{f-1} \right. \\ \left. - \frac{(1+i) \cot \lambda_1 A}{\sqrt{2} (3-f)^{3/2}} (\mu_{11}^0) J_0(\mu_{11}^0) S_1 \right. \\ \left. - \frac{(1-i) \cot \lambda_2 A}{\sqrt{2} (1+f)^{3/2}} (\mu_{11}^0)^2 J_0^2(\mu_{11}^0) S_2 \right] \\ = \frac{-2if}{\mu_{11}^0} J_2(\mu_{11}^0)\end{aligned}\quad (2.24)$$

where

$$\begin{aligned}S_1 &= \sum_1^\infty \frac{s_n z_n}{[(\mu_{11}^0)^2 - s_n^2] J_1(s_n)} \\ S_2 &= \sum_1^\infty \frac{t_n^2}{[(\mu_{11}^0)^2 - t_n^2]^2}\end{aligned}$$

Here $z_n = \int_0^1 r J_2'(s_n r) J_1(\mu_{11}^0 r) dr$. We note that in (2.24) $\sqrt{Re} \cot k_{11} A \sim O(Re^0)$ because f differs from $\bar{f}$ by $O(Re^{-\frac{1}{2}})$. Finally we note that $\cot \lambda_1 A$ and $\cot \lambda_2 A$ to this order can be

replaced by $-i$ and $+i$ respectively. Thus with ϵ_1 given by (2.24) the pressure perturbation is

$$p = \epsilon_1 \sqrt{Re} J_1(\mu_{11}^0 r) \quad (2.25)$$

which is of course $O(Re^{\frac{1}{2}})$ larger than that in the non-resonant case. It can be shown from (2.24) that when $1 \gg f - \bar{f} \gg Re^{-\frac{1}{2}}$ the equation (2.25) is consistent with that obtained from (2.19) by taking the limit $f \rightarrow \bar{f}$. Thus the formula

$$p = \epsilon_1 \sqrt{Re} J_1(\mu_{11}^0 r) + \sum_2^{\infty} \frac{4f(1-f)^2}{(1+f)} \frac{\tan k_{1n} A J_1(\mu_{1n}^0 r)}{(k_{1n}^0)[1 + (k_{1n}^0)^2] J_1(\mu_{1n}^0)} \quad (2.26)$$

gives a correct zeroth order approximation to p for all f . The pressure coefficient c_p is thus given by

$$c_p = \left| \frac{4f(1-f)^2}{1+f} \sum_2^{\infty} \frac{\tan k_{1n}^0 A J_1(\mu_{1n}^0 r)}{(k_{1n}^0)[1 + (k_{1n}^0)^2] J_1(\mu_{1n}^0)} + f^2 r A + \epsilon_1 \sqrt{Re} J_1(\mu_{11}^0 r) \right| \quad (2.27)$$

with ϵ_1 given by (2.24).

We close this section with a discussion of the relationship between our resonant expansion procedure discussed above with the viscous correction to Stewartson's inviscid theory given by Wedemeyer^[6]. In the latter paper Wedemeyer showed how a viscous correction to the inviscid theory could be obtained by considering viscous boundary layers at the ends and sidewall of the cylinder. (We observe that α and β of Eqs.(33a) and (33b) of Reference 6 are related to λ_1 and λ_2 on p.9 of this paper as follows: $\lambda_1 = i\alpha\alpha$, $\lambda_2 = -i\alpha\beta$.) We note that away from the endwalls the terms that appear inside the summation signs in (2.22) are exponentially small. The term that remains in the expressions for the velocity and pressure corresponds to Stewartson's eigenvalue. However the axial wavenumber used there is k_{11} correct to order $Re^{-1/2}$. Thus there is a viscous correction in our expansion that arises from the fact that we choose to continue the inviscid expansion in order to take care of the viscous sidewall layer. Thus the modified eigenfunction we have used is precisely that implied by Wedemeyer's equation (16). At the ends of the cylinder all the ratios of the trigonometric functions are $O(1)$ so that all the terms are to be retained. Now the terms proportional to $\sin(\lambda_j x)$, $\cos(\lambda_j x)$ are needed in order to satisfy the endwall conditions. If we choose to look at the unforced problem, then it can be shown that the effect of these terms is exactly equivalent, in a formal asymptotic sense, to Wedemeyer's conditions (18,19). Thus, for the unforced problem at the resonant frequency, our expansion is equivalent to that of Wedemeyer. Notice, however, that the form of our expansion

enables us to determine in a precise asymptotic form the forced velocity field both at the resonant frequency and elsewhere. That is not the case for Wedemeyer's theory so we can interpret (2.22) as being a generalization of Wedemeyer's work to the forced problem at all frequencies.

3. Results and Discussion

We shall give results only for the composite expansion (2.27) rather than plotting separately the resonant and non-resonant expansions. In particular we will compare our results with those of Whiting^[7] who investigated the problem experimentally. In Figures 3a,b,c we compare with experimental data for the pressure coefficient measured at $r = .668$ with $A = 3.1481$ at three different values of the Reynolds number. Also shown in these figures are curves representing theoretical results from Gerber et al.^[4] and Murphy^[8]. Both of the latter calculations used approximate endwall conditions obtained using Wedemeyer's approach. At $Re = 5 \times 10^3$ the results of the present approach and Reference 7 agree more closely with experiment than the results of Reference 4 theory. At $Re = 10^5$ the present theory gives as good agreement with the theory as the other approaches but at the highest value of the Reynolds number our method overestimates the pressure coefficient at the resonant frequency. Apart from the immediate neighborhood of the resonant frequency the different theoretical approaches at the two highest values of Re are in agreement. The difference between the predicted resonant response and the experimental data could presumably be eliminated by proceeding with the present calculation to the next order.

In Figure 4 we show a similar comparison at $A = 1.0509$; again we see that the differences in the two approaches are small except for a small interval near the resonant frequency. In Figure 5 we show the variation of the resonant frequency with Reynolds number for $A = 3.1481$.

In Figure 6 we show 3D plots of the pressure coefficient at two different Reynolds numbers over a portion of the aspect ratio-frequency plane. Equal intervals in f and A are used in these plots. The localized peaks correspond to resonant responses caused by the excitation of one of the Stewartson modes. Such plots are useful in picking out portions of parameter space with certain properties; more detailed plots or variations can be obtained from the program written to calculate c_p . A few points about the approach that we have developed here are in order:

1. The composite expansion is relevant whenever any one of the Stewartson eigenvalues occurs; the eigenvalue whose eigenfunction is rescaled by a factor $\sqrt{Re}$ is taken to be that eigenvalue.

2. The computational time for the present approach is comparable to that required for the different approaches mentioned above. For example all the calculations used to supply the data for the 3D plots (with 2500 nodes) were carried out in about one hour on an Olivetti M24 PC.

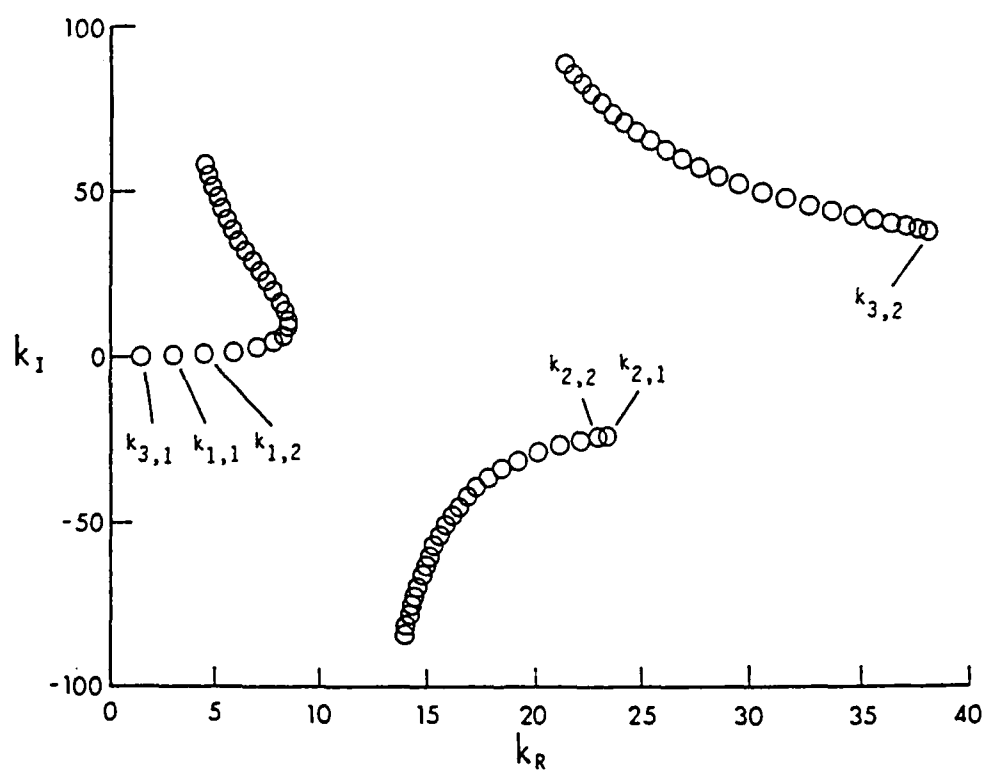


Figure 1. The first 72 eigenvalues for $Re = 1000, f = 0.1$
(taken from Reference 2).

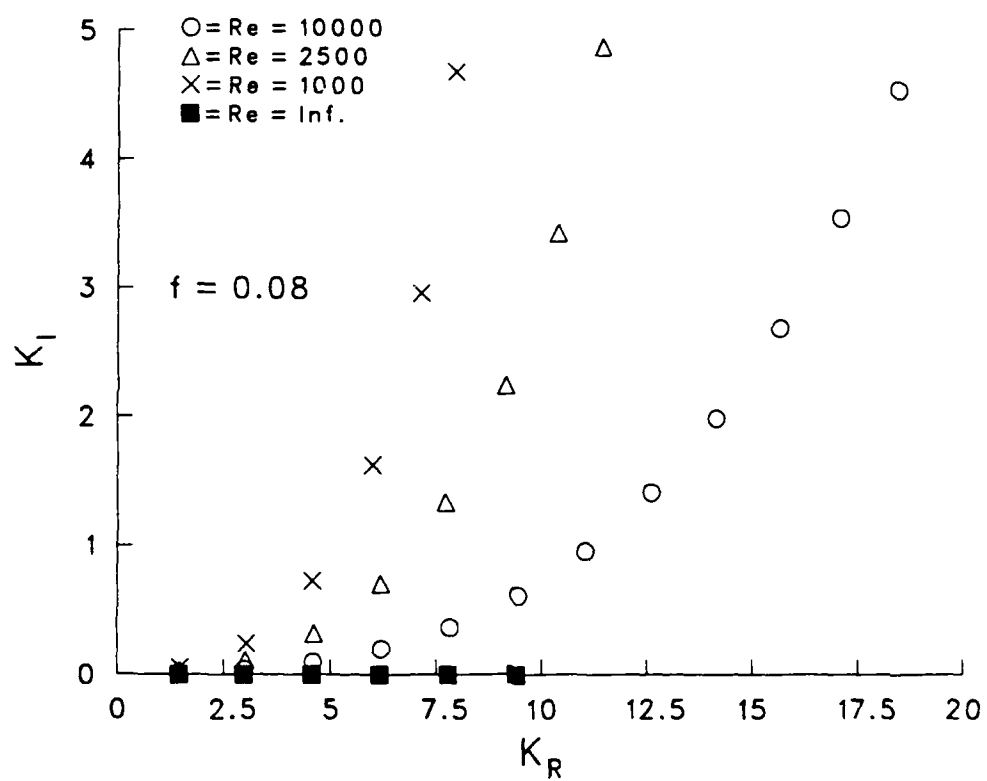


Figure 2. The eigenvalue spectrum for some large values of Re and $f = 0.08$

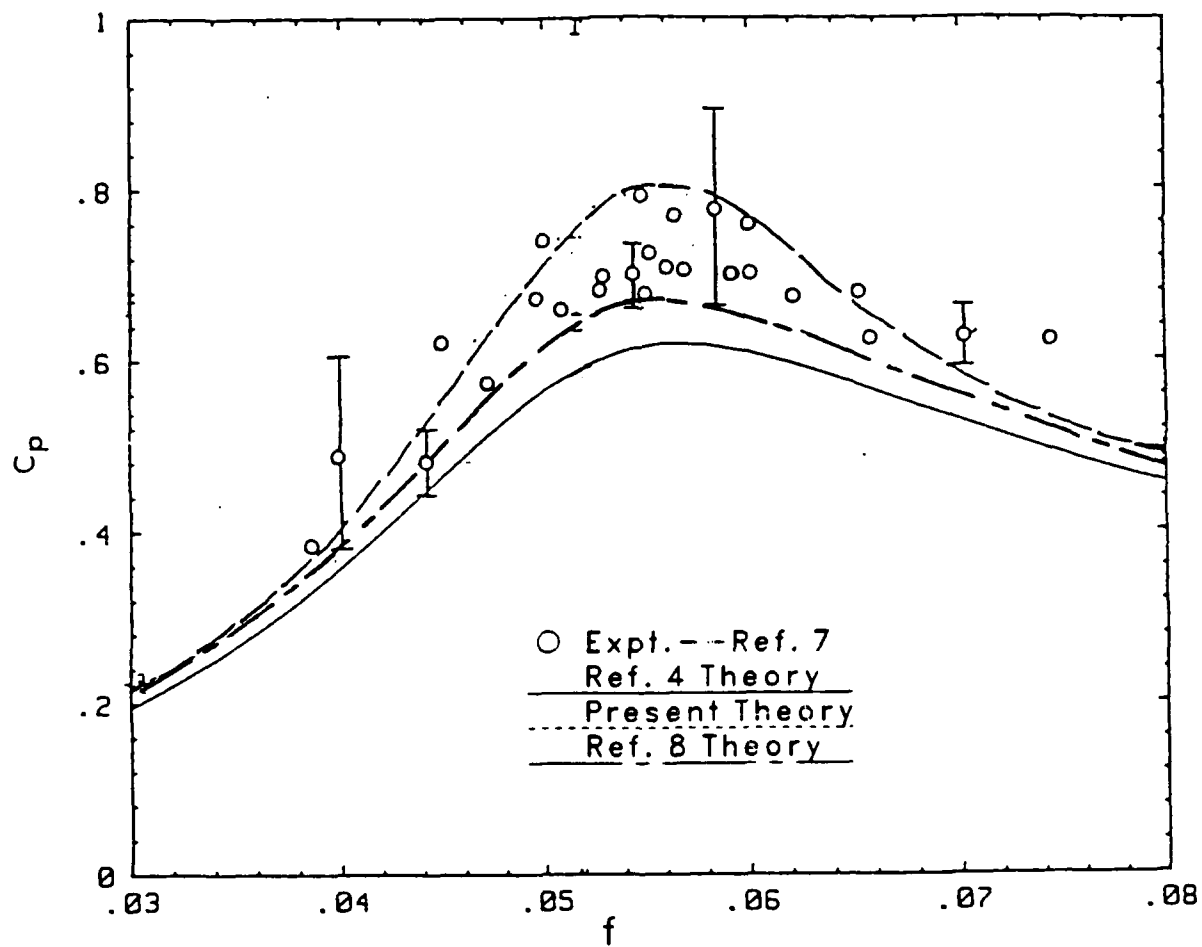


Figure 3a. The behavior of c_p calculated using the composite expansion with $A = 3.1481$, $Re = 5000$. The experimental points are from Reference 7.

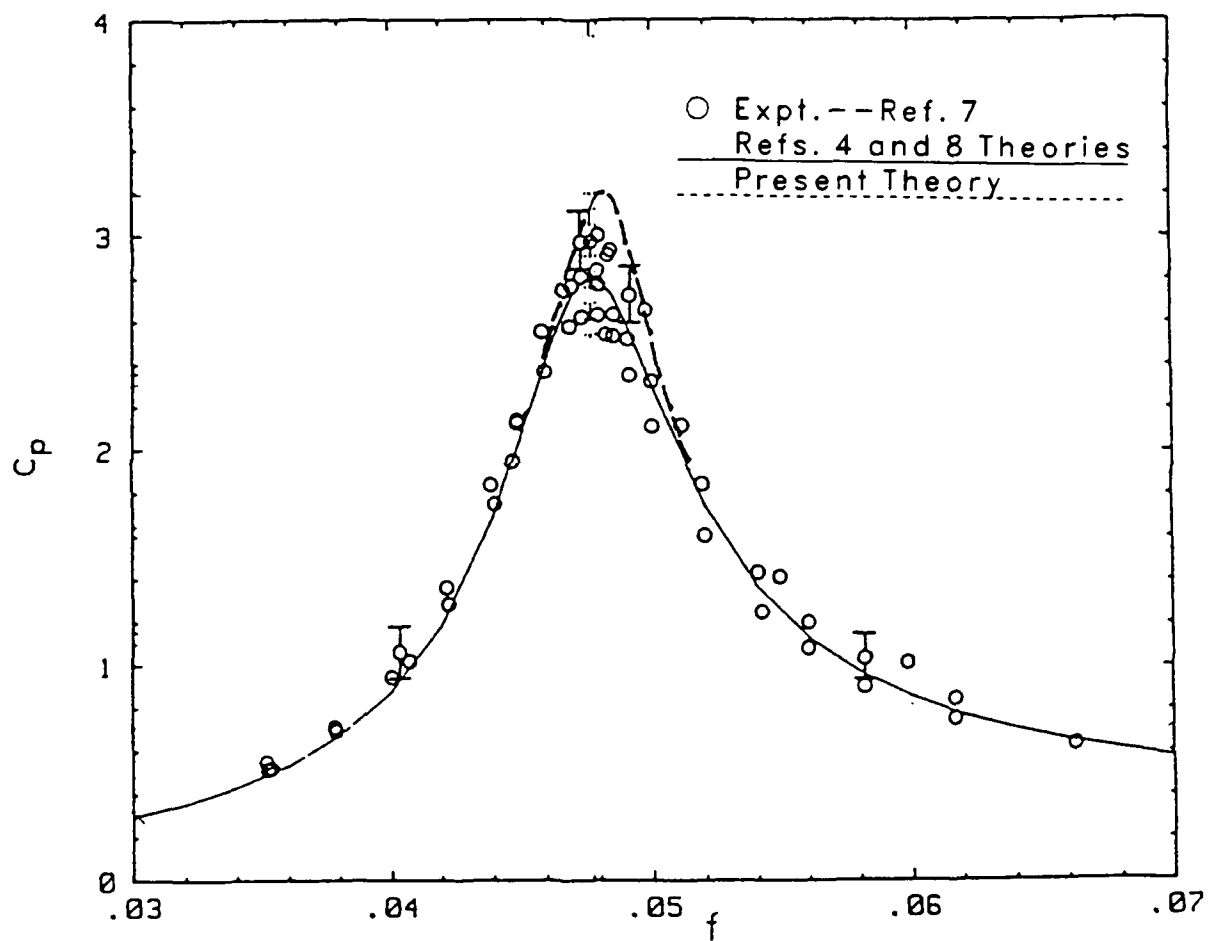


Figure 3b. The behavior of c_p calculated using the composite expansion with $A = 3.1481$, $Re = 100,000$.

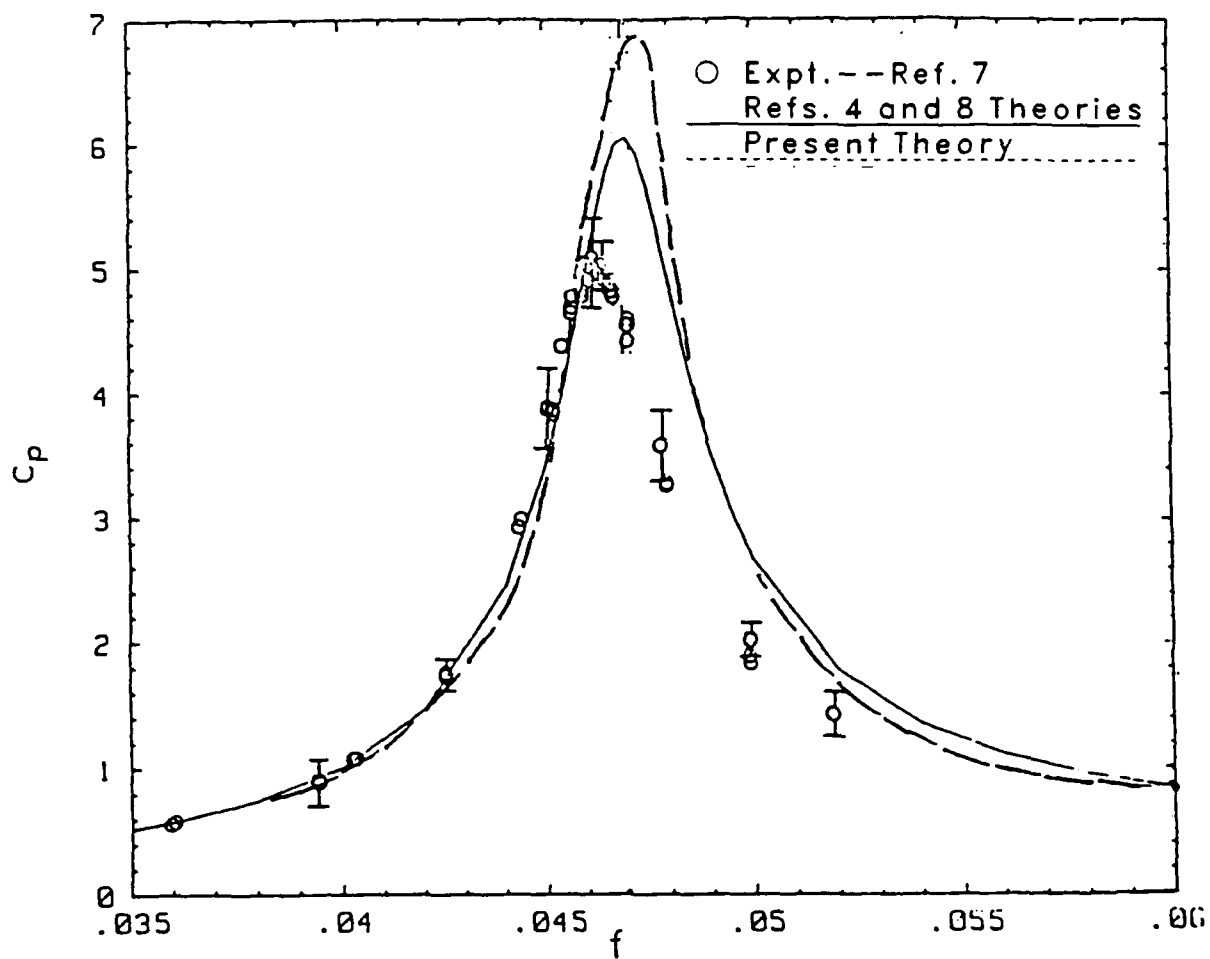


Figure 3c. The behavior of c_p calculated using the composite expansion with $A = 3.1481$, $Re = 500,000$.

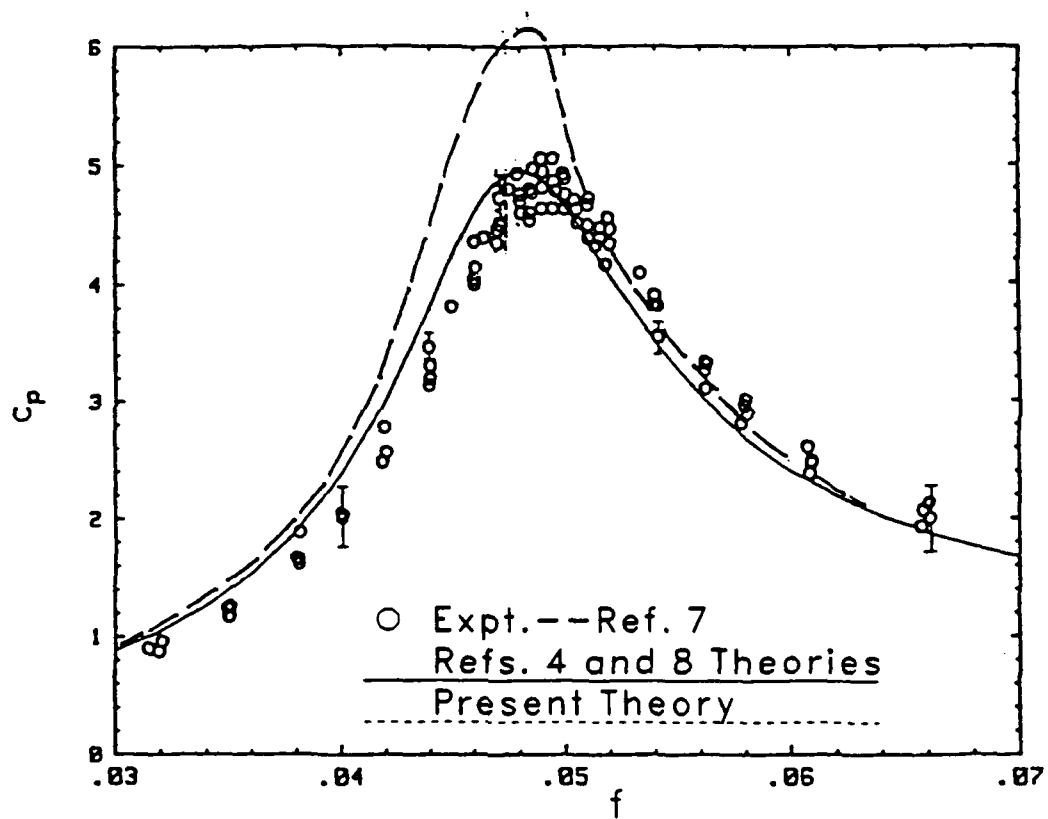


Figure 4a. The behavior of c_p calculated using the composite expansion with $A = 1.0509$, $Re = 100,000$.

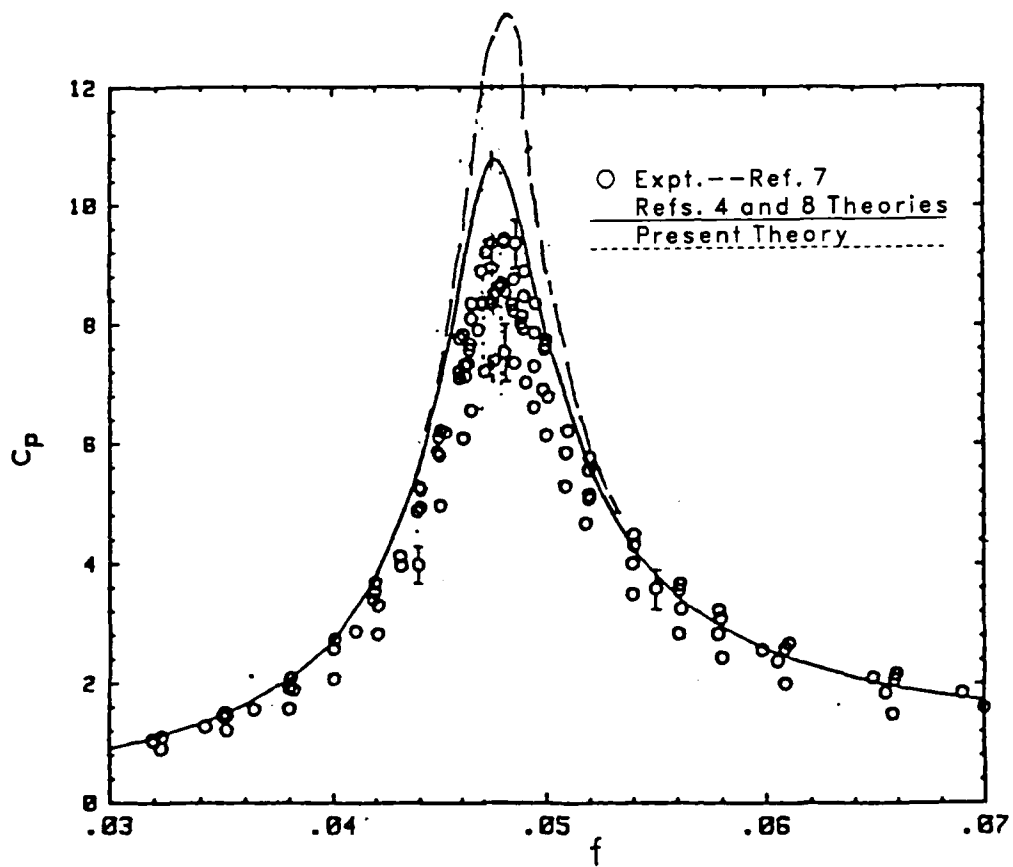


Figure 4b. The behavior of c_p calculated using the composite expansion with $A = 1.0509$, $Re = 500,000$.

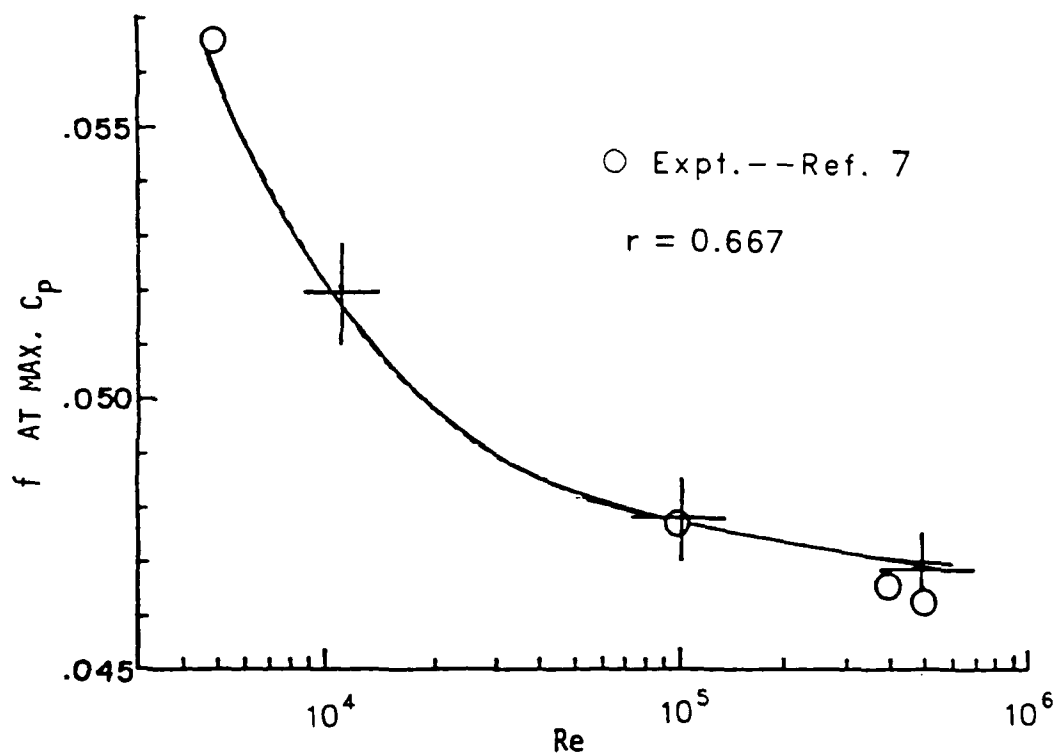


Figure 5. Variation of the frequency of peak response with Re for $A = 3.148$. Solid line is from the approach of Reference 4. The crosses are from the present calculations.

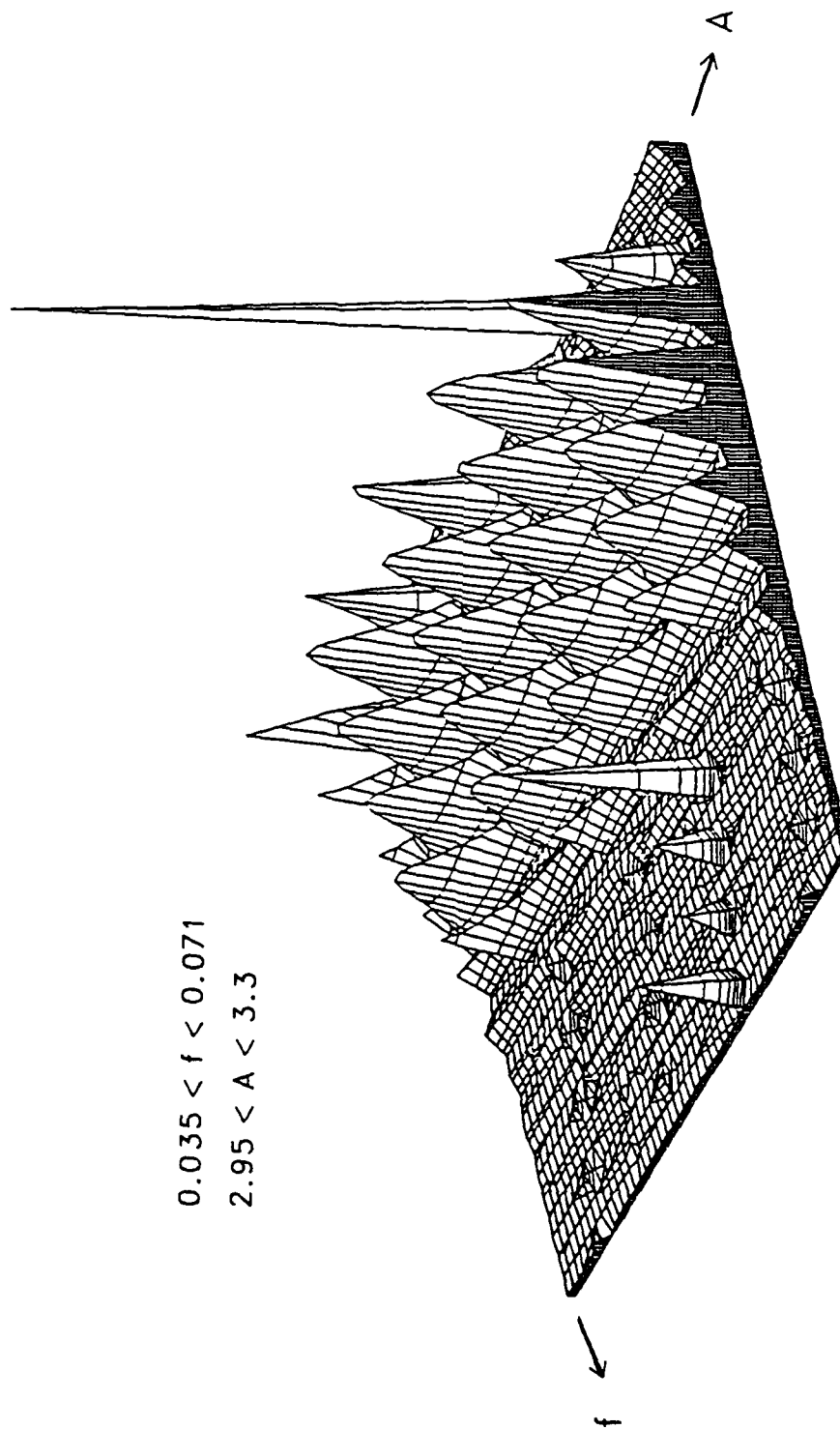


Figure 6a. A 3D plot of c_p against A and f for the case $Re = 100,000$.

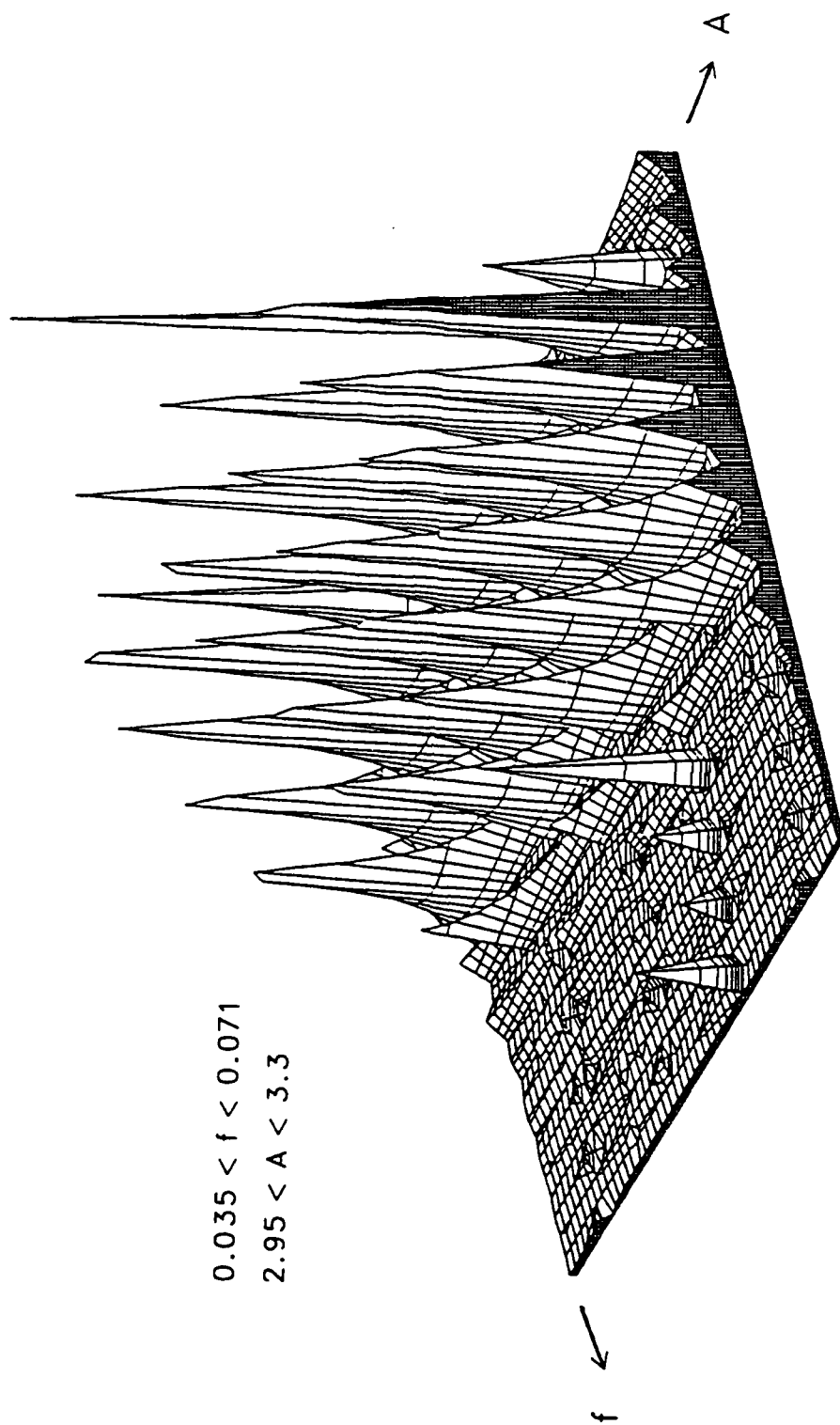


Figure 6b. A 3D plot of c_p against A and f for the case $Re = 500,000$.

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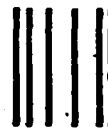
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