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ROTATIONAL TEMPERATURE ESTIMATION OF CO  
AT HIGH TEMPERATURES BY GRAPHICAL METHODS  
USING FTIR SPECTROMETRY

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OCTOBER 1990

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## I. INTRODUCTION

The theory and technique for determining the temperatures of gases from their Fourier transform infrared spectra, in emission and absorption, has been well documented in the literature.<sup>1</sup> Frequently, temperature determination involves fitting peak absorbances of rotational fine structure within a vibrational transition to a Boltzmann-type population expression,<sup>2</sup> since the relative absorbances are strongly temperature dependent. One of the difficulties associated with calculating the temperature of a gas from the peak absorbances of ro-vibrational transitions is instrumental distortion of peak absorbance values. This distortion results from the convolution of the source intensity (modulated by the sample) and the instrument lineshape function (ILS), and is most important in spectra obtained at moderate to low resolution (here taken to correspond to optical retardations of less than  $\sim 1.0$  cm). The method for correcting the error in observed peak absorbance due to this convolution as applied to FTIR spectroscopy was first described by Griffiths.<sup>3</sup>

The analytical treatment of most physically encountered convolutions is difficult.<sup>4</sup> To our knowledge, no analytical solution to the resulting "convolution integral" for transmission spectra obtained on a FTIR spectrometer has been reported. We have recently investigated and quantified the relationship between observed and undistorted peak absorbances measured by FTIR spectroscopy. This relationship was first quantified by Griffiths and Anderson.<sup>3</sup> Our goal is to develop a reasonably accurate (within  $\sim 5\%$ ) graphical method to obtain the temperature of a gas sample at elevated temperatures ( $> 500$  K) from absorption spectra in the infrared using resources typically available in an analytical laboratory. We assume this to be a Fourier transform IR spectrometer capable of  $1 \text{ cm}^{-1}$  resolution and a microcomputer of virtually any type. The graphical method is not intended to supplant the more accurate methods already in the literature,<sup>1,5</sup> but rather to provide a fast and convenient complementary view of the data. To date, we are unaware of any successful temperature estimations from peak absorbances of ro-vibrational transitions which have employed graphical methods over a temperature range of several hundred degrees.

## II. BACKGROUND

For a Fourier transform spectrometer employing triangular apodization of the interferogram, the observed minimum transmittance,  $T_m$ , of an absorption feature with a Lorentzian profile centered at frequency  $\nu_m$ , may be given by:<sup>5</sup>

$$T_m = \int_{-\infty}^{\infty} \exp\{-2.303 A_{pm} a_m^2 / ((\nu - \nu_m)^2 + a_m^2)\} D \sin^2[\pi D(\nu - \nu_m)] / [\pi D(\nu - \nu_m)]^2 d\nu \quad (1)$$

where  $A_{pm}$  is the true peak absorbance of line  $p$  at frequency  $\nu_m$ ,  $a_m$  is the linewidth,  $D$  is the maximum optical retardation of the spectrometer, and  $\pi$  is 3.1416. Here,  $m$  is the line index number, where  $m = -J''$  for P branch transitions and  $m = J'' + 1$  for R branch transitions.<sup>6</sup> This equation describes the minimum transmittance of an isolated pressure broadened line where the absorption feature is unobscured by any neighboring lines. Equation (1) is a function of temperature through the dependence of the linewidth,  $a_m$ , on

temperature. The variation of  $a_m$  with temperature and pressure may be calculated by the approximation:

$$a_m(T,P) = a_{m0}(P/P_0) (T_0/T)^n \quad (2)$$

where  $P_0$  and  $T_0$  are a pressure and temperature for which the linewidth  $a_{m0}$  is known,  $P$  and  $T$  are the experimental pressure and temperature, and  $n$  is the empirically determined temperature broadening exponent. The coefficient  $n$  has been shown<sup>8</sup> to be temperature and line dependent. For the sake of simplification,  $n$  is usually set to 0.75, although there is evidence<sup>8</sup> that a value of 0.66 may be more accurate at room temperature.

True peak absorbance may be fit to temperature,  $T$ , using:<sup>9</sup>

$$A_{pm} \propto |m| F(m) v_m \exp\{-E(m)/kT\} / a_m \quad (3)$$

where  $a_m$  varies as in Eq. (2),  $E(m)$  is the rotational energy of the lower state in the transition,  $F(m)$  is the Hermann-Wallis factor for interaction of rotation and vibration,<sup>6</sup> and  $|m|$  is the absolute value of the line index number. The linewidth appears in the denominator of Eq. (3) to take into account the dependence of the peak height on peak width for a spectral absorption feature with a Lorentzian profile.<sup>10</sup>

By substituting Eq. (2) into Eq. (1) and obtaining an analytical solution to the integral, an expression relating observed minimum transmittance (and hence observed peak absorbance) to true peak absorbance may be obtained. If we denote this expression as:

$$A_{pm} = f(T_m, a_m, D) \quad (4)$$

and combine this with Eq. (3) we obtain:

$$f(T_m, a_m, D) \propto |m| F(m) v_m \exp\{-E(m)/kT\} / a_m \quad (5)$$

The first step in linearizing Eq. (5) involves finding an approximation to Eq. (1). Griffiths gives the expression:<sup>3,5</sup>

$$A_{pm} \propto A_{obs} x_R^y \quad (6)$$

where  $R$ , the resolution parameter, is equal to  $1/(2Da_m)$  and  $x$  and  $y$  are set equal to 2. If this expression is substituted into Eq. (5), and the temperature broadening exponent  $n$  is assumed to be invariant with  $m$ , one obtains:

$$A_{obs}^x / \{(2Da_{m0})^{y-1} (|m| F(m) v_m)\} \propto \exp\{-E(m)/kT\} \quad (7)$$

When ambient temperature spectra are collected (for which the  $a_m$  for many diatomics are accurately known)<sup>11</sup> and the resolution parameter for the ro-vibrational lines being used in the calculation is such that there is a near linear relationship<sup>3</sup> between  $\log(A_{pm})$  and  $\log(A_{obs})$ , a graph of  $E(m)/k$  vs.  $\ln[A_{obs}^x / \{(2Da_{m0})^{y-1} (|m| F(m) v_m)\}]$  produces a straight line with a slope  $(1/T)$  the inverse of which usually falls within a percent of the true gas temperature.<sup>9</sup> As the temperature of the gas being investigated is increased, the error between the calculated and actual temperature increases.<sup>5</sup> This

increase in error is caused by the change in  $a_m$  with  $T$ , which affects the resolution parameter  $R$ , and may shift the relationship between  $\log(A_{pm})$  and  $\log(A_{obs})$  into a nonlinear region. Also, as stated earlier, the temperature broadening exponent,  $n$ , from Eq. (2) changes with temperature and, for a given temperature, may vary with line index number.

## III. CALCULATIONS

We have fit input parameters and the result of the numerical integration of Eq. (1) to an equation of the form:

$$A_{pm} = k A_{obs}^{xR^y} \quad (8)$$

where  $k$ ,  $x$ , and  $y$  are constants and  $R$  is the spectral resolution parameter. The integral in Eq. (1) was evaluated numerically for all combinations of optical retardations of 1, 2, 4, 8, and 10 cm, true peak absorbances of 0.1 to 10 (in intervals of 0.1) with absorbing lines for which  $a_m$  ranged from 0.01 to 0.1  $\text{cm}^{-1}$  (in intervals of 0.01  $\text{cm}^{-1}$ ). A weighted line - least squares routine was used to fit the data from portions of the numerical integration of Eq. (1) to a linearized form of Eq. (8).

The integration routine assumed that  $\nu_m$  was located at 2000  $\text{cm}^{-1}$ , and elevated Eq. (1) using Simpson's  $1/3$  rule for 50  $\text{cm}^{-1}$  on either side of  $\nu_m$ . The interval to be numerically integrated was divided into 10,000 segments. The choice of value for the center of absorption,  $\nu_m$ , is arbitrary, as may be verified by substituting  $\nu - \nu_m$  in Eq. (1). To determine if numerical integration gave values dependent on the value of  $\nu_m$  ranging from 500 to 3500  $\text{cm}^{-1}$  showed the value of the integral to be independent of the value for this parameter.

The application of our numerical integration routine to the function  $\sin^2 x/x^2$  (which is similar in behavior to Eq. (1) yet has an analytical solution) using parameters similar to those in the numerical integration of (1) gave results within 0.001% of the analytical solution. We believe our numerical integration of Eq. (1) gives similar accuracy.

For all ranges of  $A_{pm}$ , values used in the integration of Eq. (1), the average deviation between the numerically integrated value  $-\log(T_m)$  and the fitted approximation presented here was less than that of Eq. (6) with  $x$  and  $y$  equal to 2. In fairness, the authors of Reference 5 presented their approximation without derivation and in no way implied it to be valid over any range other than that specified in the paper. In addition, for the range over which the approximation of Reference 5 was used, that approximation and the fitted approximation developed here are in good agreement.

For spectra of gases below 800 K, the integral in Eq. (1) was evaluated for true peak absorbances,  $A_{pm}$ , ranging from 1 to 5 absorbance units, an optical retardation of 1 cm, and  $a_m$  values ranging from 0.01 to 0.1  $\text{cm}^{-1}$ . These parameters are typical for CO at pressures of several hundred torr.<sup>5</sup> The best fit values obtained for  $\ln(k)$ ,  $x$ , and  $y$  in Eq. (8) were 0.122 (.014), 1.818 (.014), and 1.748 (.015), respectively, where the numbers in parentheses are equal to one standard deviation. For the same parameters mentioned above, but for peak absorbances ranging from 0.1 to 1.0 absorbance units, the best fit values of  $\ln(k)$ ,  $x$ , and  $y$  were respectively, 0.211 (.00004), 1.287

(.00002), and 1.298 (.00002). The spectroscopic constants used to calculate  $E(m)$  and  $F(m)$  may be found in Reference 13. Values of  $v_m$  used in the calculation are those reported by the spectrometer. Values of  $a_m$  were obtained from Reference 5. For this work, it was assumed that all gases had the same broadening effect on the  $a_m$  for CO as did CO. The additional error<sup>11</sup> probably is insignificant given the assumption that  $n$  is constant.

At temperatures below 600 K, 20 lines from the R branch (because it is more intense than the P branch and for reasons given below) in the fundamental absorption band of CO were used in the calculation. To calculate the temperature, Eq. (7) was first linearized, and a weighted linear least squares<sup>14</sup> routine used to determine  $1/T$ .

To decide which lines will be used to calculate temperatures from FTIR spectra, it is necessary to establish a criteria for line acceptability. Lines which are significantly overlapped by absorbances due to hot band transitions or to absorbances by isotopically substituted homologues should be excluded from the computation. Figure 1 shows a portion of the CO spectrum in a cell held at 866 K. Visible in this figure are absorbances due to  $^{13}\text{CO}$  and those due to the 1-2 transition in  $^{12}\text{CO}$ , as well as the much more intense fundamental transition. The error caused by hot band overlap is very temperature sensitive and becomes most important at high temperature. Figures 2 and 3 show the line center separation in wavenumbers between fundamental and hot ( $v=1-2$ ) band transitions in ordinary CO between fundamental transitions in normal and  $^{13}\text{C}$  substituted CO, respectively. These figures were generated by calculating rotational energy levels from spectroscopic constants<sup>13</sup> for the appropriate species/vibrational level, calculating transition frequencies using conventional selection rules,<sup>15</sup> ordering these transitions with respect to the energy of transition, and then taking the difference in energy between each successive level. This energy difference is plotted versus the energy of the transition with the lower energy. For each species considered,  $J''$  values from 0 or 1 to 30 were used. At a glance, each of these figures provides a good idea of where regions of minimum line overlap exist for the fundamental absorption of CO. Because these figures do not include intensities of individual transitions, they cannot indicate where a weak fundamental absorption line used in the calculation is overlapped by a significantly strong  $^{13}\text{CO}$  line, for example, or vice versa. For CO, typical ranges of lines to use in the calculation are the 20 lines in the regions  $2059-2139\text{ cm}^{-1}$  (P-branch  $J''=1-20$ ) and/or  $2147-2212\text{ cm}^{-1}$  (R-branch  $J''=0-19$ ). An examination of Figure 2 shows that these two regions of the fundamental absorption in CO significantly overlap with the  $v=1-2$  transitions in CO, with the P branch being slightly less obscured. However, since the 1-2 envelope is shifted by about  $25\text{ cm}^{-1}$  to lower energy than the 0-1 envelope (not apparent from the figure), the R branch of the 0-1 transition is relatively unobscured by those 1-2 transitions which have appreciable intensity at normal temperatures.

Figure 3 shows that both branches in the fundamental absorption of  $^{12}\text{CO}$  have several lines with separation between line centers for overlap with  $^{13}\text{CO}$  of less than  $0.5\text{ cm}^{-1}$ . However, because the band centers differ between CO and  $^{13}\text{CO}$  by about  $45\text{ cm}^{-1}$ , the R branch of CO is relatively unobscured by overlap from the R branch of  $^{13}\text{CO}$  (since the R branch of  $^{13}\text{CO}$  overlaps the P branch of CO). Therefore, Figures 2 and 3 suggest using the R-branch of CO for temperature determinations at temperatures where hot band transitions are not important. At higher temperatures, Figure 3 suggests a switch to the

# CO at 593 C using .125 cm-1 resolution

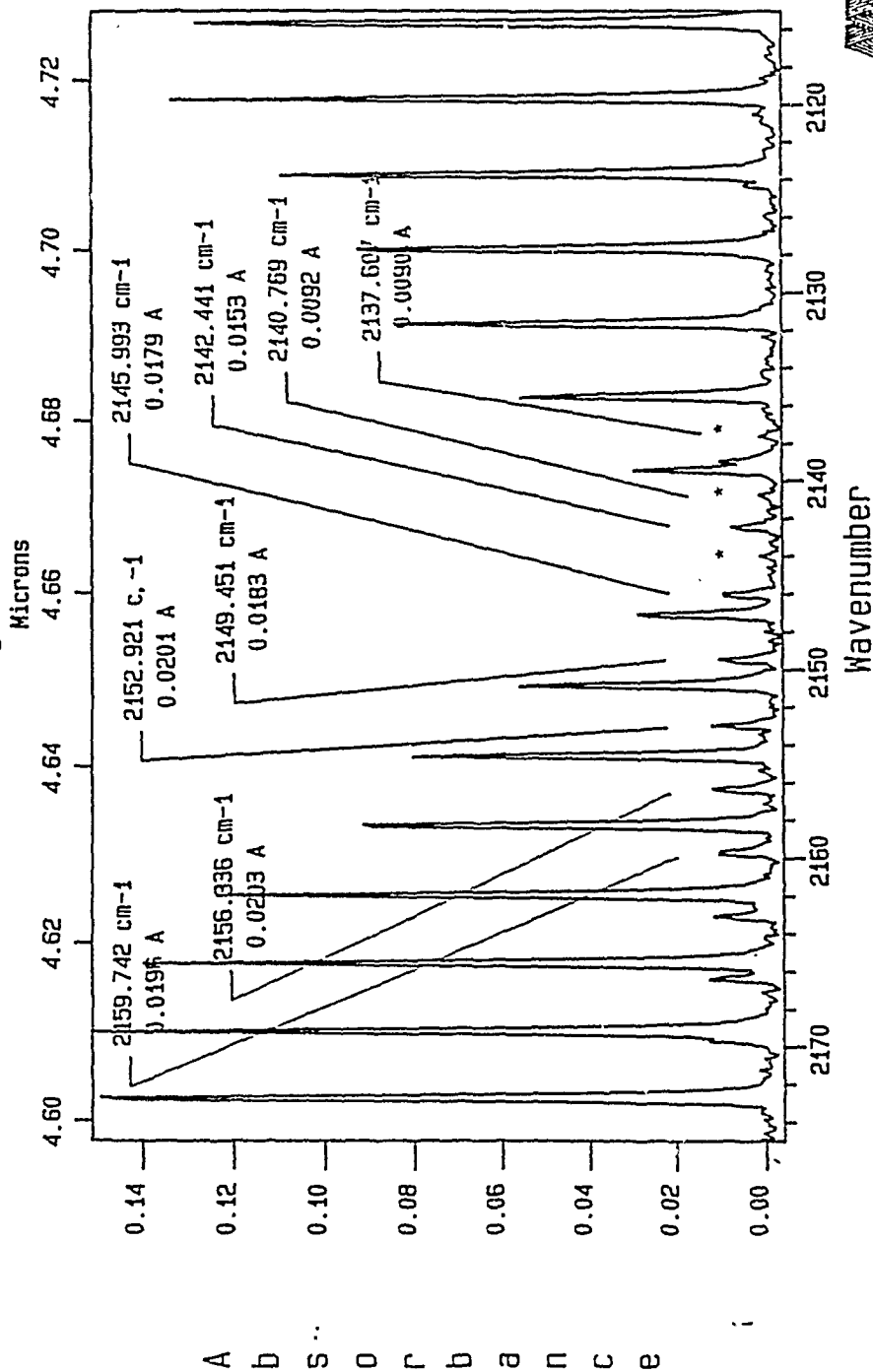


Figure 1. A portion of the IR Spectrum of CO at 866 K. Absorbances marked \* are due to <sup>13</sup>CO; otherwise, annotated peaks are due to the v=1-2 transition in CO. The most intense peaks are due to the fundamental absorption in CO.

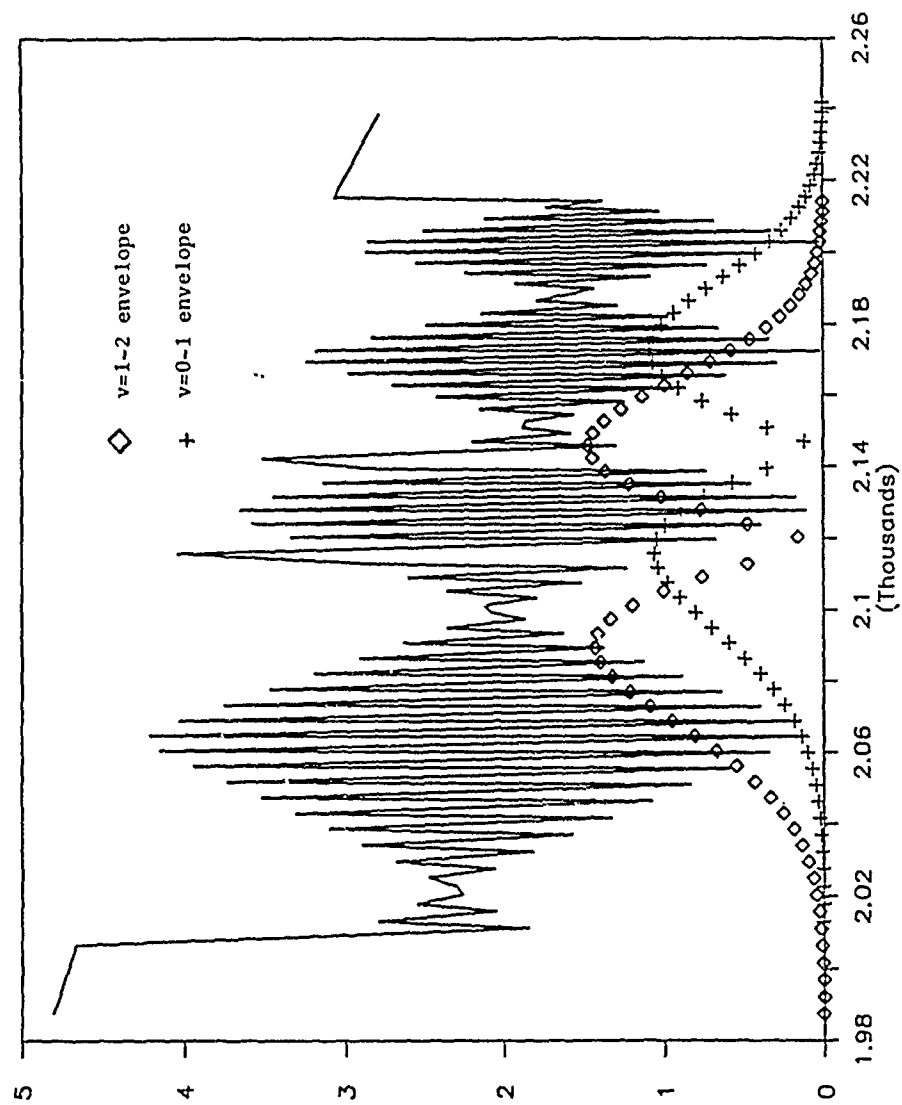


Figure 2. Line Center Separation Between the  $v=0-1$  and  $v=1-2$  Transitions in CO (see text)

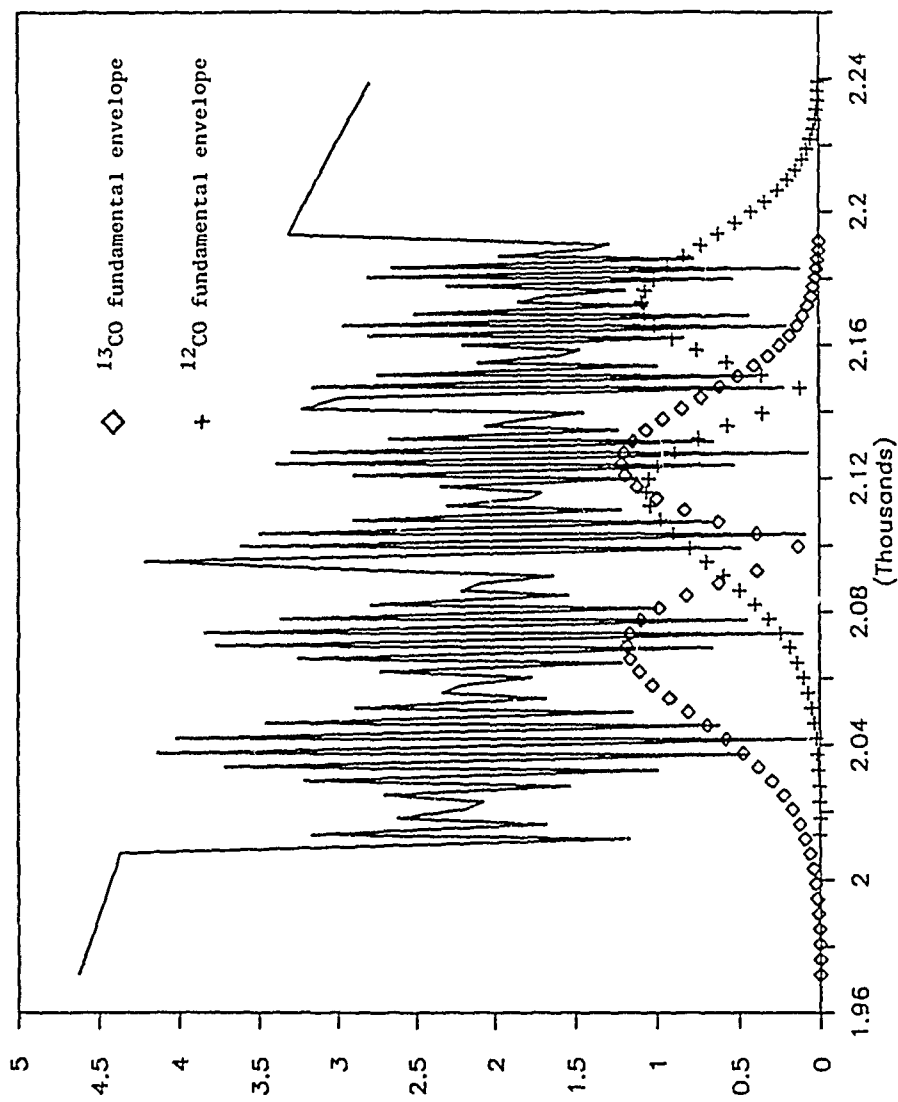


Figure 3. Line Center Separation Between the  $v=0-1$  Transition in  $^{12}\text{CO}$  and the  $v=0-1$  Transition in  $^{13}\text{CO}$  (see text)

P-branch of CO for temperature determinations when 1-2 transitions become more troublesome than  $^{13}\text{CO}$  transitions.

#### IV. EXPERIMENTAL

All temperatures reported here were obtained from absorption spectra. Since the heated cell/burner was located in the sample compartment of an otherwise unmodified FTIR, infrared emission was unmodulated, and no attempt was made to collect emission spectra. No correction was made to the spectra to account for emission. Experiments were performed on a Mattson Sirius 100 FTIR spectrometer using a HgCdTe liquid nitrogen cooled detector. All spectra were obtained with 100 co-added scans obtained at  $1\text{ cm}^{-1}$  resolution. Interferograms were triangularly apodized and zero-filled to the limit of our computer memory (131072 points) prior to transformation.

Temperature controlled stainless steel cells were equipped with NaCl windows. The pathlength for absorption in each cell was approximately 12 cm. For the temperature range 300 K to 473 K, the NaCl windows were attached to the cell body using metal flanges and fiber gaskets. The leak rate of this cell at all temperatures used was always less than 500 mtorr/hr. Pressures used in these low temperature experiments ( $<473\text{ K}$ ) ranged from 115 to 197 torr. For the temperature range from 473 K to 773 K the NaCl windows were attached directly to the cell body by pressure supplied by flanges. To prevent heat loss, the high temperature cell was encased in asbestos which extended approximately 1 inch past the windows on either side of the cell. No other precautions were taken to avoid temperature gradients along the optical pathlength. The leak rate of this cell was such that all high temperature experiments were performed at 760 torr total pressure using a 4:1 mixture of air:CO. A check of the instrument after filling this cell, allowing it to stand for a typical time of one experiment ( $\sim 10$  minutes) and then evacuating the cell gave no CO absorption due to CO in the optical compartment of the spectrometer.

Temperature was controlled by an Omega Model 2010 temperature controller using heating cartridges placed in cavities in the cell body and monitored by copper-constantin thermocouples placed near the inner wall of the cell and placed within the gas volume but out of the optical path. A check of the cell temperature by an Omega Model 2168A digital thermometer inserted into the body of the cell gave temperature readings within 5 K of those reported by the controller thermocouple. All temperatures reported as actual are the temperatures of the cell body. After filling the cell, the gas was allowed to stand for three to five minutes to reach the temperature of the cell. Waiting for longer periods of time gave no significant change in calculated temperature. No correction was made for temperature variations in the vicinity of the cell windows. At temperatures greater than 550 K, some conversion of CO to  $\text{CO}_2$  was observed for the cell containing CO plus air.

The experimental setup for the low pressure flame experiments is shown in Figure 4. The optical pathlength within the burner chamber was approximately 6 cm. The burner head was constructed entirely of Pyrex, with the burner housing constructed of brass and aluminum. Gas flow to the mixing section below the burner head was regulated by calibrated floating-ball type flowmeters. Pressure was measured using an MKS Instruments Baratron capacitance manometer and regulated by adjusting a valve downstream from the



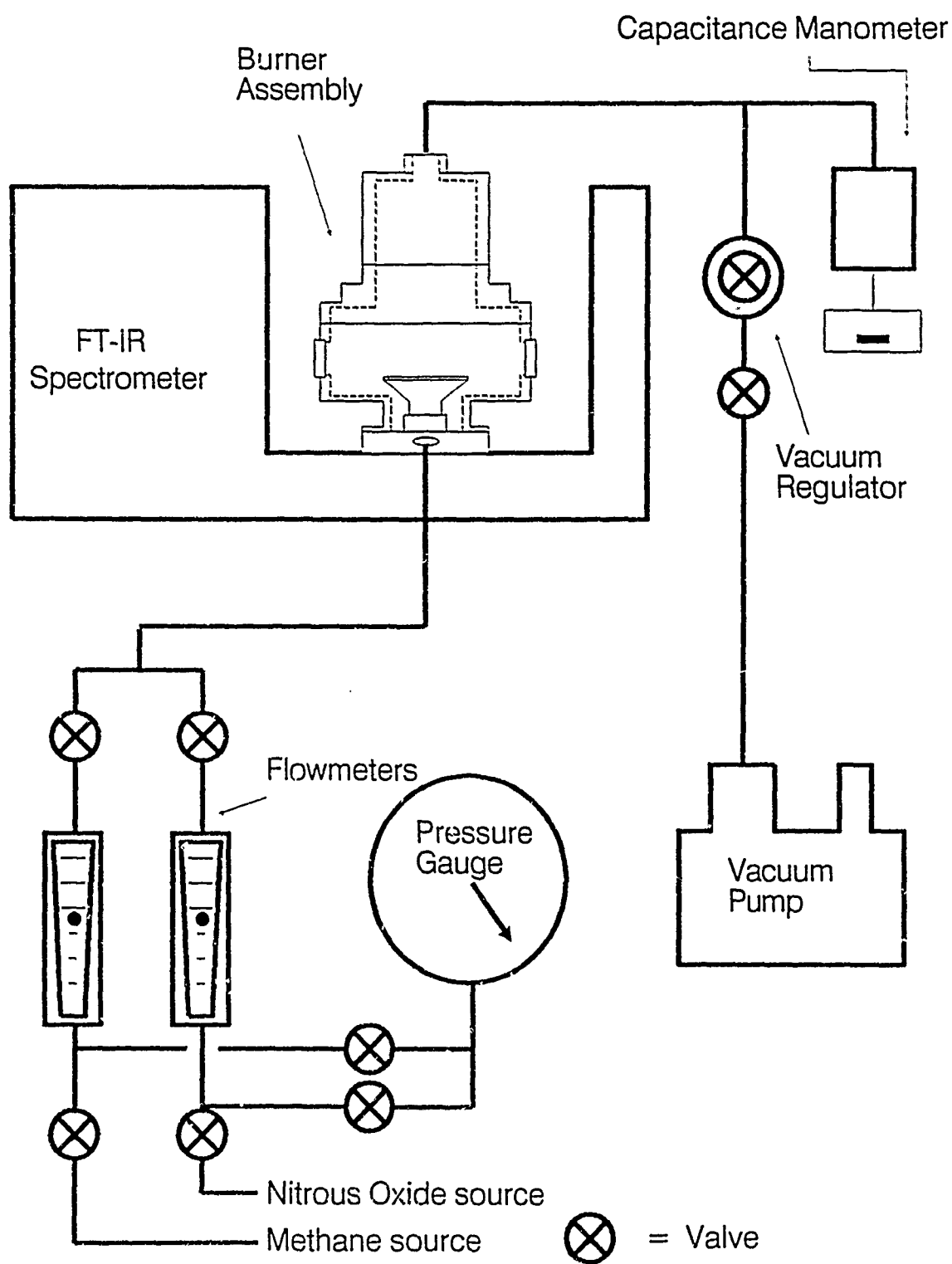


Figure 4. Experimental Setup for Burner Flame Experiments

flowmeters. Windows were of LiF. A low pressure (<100 torr operation), the flame is lifted off the burner surface by 2-4 mm, preventing damage to the Pyrex burner frit. Data was collected at the maximum iris setting of the spectrometer. Although maximizing sample compartment aperture increases the likelihood of resolution errors due to beam divergence, the beam waist at all times was kept within the limits built into the spectrometer. No correction was applied to reflect any non-collimated beam related errors.

## V. RESULTS AND DISCUSSION

The results of temperature calculations using the temperature controlled gas cells for the range 300 K to 773 K are shown in Table 1. Values of  $x$  and  $y$  used in calculations over this temperature range were 1.818 and 1.748, respectively. At temperatures above 600 K, the P branch of the ro-vibrational envelope of CO was used in the calculations (see discussion above). Figure 5 shows a plot of  $E(n)/k$  vs.  $\ln(A_{\text{obs}}^x / \{(2Da_{m,0})^{y-1} |n| F(n) v_m\})$  for samples of CO held in the temperature controlled cells over the temperature range 302 K to 773 K. The slope of the best straight line through each set of data obtained by the weighted linear least squares routine<sup>14</sup> is equal to  $1/T$ . In all cases, calculations yield gas temperatures within 10% of the cell temperature. Neglecting the one calculated point at 620 K, calculated temperatures are within 5% below 700 K. The best results were obtained when the temperature of the gas was near 400 K. This is in contrast to earlier work<sup>5</sup> in which the best results were found at room temperature. The reason for this difference is that our fit to Eq. (1) is more accurate for small values of  $A_{\text{pm}}$  (<3) than for cases where  $x$  and  $y$  in Eq. (6) are equal to 2. As average peak absorbances of individual lines decrease with increasing temperature (total number of lines increases), an approximation to Eq. (1) which emphasizes small values of  $A_{\text{pm}}$  yields a temperature that is closer to actual than when  $x$  and  $y$  in Eq. (6) are equal to 2.

The curvature in the far right-hand side of the plots in Figure 5 in the region corresponding to low values of  $J''$  is analogous to that seen in Reference 5. This arises from the fact that CO fundamental absorbances with low values of  $J''$  have the largest  $a_m$  and smallest value of the resolution parameter,  $R$ . Spectra of lines with small resolution parameters and small absorbances have been shown<sup>3,5</sup> not to adhere to a linear relationship between  $\log(A_{\text{pm}})$  and  $\log(A_{\text{obs}})$  when obtained using triangular or sinc<sup>2</sup> instrument line shape (ILS) functions.

Figure 6 shows a portion of the absorbance spectrum of a rich low pressure (64 torr) premixed methane/nitrous oxide burner flame used in the calculation of rotational temperatures of CO. The CO hot band ( $v=1-2$ ) P branch transitions were chosen for the analysis because the hot band transitions are shifted to lower energy than those transitions of the fundamental. Also, interference from CO<sub>2</sub> bands around 2349 cm<sup>-1</sup> and from N<sub>2</sub>O bands around 2223 cm<sup>-1</sup> makes finding suitable lines in the fundamental of CO difficult. Hot band transitions were chosen to discriminate against any cold CO near the inner surfaces of the burner chamber windows. Peaks were identified by calculating the peak locations for the  $v=1-2$  ro-vibrational envelope of CO from spectroscopic constants.<sup>13</sup> All calculations agreed within a few hundredths of a wavenumber with the peak locations reported by the

Mattson peak picking software. Table ^ shows calculated CO rotational temperatures as a function of the distance of the center of the IR beam above the burner frit surface.

Table 1. Calculated and Actual CO Temperatures for the Temperature Controlled Gas Cells Used in These Experiments Over the Range 302 to 773 K

| <u>Actual<br/>Temperature<br/>(Kelvin)</u> | <u>Calculated<br/>Temperature<br/>(Kelvin)</u> | <u>Standard<br/>Deviation<br/>(Kelvin)</u> | <u>Total<br/>Pressure<br/>(torr)</u> | <u>Actual-Observed<br/>Temperature<br/>(Kelvin)</u> |
|--------------------------------------------|------------------------------------------------|--------------------------------------------|--------------------------------------|-----------------------------------------------------|
| 302                                        | 299                                            | 4                                          | 166.8                                | 3                                                   |
| 302                                        | 297                                            | 4                                          | 164.7                                | 5                                                   |
| 302                                        | 296                                            | 4                                          | 166.8                                | 6                                                   |
| 313                                        | 306                                            | 5                                          | 760                                  | 7                                                   |
| 322                                        | 313                                            | 3                                          | 196.7                                | 9                                                   |
| 323                                        | 319                                            | 5                                          | 163                                  | 4                                                   |
| 325                                        | 323                                            | 6                                          | 141.8                                | 2                                                   |
| 332                                        | 324                                            | 4                                          | 184.7                                | 8                                                   |
| 343                                        | 335                                            | 4                                          | 195.1                                | 8                                                   |
| 343                                        | 337                                            | 4                                          | 189.6                                | 6                                                   |
| 344                                        | 340                                            | 6                                          | 152.1                                | 4                                                   |
| 363                                        | 364                                            | 7                                          | 127.3                                | -1                                                  |
| 364                                        | 360                                            | 5                                          | 190.3                                | 4                                                   |
| 364                                        | 360                                            | 4                                          | 193.5                                | 4                                                   |
| 373                                        | 369                                            | 5                                          | 194                                  | 4                                                   |
| 373                                        | 369                                            | 5                                          | 196.2                                | 4                                                   |
| 374                                        | 373                                            | 6                                          | 164                                  | 1                                                   |
| 413                                        | 415                                            | 9                                          | 140.9                                | -2                                                  |
| 413                                        | 416                                            | 9                                          | 141.4                                | -3                                                  |
| 432                                        | 432                                            | 8                                          | 164.8                                | 0                                                   |
| 433                                        | 436                                            | 11                                         | 115.1                                | -3                                                  |
| 438                                        | 437                                            | 8                                          | 184.6                                | 1                                                   |
| 473                                        | 477                                            | 3                                          | 760                                  | -4                                                  |
| 473                                        | 478                                            | 5                                          | 760                                  | -5                                                  |
| 576                                        | 602                                            | 9                                          | 760                                  | -26                                                 |
| 592                                        | 594                                            | 16                                         | 760                                  | -2                                                  |
| 674                                        | 620                                            | 61                                         | 760                                  | 54                                                  |
| 674                                        | 672                                            | 13                                         | 760                                  | 2                                                   |
| 674                                        | 673                                            | 40                                         | 760                                  | 1                                                   |
| 773                                        | 758                                            | 23                                         | 760                                  | 15                                                  |
| 773                                        | 813                                            | 48                                         | 760                                  | -40                                                 |
| 773                                        | 817                                            | 35                                         | 760                                  | -44                                                 |

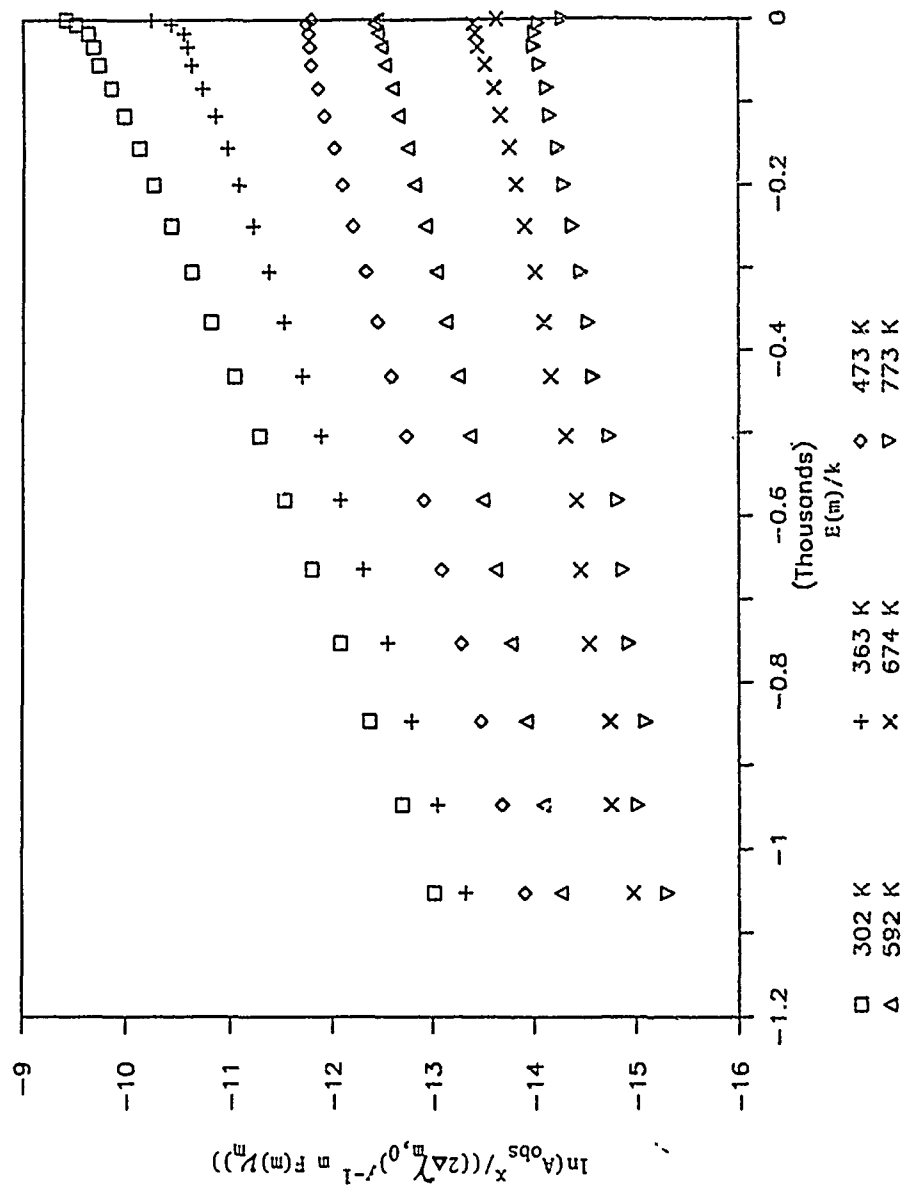


Figure 5. A Plot of  $-\ln(A_{obs}^x / ((2\Delta\chi_{m,0})^{x-1} F(m) V_m^x))$  for Samples of CO Held in the Temperature Controlled Cells Over the Temperature Range 302 K to 773 K. The slope of each line is equal to  $T^{-1}$ . The abscissa for each set of data has been shifted to eliminate overlap of data sets.

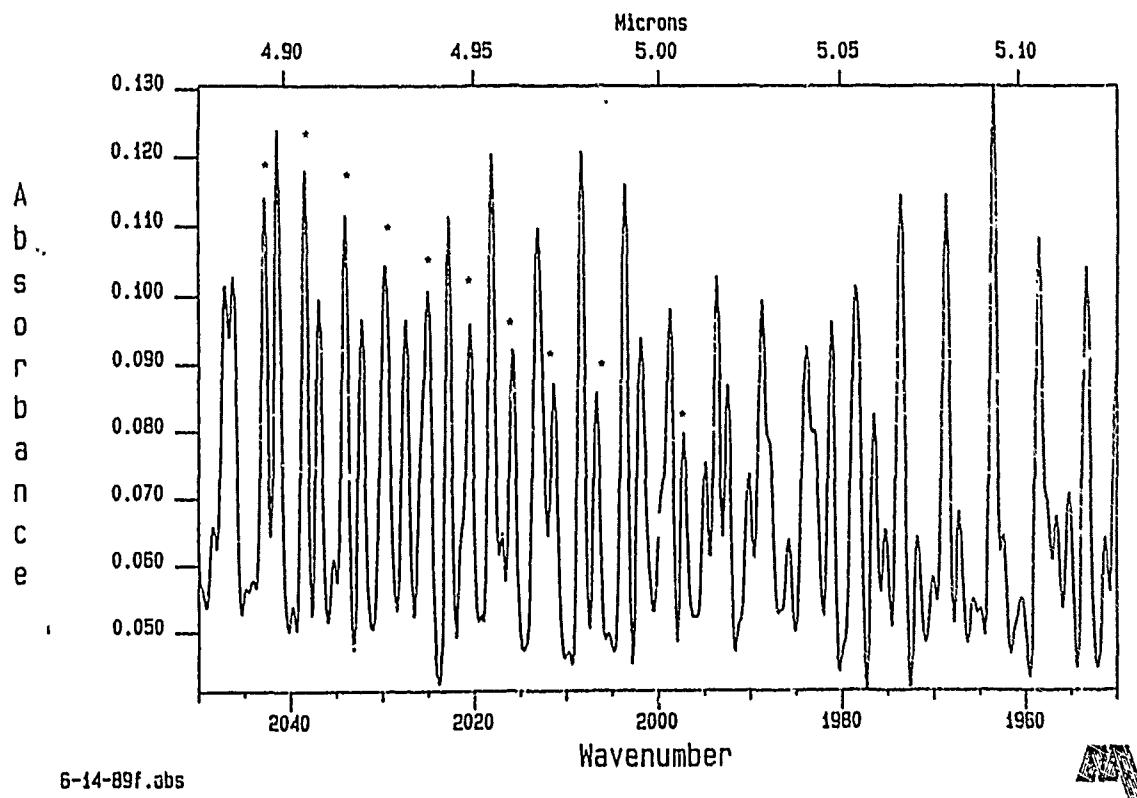


Figure 6. Spectrum of a Rich  $\text{CH}_4/\text{N}_2\text{O}$  Flame at 64 Torr. Peaks marked with \* are due to the P branch of the  $v=1-2$  transition in CO.

Table 2. Calculated CO Rotational Temperatures as a Function of the Distance Between the Center of the IR Beam and the Burner Frit Surface

| Burner Surface<br>to Beam Center<br>(cm) | Observed<br>Temperature<br>(Kelvin) | Standard<br>Deviation<br>(Kelvin) |
|------------------------------------------|-------------------------------------|-----------------------------------|
| 0.5                                      | 1524                                | 107                               |
| 0.9                                      | 1516                                | 107                               |
| 1.3                                      | 1517                                | 106                               |
| 1.7                                      | 1515                                | 113                               |
| 2.1                                      | 1518                                | 111                               |
| 2.5                                      | 1542                                | 111                               |
| 2.9                                      | 1559                                | 105                               |
| 3.3                                      | 1579                                | 107                               |
| 3.7                                      | 1639                                | 105                               |
| 4.1                                      | 1704                                | 110                               |

Figure 7 shows a plot of  $E(m)/k$  vs.  $\ln(A_{\text{obs}}^x / \{(2Da_{m,0})^{y-1} |m| F(m) v_m\})$  for the burner flame experiments. Since the observed absorbances are very weak (maximum less than 0.12), we used the approximation to true peak absorbance from observed peak absorbance for the range of true peak absorbance from 0.1 to 1.0 ( $x=1.287$ ,  $y=1.208$ ). The flame temperatures reported here are lower than adiabatic for the methane/nitrous oxide flame system.<sup>16</sup> We believe this is because the small burner chamber acted as a heat sink for the flame. No thermocouple measurements were made of the flame temperature within the burner chamber. Also, because we were unable to find published data for the  $a_m$  of the hot band spectral, we used the values for the fundamental. Finally, we made no provision for broadening of the CO lines by gases other than CO, so in effect, we were treating our data as if it were obtained in a chamber of pure CO. Our reasons for this stemmed from the many different species present in the flame and our inability at this time to quantify many of these species, particularly nitrogen.

## VI. CONCLUSION

Calculations of rotational temperatures of CO at high temperatures and of CO participating in combustion reactions by use of simple graphical methods applied to FTIR spectra have been shown to be accurate within 10% (and usually much less) of actual temperature up to 800 K. Graphical analysis of the ro-vibrational structure of the fundamental absorption of CO gives results which indicate that the simple method employed here has practical applications, especially when an accurate approximation to true peak absorbances is used. Results from spectra of flames are encouraging but more work is needed. We are presently engaged in pursuing this area of research.

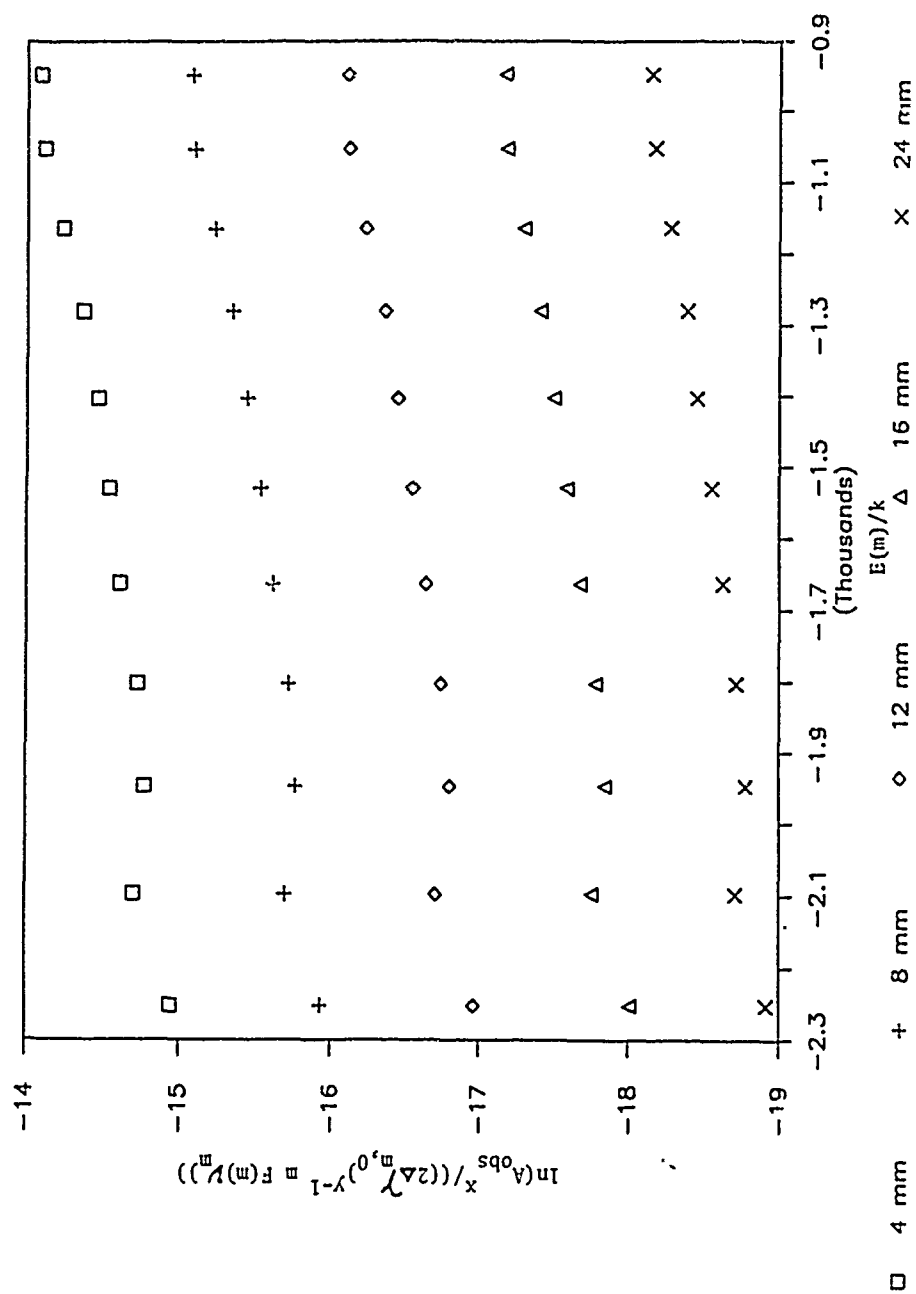


Figure 7. A Plot of  $-\ln(A_{\text{obs}}^x / \{(2\Delta\gamma_{m,0})^{y-1} [n|F(m)v_n]\})$  for CO v=1-2 P Branch Transitions in a Low Pressure (64 torr)  $\text{CH}_4/\text{N}_2\text{O}$  Flame. The abscissa for each set of data has been shifted to eliminate overlap of data sets.

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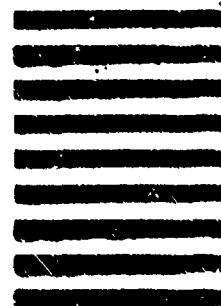
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