

# ESTIMATES OF $P_n$ AND $S_n$ ACROSS EURASIA

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## ABSTRACT

As an intermediate step to producing a integrated  $P$  and  $S$  model of the crust and upper mantle beneath most of Eurasia to predict regional station specific travel time correction surfaces, we present maps of isotropic  $P_n$  and  $S_n$  velocities across Eurasia. These maps are constructed using a groomed version of the ISC/NEIC data that are described, in part, by Engdahl et al. (1998). The locations and origin times of the events are replaced with local ground truth locations whenever possible (e.g., Sultanov et al., 1999). We use the tomographic method of Barmin et al. (2000) to estimate  $P_n$  and  $S_n$  maps on a  $1.5^\circ \times 1.5^\circ$  grid, augmented to include station and event static corrections and an epicentral distance or range correction. The starting model is CRUST5.1 (Mooney *et al.*, 1998). We call the  $P_n$  and  $S_n$  maps together with these corrections the CU  $P_n/S_n$  model. Data are used in the inversion if the residual relative to the prediction from the spherical model AK135 (Kennett et al., 1995) is less than 7.5 s for P and 15 s for S, if the event depth is within the crust or less than 50 km deep, if the azimuthal gap to all reporting stations for the event is less than 180 degrees, and if the nominal error ellipse is less than 1000 km<sup>2</sup> in area. We define the phases  $P_n$  and  $S_n$  as arriving between epicentral distances of  $3^\circ$  and  $15^\circ$ . The resulting data set consists of about 1,600,000  $P_n$  and 400,000  $S_n$  travel times inhomogeneously distributed across Eurasia. The  $P_n$  and  $S_n$  maps compare favorably with recent 3-D models  $P$  and  $S$  in the uppermost mantle estimated from different data sets (Bijwaard *et al.*, 1998; Villaseñor et al., 2000a) and with recently compiled teleseismic station corrections across the region (Engdahl and Ritzwoller, 2000). The rms misfit to the entire Eurasian data set from the CU  $P_n/S_n$  model is 1.8 s for  $P_n$  and 3.8 s for  $S_n$ . The  $P_n$  rms misfit reduces to  $\sim 1.7$  s for certain high quality subsets of the entire Eurasian data set. We present examples of travel time correction surfaces for several IMS stations or surrogates computed to epicentral distances of  $15^\circ$ . This method will be applied to compute travel time correction surfaces for the remaining IMS stations or surrogates across Eurasia. These correction surfaces will provide an important reference for 3-D models to match and extend.

**Key Words:** location, Eurasia,  $P_n$ ,  $S_n$ , correction surfaces

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## OBJECTIVE

The objective of this research is to develop and test a method to produce  $P_n$  and  $S_n$  station specific travel time correction surfaces across much of Eurasia. This method will be applied systematically to compute correction surfaces for the remaining IMS stations or surrogates across Eurasia to epicentral distances of  $15^\circ$ . These correction surfaces will provide an important reference for emerging 3-D models to match and extend.

## RESEARCH ACCOMPLISHED

### Introduction

Determination of accurate seismic locations and uncertainties is of prime importance in monitoring the Comprehensive Nuclear Test-Ban Treaty (CTBT). Small magnitude events will only be recorded at a sparse subset of the International Monitoring System (IMS) at regional distances less than  $20^\circ$  -  $30^\circ$ . Sparse network locations are subject to significant bias due to regional variations in the structure of the crust and upper mantle. To meet the goals of the CTBT for these small events, this bias must be substantially reduced in regions of significant structural variability such as that in evidence across much of Eurasia. To do so will require either a model of the 3-D structures or the effects of the structures on the relevant travel-times.

Our ultimate goal is to produce a 3-D  $P$  and  $S$  model of the crust and uppermost mantle on a  $1^\circ \times 1^\circ$  grid beneath N. Africa, the Middle East, and Central Asia that will accurately predict regional travel time variations and improve regional location capabilities. Villaseñor *et al.* (2000a) describe the construction of an  $S$  model from surface wave measurements. In order to predict regional  $S$  travel times accurately and to convert this model to  $P$ , the model will require further calibration using regional body wave travel time data. The study presented here is the first step in this direction. Our approach is not to begin by calibrating the model itself, but rather to calibrate the travel time correction surfaces (e.g., Myers and Schultz, 2000) computed from the model. To do this requires independent estimates of correction surfaces using regional  $P_n$  and  $S_n$  data. In this paper, we (1) discuss the data set and method used to estimate these correction surfaces, (2) show the first continental scale images of  $P_n$  and  $S_n$  variations across Eurasia, (3) display the resulting correction surfaces for several IMS surrogate stations, and (4) discuss the fit to the regional phase data.

### Data

$P_n$  and  $S_n$  travel times are taken from a groomed version of the ISC and NEIC data bases described, in part, by Engdahl *et al.* (1998). ISC travel times are for events that occurred from 1964 through 1997 and NEIC data are from 1998 and 1999. The locations of explosions are replaced with “ground truth” locations whenever possible (e.g., Sultanov *et al.*, 1999). We define the phases  $P_n$  and  $S_n$  as arriving between epicentral distances of  $3^\circ$  and  $15^\circ$ . At epicentral distances beyond about  $10^\circ$ , however,  $P_n$  and  $S_n$  may dip into the mantle. The depth of penetration will depend on the radial gradient of velocity, which will vary spatially. Truncation of the data set to include rays only if epicentral distances are less than  $\sim 12^\circ$ , as in some other studies (e.g., Hearn and James, 1994), severely restricts path coverage in some areas of Eurasia. To utilize longer paths it is desirable to correct for the effect of ray penetration into the uppermost mantle. We discuss the range correction in the next section entitled “Method”. This data set totals 3,672,268  $P_n$  phases and 1,346,676  $S_n$  phases for 5,418 stations and 149,929 events worldwide. Data are used in the inversion if the residual relative to the prediction from the spherical model AK135 (Kennett *et al.*, 1995) is less than 7.5 s for P and 15 s for S, if the event depth is within the crust or less than 50 km deep, if the azimuthal gap to all reporting stations for the event is less than 180 degrees, and if the nominal error ellipse is less than 1000 km<sup>2</sup> in area. These selection criteria reduce the data set to 1,636,430  $P_n$  and 493,734  $S_n$  phases worldwide. Most of these paths cross Eurasia ( $0^\circ$  -  $80^\circ$ N latitude and  $-10^\circ$ W -  $180^\circ$ E longitude plus a buffer zone), the numbers being 1,257,052 for  $P_n$  and 422,634 for  $S_n$ .

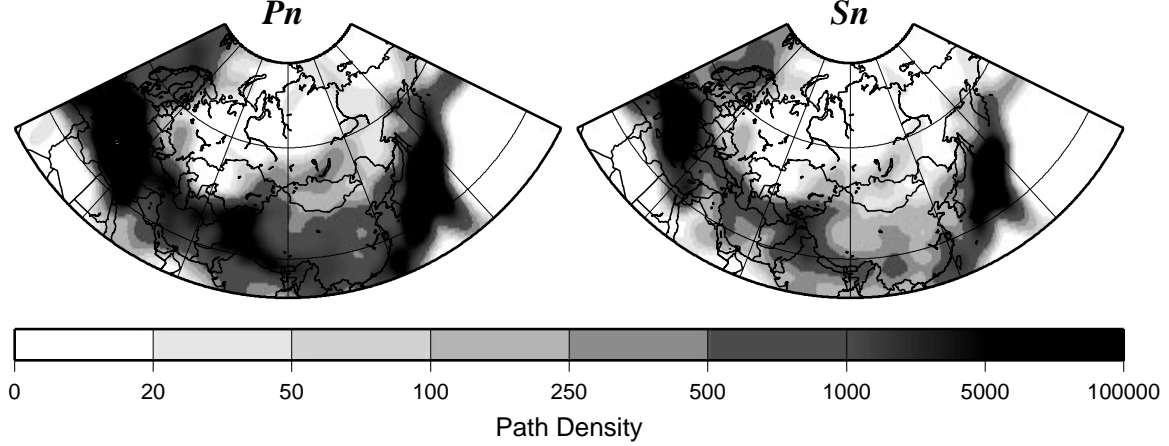


Figure 1: Path density for the  $P_n$  and  $S_n$  data used here, defined as the number of paths intersecting a  $2^\circ \times 2^\circ$  cell ( $\sim 50,000 \text{ km}^2$ ). Because path lengths for these phases are by definition short ( $< 15^\circ$ ), paths only exist in regions where both sources and receivers are common. Thus, the path distribution is highly heterogeneous across the region.

### Method

The  $P_n$  and  $S_n$  maps are defined over a two-dimensional surface and, therefore, may be estimated with the same method we developed for surface wave tomography (Barmin *et al.*, 2000). In this method, the model is constructed on an equally spaced grid such that the following figure-of-merit is minimized:

$$(\mathbf{G}\mathbf{m} - \mathbf{d})^T \mathbf{C}^{-1} (\mathbf{G}\mathbf{m} - \mathbf{d}) + \sum_{k=0}^n \alpha_k^2 \|F_k(\mathbf{m})\|^2 + \sum_{k=0}^n \beta_k^2 \|H_k(\mathbf{m})\|^2, \quad (1)$$

which is a linear combination of data misfit, model roughness, and the amplitude of the perturbation to a reference model. The vector  $\mathbf{m}$  represents the estimated model which is a perturbation relative to a reference across the region of interest,  $\mathbf{G}$  is the forward operator that computes travel time from the estimated model,  $\mathbf{d}$  is the data vector,  $\mathbf{C}$  is the data covariance matrix or matrix of data weights,  $F$  is a Gaussian smoothing operator, and  $H$  is an operator that penalizes the norm of the model in regions of poor path coverage. The spatial smoothing operator is defined over the 2-D model as follows

$$F_k(\mathbf{m}) = m_k(\mathbf{r}) - \int_S S_k(\mathbf{r}, \mathbf{r}') m_k(\mathbf{r}') d\mathbf{r}', \quad (2)$$

where  $S_k$  is a smoothing kernel:

$$S_k(\mathbf{r}, \mathbf{r}') = K_{0k} \exp\left(-\frac{|\mathbf{r} - \mathbf{r}'|^2}{2\sigma_k^2}\right) \quad (3)$$

$$\int_S S_k(\mathbf{r}, \mathbf{r}') d\mathbf{r}' = 1, \quad (4)$$

and  $\sigma_k$  is the spatial smoothing width or correlation length. The minimization of the expression in equation (2) explicitly ensures that the estimated model approximates a smoothed version of the model. The maps are estimated on a  $1.5^\circ \times 1.5^\circ$  grid across Eurasia.

Barmin *et al.* (2000) discuss tomographic inversions of group and phase velocity data in detail. The inversion for  $P_n$  and  $S_n$  is essentially the same except that from the reference crustal model we compute source and receiver side Moho penetration points and use these points as the starting and ending points of the ray during inversion.  $P_n$  and  $S_n$  rays are assumed to be horizontal in a spherical mantle. We have used CRUST5.1 as the reference model in the crust and for  $P_n$  and  $S_n$ . In addition, we estimate a

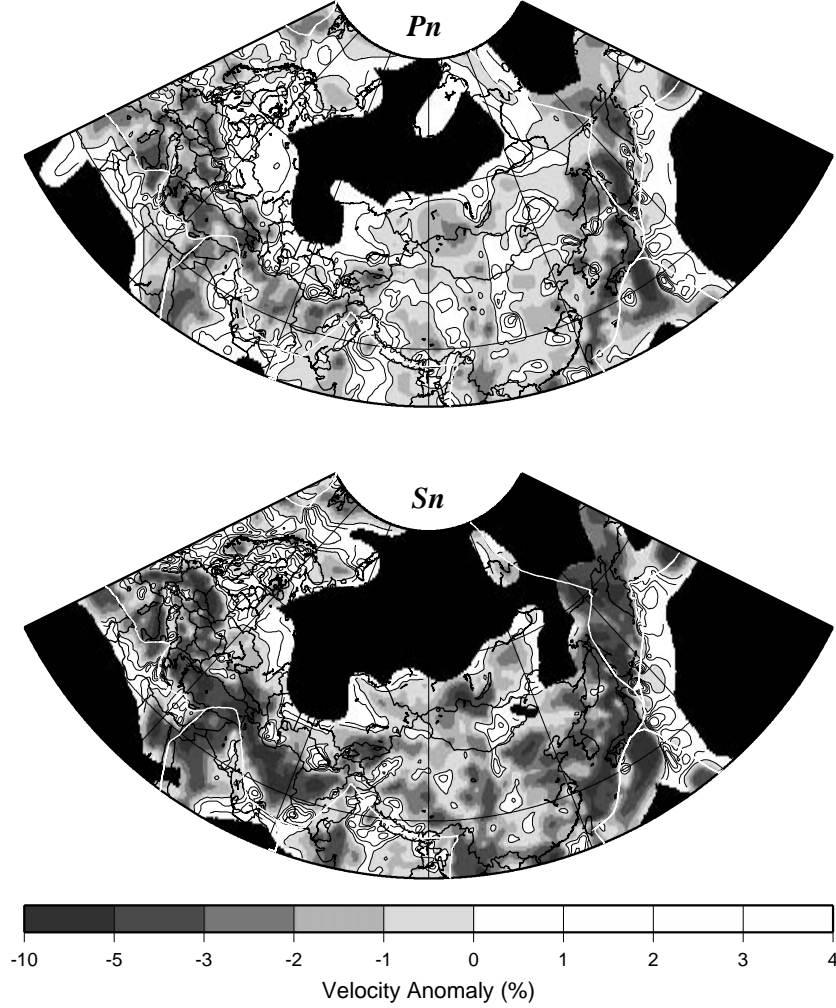


Figure 2:  $P_n$  and  $S_n$  maps estimated across Eurasia represented as the deviation from the mean across each map in percent.  $P_n$  is relative to 8.052 km/s, and  $S_n$  is relative to 4.583 km/s. We consider regions in which path density is less than 20 paths/50,000km<sup>2</sup> to be unconstrained by the data, and these regions are shaded black.

range correction and, following Hearn and collaborators (e.g., Hearn and Clayton, 1986; Hearn *et al.*, 1991; Hearn and James, 1994; and elsewhere), we also estimate event and station statics. The range correction is designed to compensate for the fact that rays are not horizontal in the uppermost mantle, but systematically dip further into the mantle as propagation distances increase. This allows us to fit data to greater ranges than is possible without this correction. With this correction, the resulting tomographic maps agree well with those produced with short path data alone (epicentral distances less than 10°) in those regions where tomographic maps can be constructed reliably using only the short path data.

The observed travel time,  $t_{obs}$ , is, therefore, modeled as follows:

$$t_{obs} = t_m + t_{crust\_sta} + t_{crust\_evt} + \delta t_{sta} + \delta t_{evt} + \delta t(\Delta) + \int_p \frac{ds}{v_m}, \quad (5)$$

where  $t_m$  is the predicted travel time for horizontal rays through the mantle part of the reference model CRUST5.1, the contributions to the travel time due to the reference crustal model CRUST5.1 on the source and event sides are  $t_{crust\_sta}$  and  $t_{crust\_evt}$ , the station and event statics are  $\delta t_{sta}$  and  $\delta t_{evt}$ ,  $\delta t(\Delta)$  is the distance correction,  $v_m$  is the estimated velocity at the top of the mantle, and  $p$  denotes the appropriate path linking source and receiver through the mantle. Thus,  $t_m$ ,  $t_{crust\_sta}$ ,  $t_{crust\_evt}$ , and  $p$  are predicted by the

reference model and  $\delta t_{sta}$ ,  $\delta t_{evt}$ ,  $\delta t(\Delta)$ , and  $v_m$  are estimated. A static correction is estimated for a station if there are phase picks from at least 7 events made at that station and an event correction is estimated for all events for which there are at least 20 reporting stations. The asymmetry in this condition is due to the fact that there are more physical phenomena modeled with the event static than with the station static (e.g., mislocation, origin time error). The station and event statics are undamped at present.

We will refer to the  $P_n$  and  $S_n$  maps together with the corrections  $\delta t_{sta}$ ,  $\delta t_{evt}$ , and  $\delta t(\Delta)$  as the CU  $P_n/S_n$  model to distinguish it from our recent 3-D models (e.g., Villaseñor *et al.*, 2000a).

### $P_n$ and $S_n$ Tomography

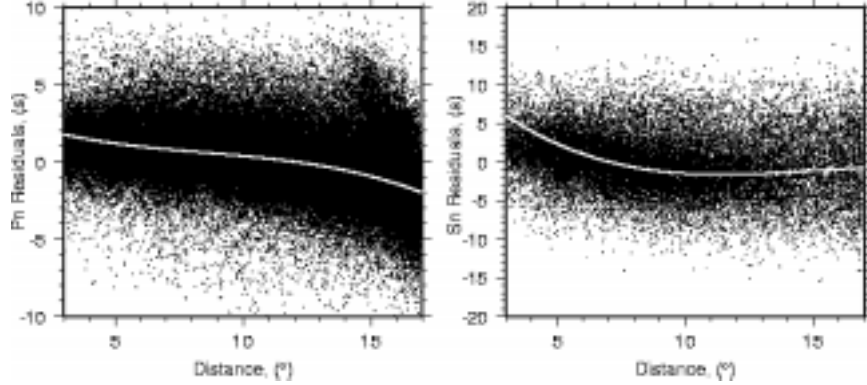


Figure 3: Raw residuals computed for the entire data set using only the  $P_n$  and  $S_n$  maps shown in Figure 2; these residuals do not include static or range corrections. The white lines are cubic polynomial fits to these data and define the range corrections  $\delta t(\Delta)$ .

**Table 1. Summary of the distribution of station and event statics world-wide.**

<i>static correction</i>	$P_n$		$S_n$	
	<i>station (#)</i>	<i>event (#)</i>	<i>station (#)</i>	<i>event (#)</i>
total	5418	149929	3781	99385
with correction	3295	73587	1780	23737
no correction	2123	76332	2001	75648
max. correction (s)	(-3.78,5.36)	(-6.156,6.89)	(-6.58,10.12)	(-14.56,16.51)
# < -2.0 (s)	4	1396	65	2534
-2.0 < # < -1.5 (s)	14	1884	49	1160
-1.5 < # < -1.0 (s)	57	3781	99	1516
-2.0 < # < -0.5 (s)	177	6801	161	1856
-0.5 < # < 0.0 (s)	580	9890	214	2239
0.0 < # < 0.5 (s)	927	12061	276	2494
0.5 < # < 1.0 (s)	796	12959	231	2358
1.0 < # < 1.5 (s)	425	10913	172	2186
1.5 < # < 2.0 (s)	178	6897	138	1867
# > 2.0 (s)	137	7015	375	5527

The estimated  $P_n$  and  $S_n$  maps are shown in Figure 2. The velocity anomalies correlate well with known tectonic features and the  $P_n$  and  $S_n$  maps compare favorably with one another and with the patterns of velocity variations in the 3-D shear velocity model of Villaseñor *et al.* (2000a) and the teleseismic  $P$  wave model of Bijwaard and Spakman (1998). Because our tomographic method penalizes the amplitude of the maps in regions of poor data coverage and the estimated maps are perturbations to a reference state, the maps revert to the reference model where data coverage is poor; i.e., less than 15 - 20 paths per each  $2^\circ \times 2^\circ$  cell. The areas of poor data coverage are identified as black regions in Figure 2. Both the  $P_n$  and  $S_n$  maps demonstrate poor coverage across the shield and platform regions of northern Russia and Kazakhstan, in the

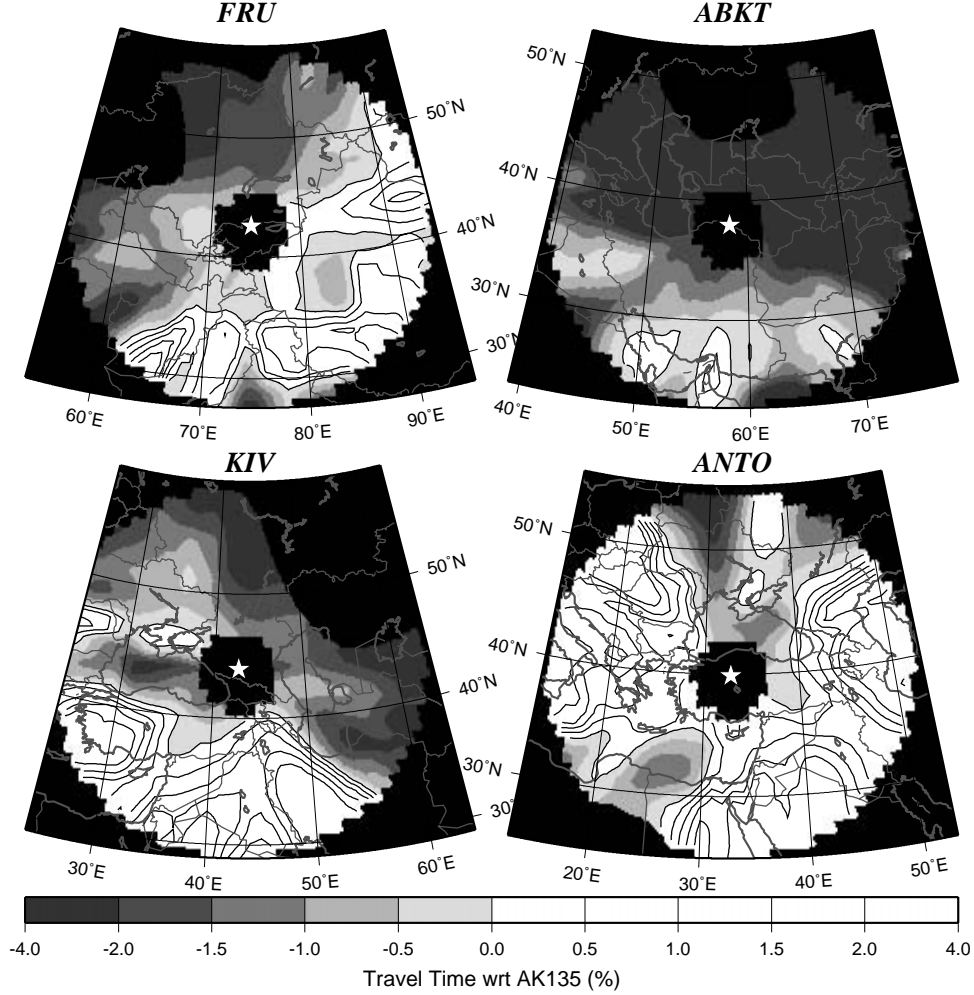


Figure 4: Station specific travel time correction surfaces for  $P_n$  for four IMS stations or surrogates ( $FRU \rightarrow AAK$ ,  $ANTO \rightarrow BRAR$ ,  $KIV \rightarrow KBZ$ ). Surfaces exist only where  $P_n$  data density is locally greater than 20 paths per 50,000  $\text{km}^2$  and to  $15^\circ$  from the station. Low data density regions are shaded black.

oceans, and across North Africa. Elsewhere the spatial resolution of the maps is estimated to be between 150 - 300 km.

The range correction is shown in Figure 3. This figure shows that mispicks from the triplication from the 400 km discontinuity begin to emerge in the data set at about  $15^\circ$ , which is the reason we have limited the  $P_n$  and  $S_n$  maps to this distance. The number of  $S_n$  picks reduces sharply after about  $10^\circ$ . Station and event statics,  $\delta_{sta}$  and  $\delta_{evt}$ , are summarized in Table 1 for the entire data set world-wide. Although for  $P_n$  we estimate static corrections only for about 60% of the stations and half of the events world-wide and for  $S_n$  the numbers are about 50% and 25%, respectively, the great majority of the measurements emanate from events and are recorded at stations that have static corrections. This is particularly true for  $P_n$ , where only 3% of the measurements are made at stations without static corrections and 12% of the measurements are for events without event corrections. For  $S_n$ , the numbers are 4% and 33%, respectively. Thus, most measurements have the full complement of corrections applied. Not surprisingly, statics for  $S_n$  are typically larger than for  $P_n$  and event statics are larger than station statics. This is because  $S$  variations in the crust and upper mantle are typically larger than  $P$  by about a factor of two and, in addition to compensating for errors in the crustal model, event statics also compensate for mislocations and origin time errors. Event statics also are more effective at reducing rms misfit. Although not shown here, the  $P_n$  and  $S_n$  station statics are geographically coherent and correlate with one another with a poorly determined  $S/P$  ratio of about 1.8.

However, because the statics are determined simultaneously their values are interdependent and difficult to interpret in abstraction from the others. In addition, because we have not constrained the values of the static corrections, in some cases they are probably too large in magnitude, particularly  $S_n$  event statics.

Figure 4 displays station specific travel time correction surfaces for several IMS stations or surrogates. These surfaces are for surface sources with the local elevation observed at the station elevation. This differs from correction surfaces as they are commonly displayed in which both the source and station are on the reference ellipsoid. The correction surface for ANTO (Ankara, Turkey) compares favorably with the surfaces from two nearby stations in Turkey (KAS, KVT) reported by Myers and Schultz (2000) using a different method.

Fit to the Data

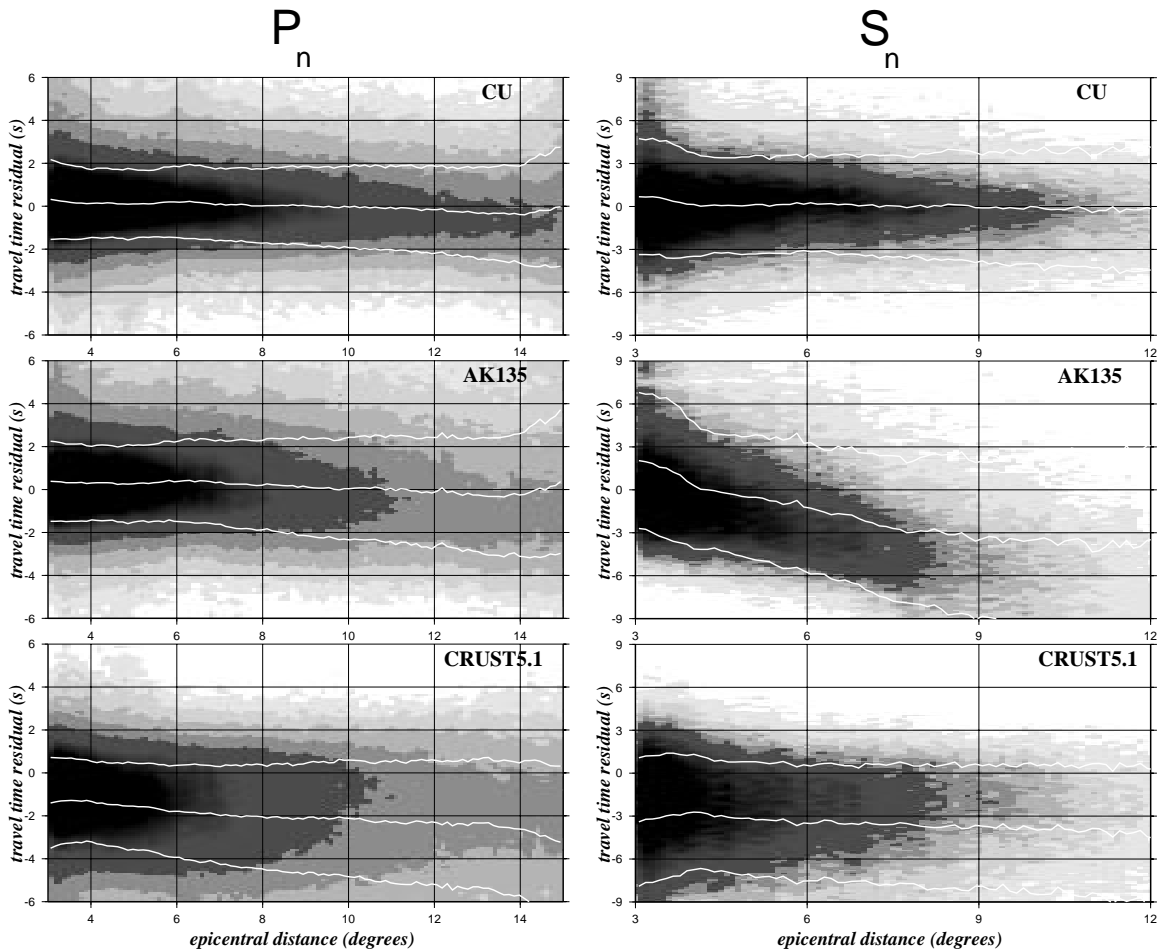


Figure 5: Shaded plots of the density of  $P_n$  and  $S_n$  travel time residuals (observed - predicted) for the Eurasian data set presented versus epicentral distance. Results for three models are shown: (TOP) the CU  $P_n/S_n$  model, (MIDDLE) the spherically symmetric model AK135 (Kennett *et al.*, 1995), and (BOTTOM) the laterally heterogeneous crustal,  $P_n$ , and  $S_n$  model CRUST5.1 (Mooney *et al.*, 1998). Darker shades indicate larger number of residuals and the white lines show the local mean and  $\pm 1\sigma$ . Means and standard deviations are summarized in Table 2.

Table 2. Summary of misfits to the whole Eurasian data set displayed in Figure 5.

model	$P_n$		$S_n$	
	mean (s)	$\sigma$ (s)	mean (s)	$\sigma$ (s)
CU $P_n/S_n$	0.06	1.81	0.12	3.83
AK135	0.23	2.06	-0.94	5.36
CRUST5.1	-1.80	2.37	-4.03	5.14

Misfits to the entire Eurasian data set for  $P_n$  and  $S_n$  are shown in Figure 5 and summary statistics are presented in Table 2. The standard deviation  $\sigma$  reported in Table 2 is computed relative to the mean, so total rms misfit must incorporate the mean. In general, short distance offsets result from errors in the crustal model either in average crustal velocities or Moho depths. Errors in the uppermost mantle velocities manifest themselves as trends with distance. The model AK135 does very well for  $P$ . Improvements afforded by the CU  $P_n/S_n$  model over AK135 on average begin only after about  $8^\circ$ . However, on average, for both  $P$  and  $S$  CRUST5.1 is too slow in the crust whereas AK135 is fine for  $P$  but too fast for  $S$ . CRUST5.1 displays only weak distance trends which indicates that the average upper mantle velocities are reasonable, but AK135 is too slow for  $S_n$ , demonstrating a very strong trend. This indicates that uppermost mantle  $S$  velocities in AK135 are too slow on average for Eurasia or the depth gradient in the uppermost mantle is not right. The overall rms misfit from our inversions is 1.8 s across all of Eurasia for  $P_n$  and approximately twice this value for  $S_n$ . These misfit statistics appear to be consistent with those reported by Myers and Schultz (2000) in a study limited to the neighborhood of the Racha earthquake sequence (Caucasus Mountains).

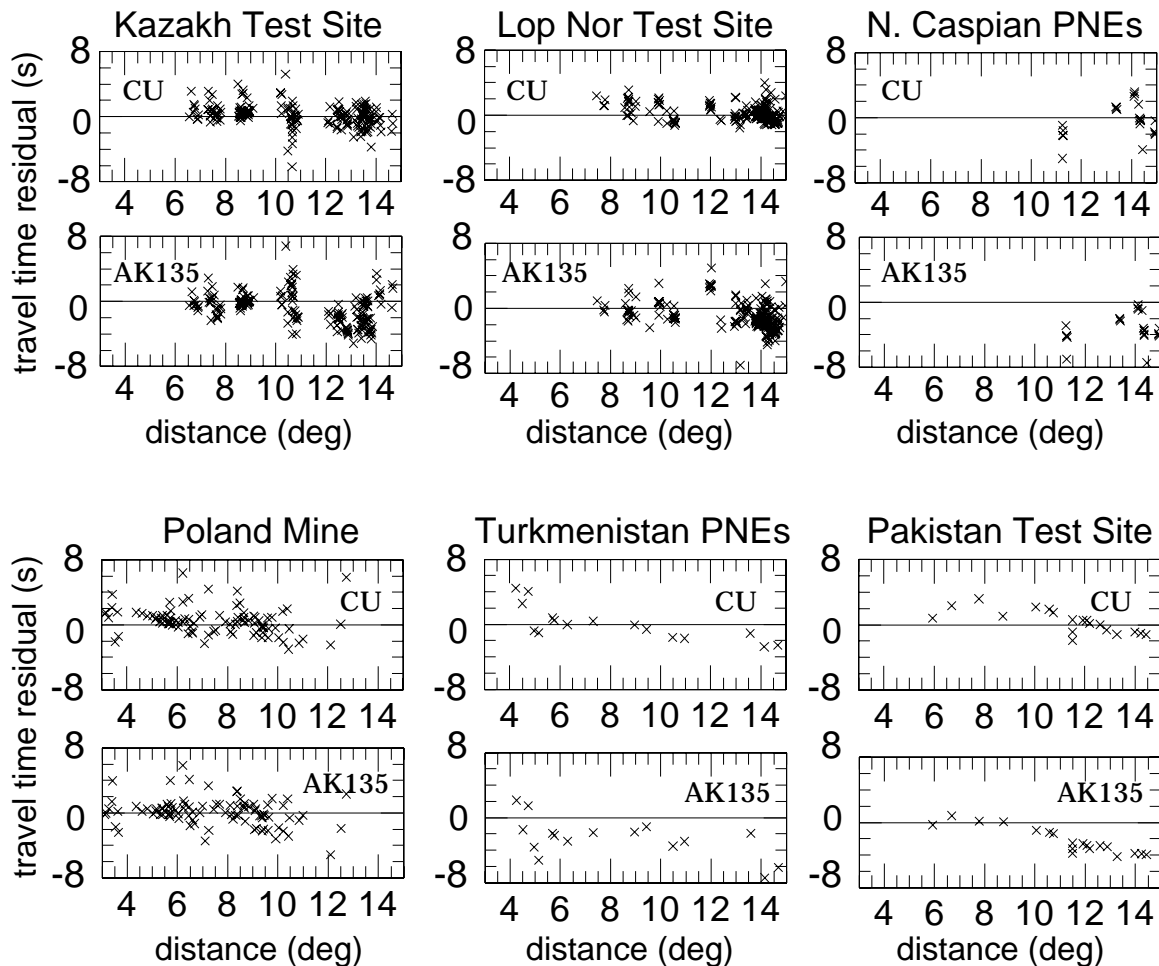


Figure 6: Misfits to  $P_n$  and  $S_n$  measurements from selected explosions that occurred at six source locations. The upper of each pair of plots for each source location is for the CU  $P_n/S_n$  model and the lower of each pair is for the model AK135. Summary statistics are presented in Table 3.

The entire Eurasian data set is very noisy and many locations and origin times are poorly known. A better estimate of expected misfit may come from explosion data in which in many cases the epicenter is well constrained, although the origin times may not be. Figure 6 displays misfits to data from six explosion regions (3 test sites, 2 PNE locations, and 1 large mine) and Table 3 summarizes these data. Only explosions with  $m_b \geq 4.6$  as reported in the PDE are used. With the exception of the large mine explosion in south-

western Poland, these events are not observed at enough stations with regional phases to have event statics. The overall misfit of the CU  $P_n/S_n$  model to these explosion data is 1.68 s for  $P_n$ , which is slightly reduced over the entire Eurasian data set. For events at the three test sites, however, misfit is better than 1.45 s for  $P$  and the fit afforded by AK135 is considerably worse. Again, most of the improvement over AK135 delivered by the CU  $P_n/S_n$  model comes for paths longer than about  $8^\circ$ . A similar rms misfit of 1.7 s results from a subset of the complete Eurasian data set that consists only of events with  $m_b \geq 4.6$  and measurements from events with an event correction measured at stations with a station correction.

**Table 3. Summary of misfits to explosion data for  $P_n$  displayed in Figure 6.**

<i>Location</i>	# meas.	# stations	# events	rms misfit (s)	
				CU $P_n/S_n$	AK135
Kazakh Test Site	189	18	42	1.38	2.08
Lop Nor Test Site	204	39	19	1.16	1.90
N. Caspian PNEs	22	7	6	2.10	3.72
Pakistan Test Site	19	18	2	1.43	2.72
Poland Mine	92	92	1	1.72	1.68
Turkmenistan PNEs	16	16	1	2.04	3.45
Total	542	-	-	1.68	2.71

## CONCLUSIONS AND RECOMMENDATIONS

### Conclusions

The method for producing  $P_n$  and  $S_n$  maps with associated parametric corrections appears to effectively summarize the information in the large groomed ISC/NEIC data base. The  $P_n$  and  $S_n$  maps correlate well with other high resolution information about structural variations in the uppermost mantle. The station specific travel time correction surfaces computed from the CU  $P_n/S_n$  model appear to be robust and fit the data with very low levels of bias at epicentral distances from  $3^\circ$  to  $15^\circ$ . Overall rms misfits across Eurasia for  $P_n$  are  $\sim 1.8$  s and for  $S_n \sim 3.8$  s and are better for data subsets chosen for their quality (e.g., explosions, large magnitude events, independent information about epicenter location and/or origin time). These misfits are considerably better than those produced by AK135 and CRUST5.1. The correction surfaces will provide an important reference for 3-D models to match and extend.

### Recommendations

- **Continued development of the method.** Small modifications to the method may continue to improve the results. For example, although a better starting model of the mantle will not appreciably change the estimated  $P_n$  and  $S_n$  maps it may help to stabilize the static corrections and the range correction. Also, the range correction, which exists to compensate for ray penetration into the mantle, may be improved if it depended on tectonic region because vertical gradients probably vary between tectonic regions.
- **Systematic application of this method to compute station specific travel time correction surfaces across Eurasia.** The method should be applied systematically across all of Eurasia to all IMS stations and surrogates as a reference for future studies.
- **Use of these correction surfaces to calibrate the correction surfaces computed from emerging 3-D models.** The computation of correction surfaces from 3-D models is technically straightforward (e.g., Villaseñor *et al.*, 2000b), but there are practical problems. For example, there are few constraints on variations in crustal velocities across most of Eurasia and the mantle model is probably more accurate in  $S$  than in  $P$ . Thus, raw predictions from existing models such as that of Villaseñor *et al.* (2000a) are not likely to be as accurate as desired. The correction surfaces shown here will help to calibrate the surfaces computed from 3-D models.

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