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14. ABSTRACT Modern optical components such as aspheres are difficult to inspect. Interferometry, the usual solution, lacks the required dynamic range. Too many fringes are produced. Adapting interferometry for null measurements currently necessitates special optical components for each inspected component. In this effort, a programmable Spatial Light Modulator (SLM) was incorporated into an optical inspection instrument to boost the dynamic range and enable null testing on almost any component using off the shelf optics. The instrument was used to measure the surface accuracy of a 6 inch tall cone with a top diameter of 8.4 inches and a bottom diameter of 8.2 inches. Measurement simulations were compared with test data and the system errors were investigated. A procedure was also developed and tested to subtract system aberrations. A dynamic range of 150 wavelengths was demonstrated.					
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Enhanced Interferometry with a Programmable Spatial Light Modulator

Presented at
Mirror Technology SBIR/STTR Workshop
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Millennium Hotel, Boulder CO

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(POC's: Geraldine Wright and David Content)

This presentation is approved for General Public release



Presentation Summary

- Summarize the problem & innovation
- Describe the hardware and software
 - Digital Interferometry with preconditioned wavefronts
 - Hybrid Hartmann/Digital Interferometry
- Typical measurements
- Potential Applications
- Future Work



The Problem Being Addressed

Modern optical components such as aspheres are difficult to inspect. Interferometry, the usual solution, lacks the required dynamic range. Too many fringes are produced. Adapting interferometry for null measurements currently necessitates special optical components for each inspected component.



Innovation/Solution

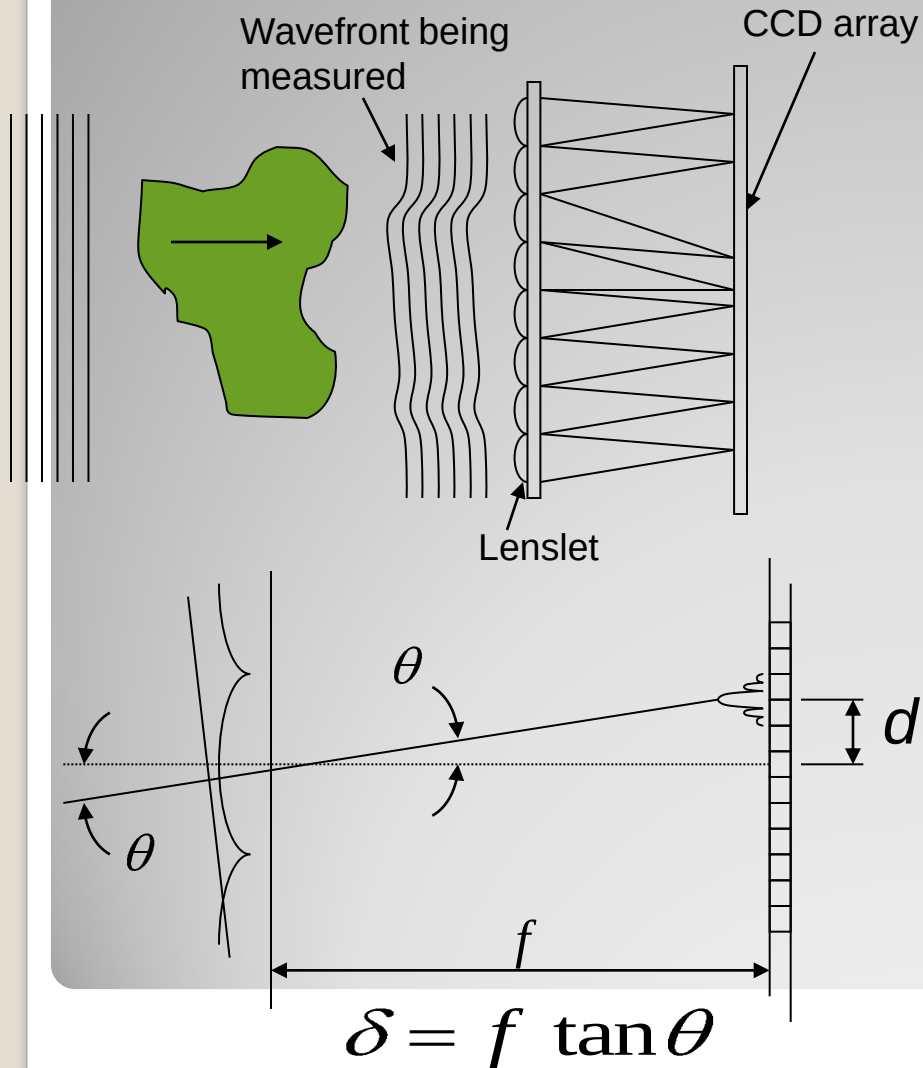
- Incorporate a programmable SLM into an optical inspection instrument enabling:
 - Combined interferometry and Hartmann Sensing
 - Virtually unlimited dynamic range
 - Extended dynamic range of interferometry by preconditioning waves
 - Null testing on almost any component using off the shelf optics.
- Incorporate Instantaneous Digital Interferometry Technology



Challenges Faced and Solved

- **Incorporating the SLM into a PhaseCam**
- **Integrating Hartmann and Digital Interferometry**
- **Calibrating the SLM**
 - **Phase-only mode by controlling polarization**
 - **Corrected gamma curve, i.e. linear phase shift versus grayscale value**
- **Impressing the required phase function on the SLM**
- **Identifying and minimizing errors and noise**

Conventional Shack-Hartmann Characteristics



- Lenslet diameters, d , define spatial resolution over the wavefront being measured.
- δ (sensitivity) proportional to f , which should be less than d to prevent confusion

Without aberration



With aberration



'Measuring aberration of the eye with wavefront technology'

Giuseppe Colicchi, et al

Zur Veröffentlichung eingereicht bei "Physics Education", 2006

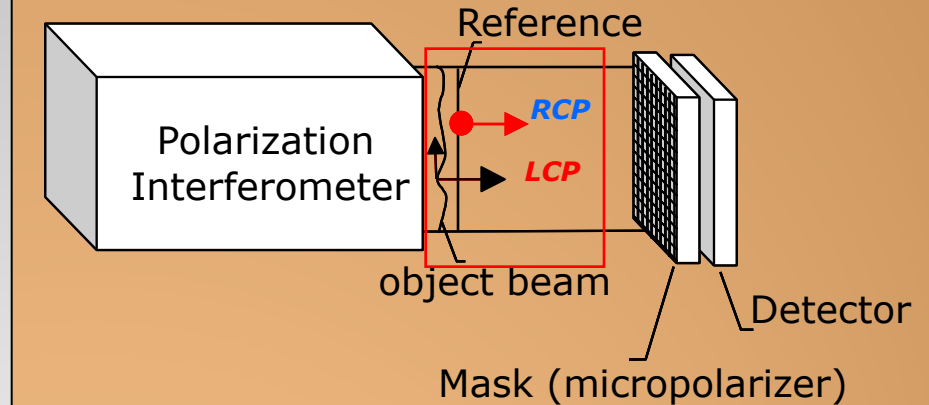
Key Components: SLM, Pixelcam*

SLM for a compensator

- Programmable holographic optical element
- Produce wavefronts of any shape and can simulate freeform optical surfaces
- Holoeye SLM
 - SLM can produce a phase up to 2π at 632.8nm
 - Assign 0 to 255 grayscale values to 0 to 2π (or 1λ)
 - Can generate higher phases by wrapping phase
 - Can provide more than 150 wave tilt
 - Pixel size ~ 19 microns.
 - 19.5 x 14.6mm size (1024 x 768 pixels)

Pixelated Phasecam for a detector

- Spatial phase shifting interferometer
- Single shot, insensitive to vibration



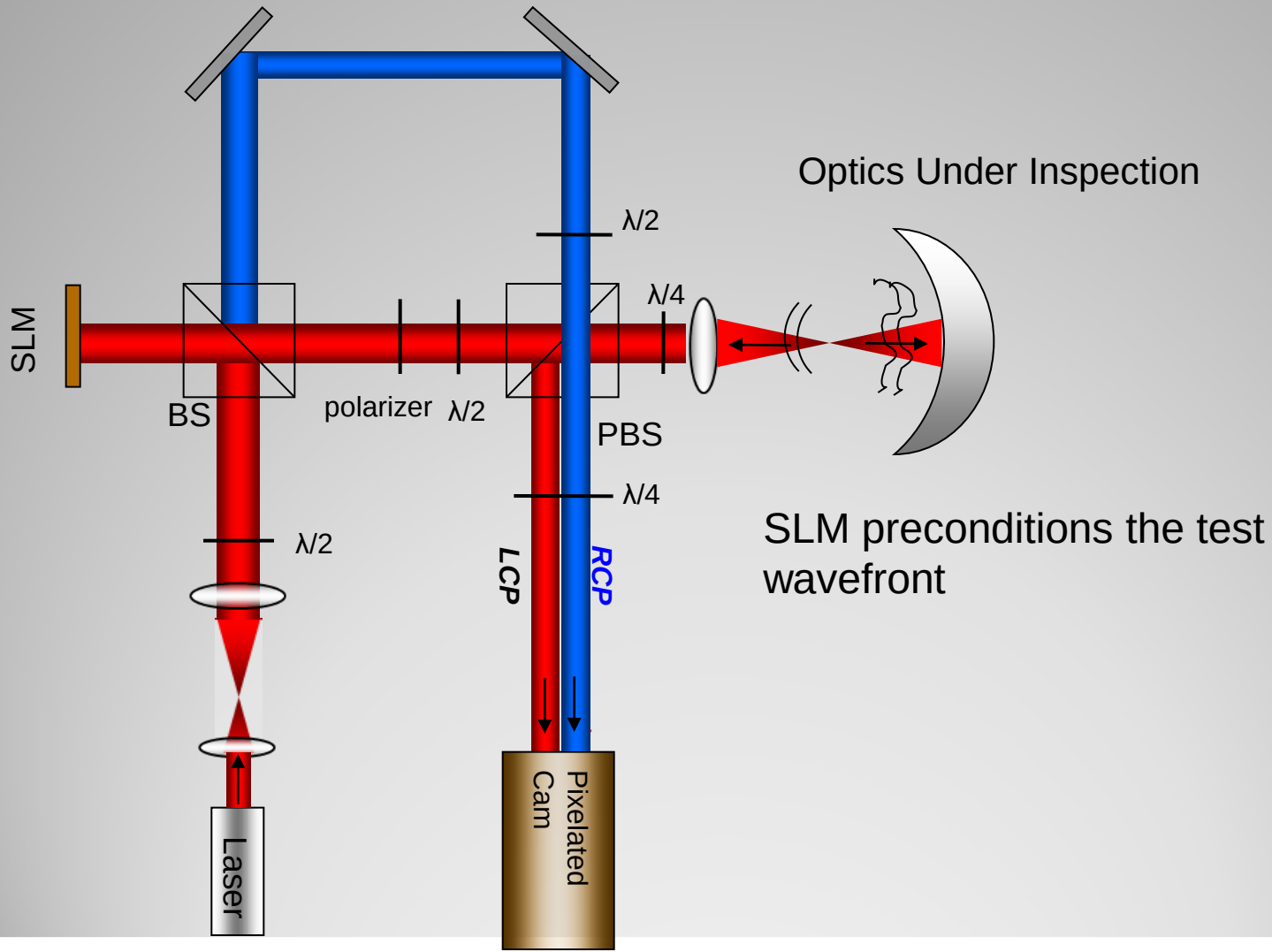
0	90
270	180

Phase information of the object, $\Delta\phi(x, y)$ can be obtained from the 4 intensities on each unit cell.

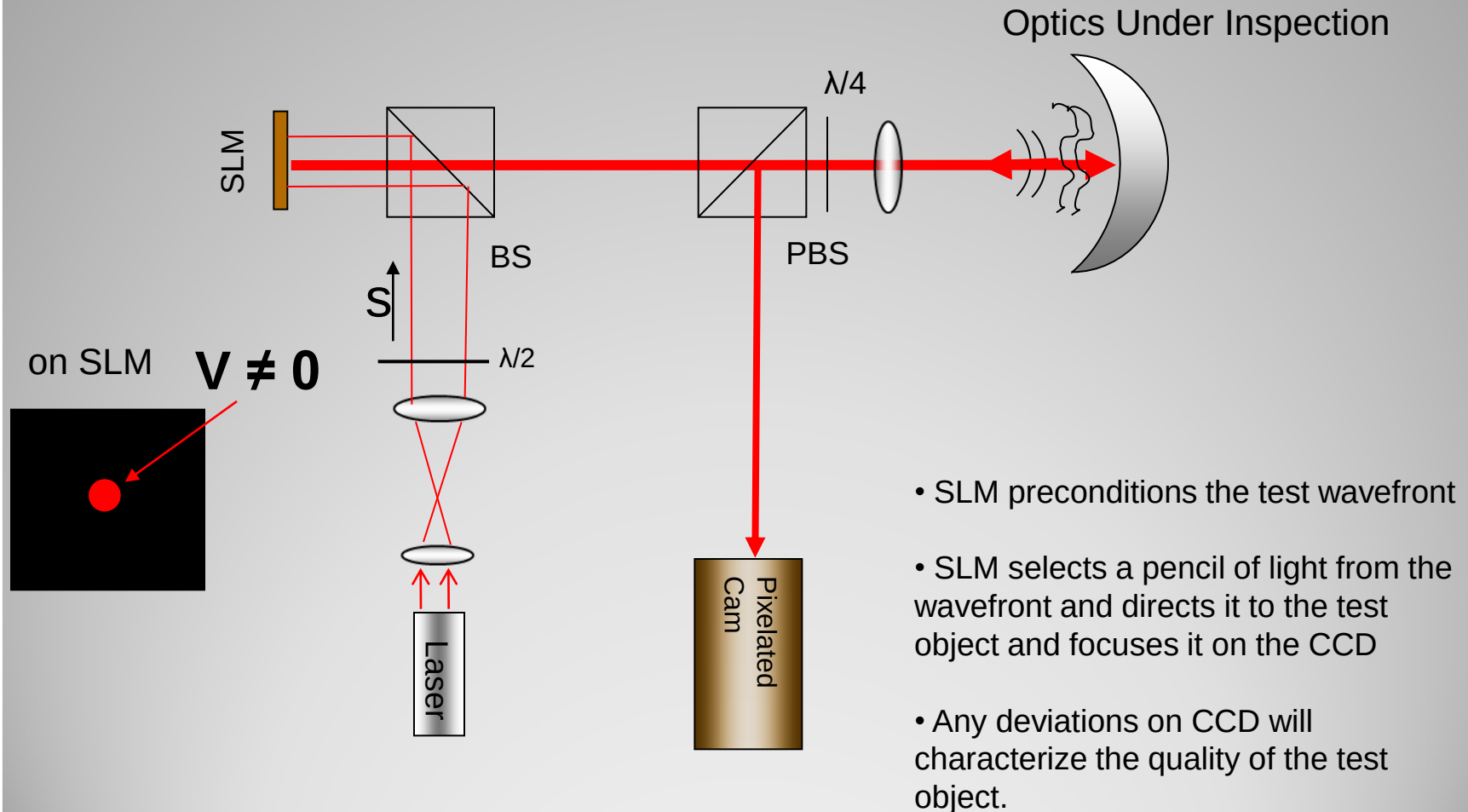
$$I(x, y) = \frac{1}{2} (I_r + I_s + 2\sqrt{I_r I_s} \cos(\Delta\phi(x, y) + 2\alpha_p))$$

*Produced and Trademark by 4D Technologies, Inc, Tucson, AZ <http://www.4dtechnology.com>

Hybrid Hartmann & Digital Interferometer

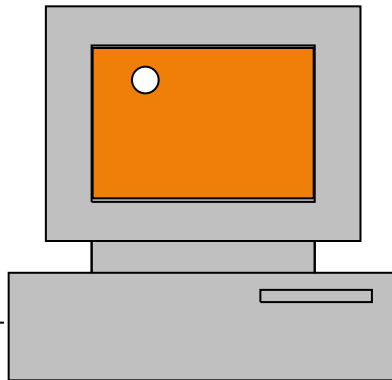


Using a SLM as a Scanning Shack-Hartmann Component



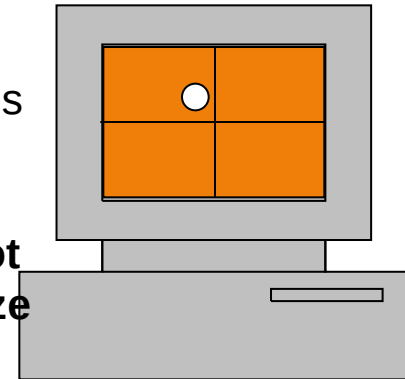
Scanning Shack-Hartmann System

Control Screen

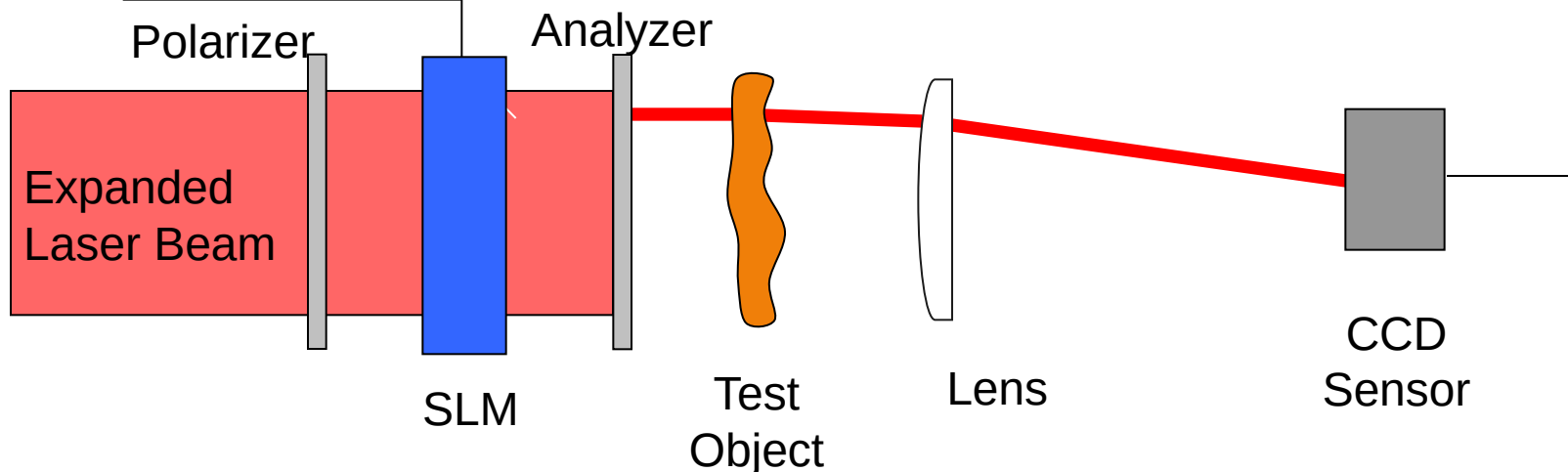


Control screen is a map of a scanned aperture (or pencils light) of the SLM.

Data Screen

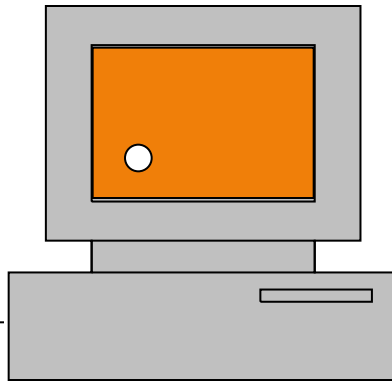


No Overlap of Focused Spot
Programmable Aperture Size



Scanning Shack-Hartmann System

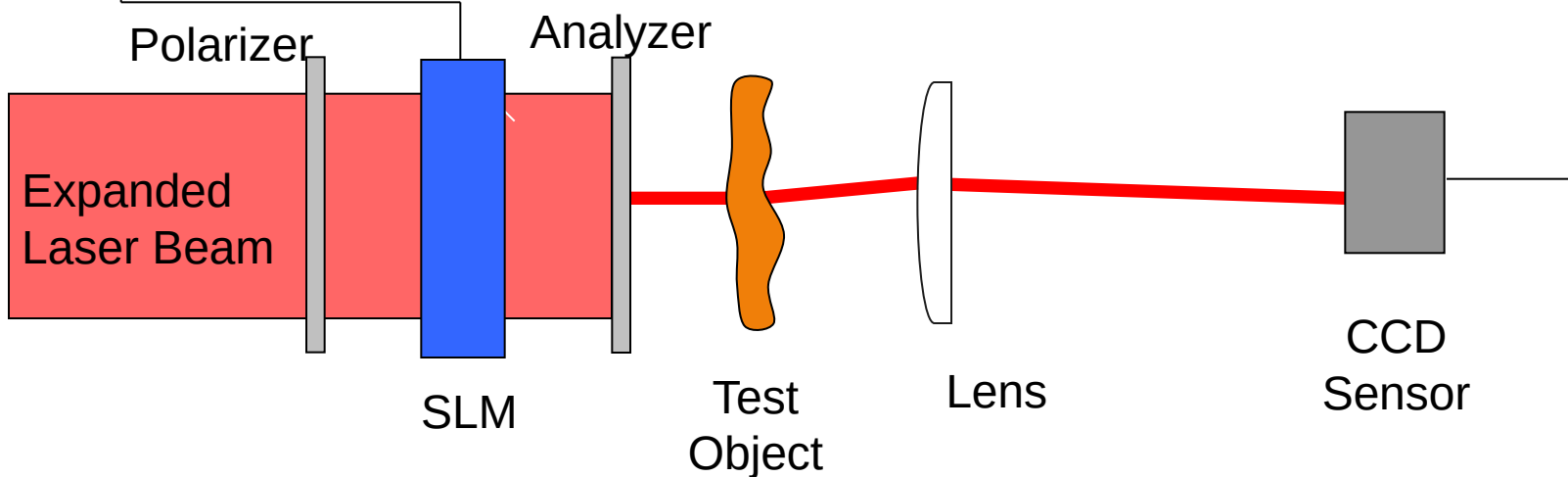
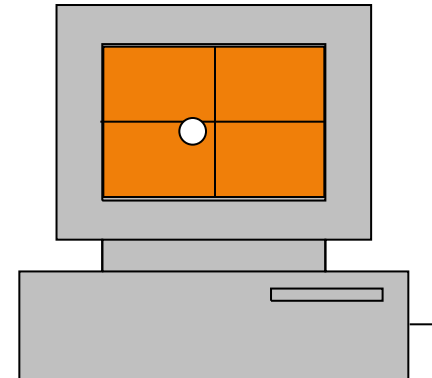
Control Screen



'Pencils' of light are deflected by the test object's slope of wavefront as they scan the object.

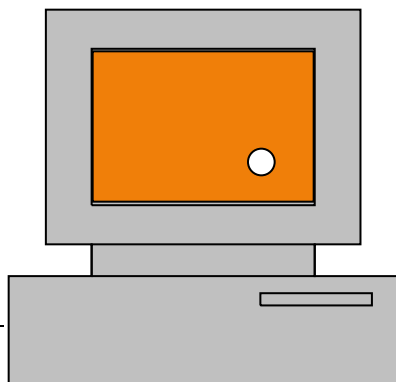
SLM corrects Aberrations

Data Screen



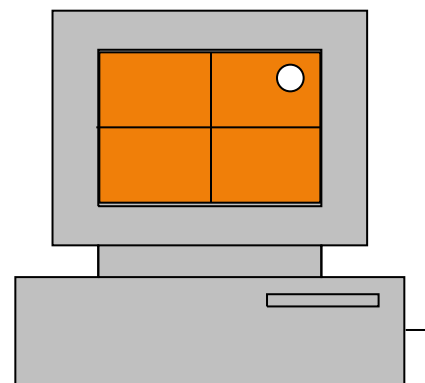
Scanning Shack-Hartmann System

Control Screen

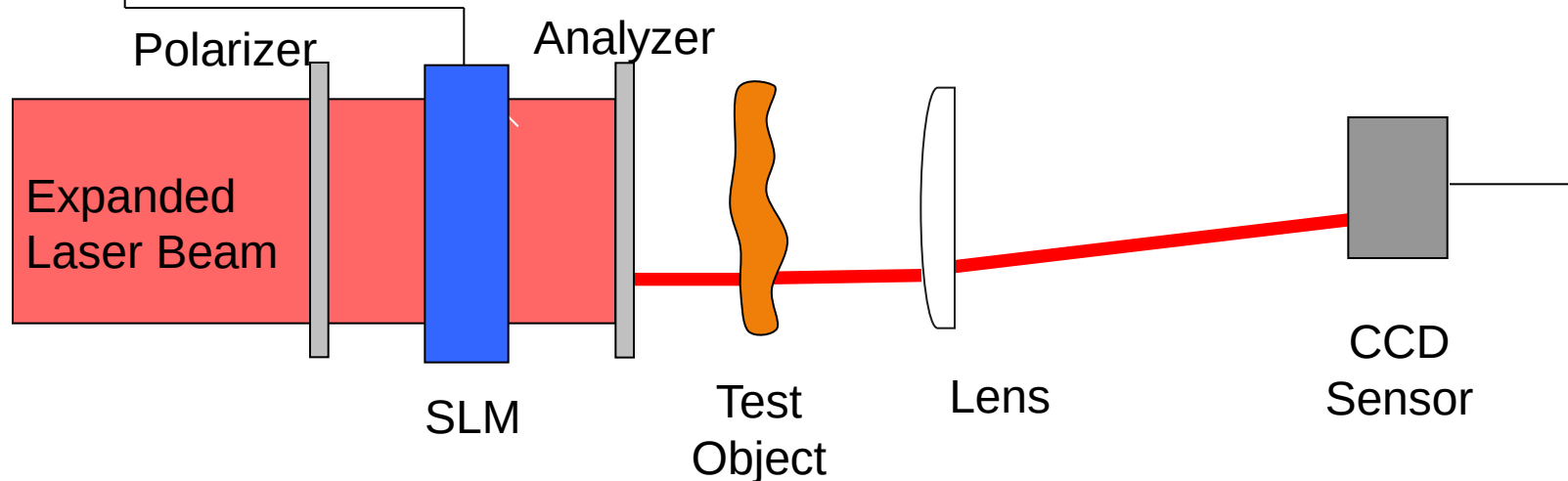


Data screen maps, in time, the sequence of angular deviations caused by the object.

Data Screen



Dynamic range is limited only by the CCD size.



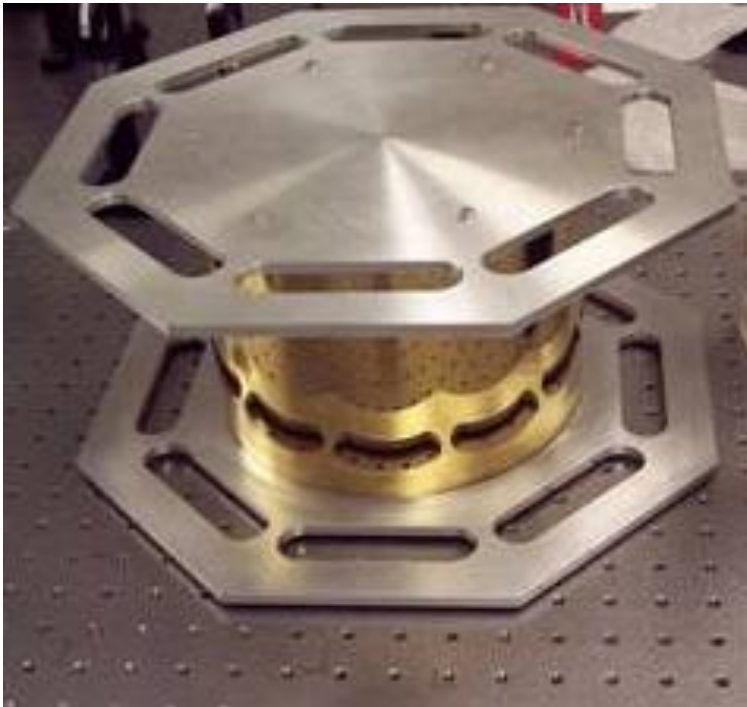
Constraints in conventional Hartmann Testing that can be obviated with this concept

- Sensitivity requires a longer focal length which can cause focused image overlap of adjacent light pencils.
> SLM enables scanning in time, NO OVERLAP
- If the wave is aberrated, the focused spot will not be round, so its centroid is more difficult to locate, further reducing accuracy
> SLM enables preconditioning/aberration correction
- Spatial resolution is limited to the diameter of the lenses in the array.
> Limited to fractional pixel size of SLM

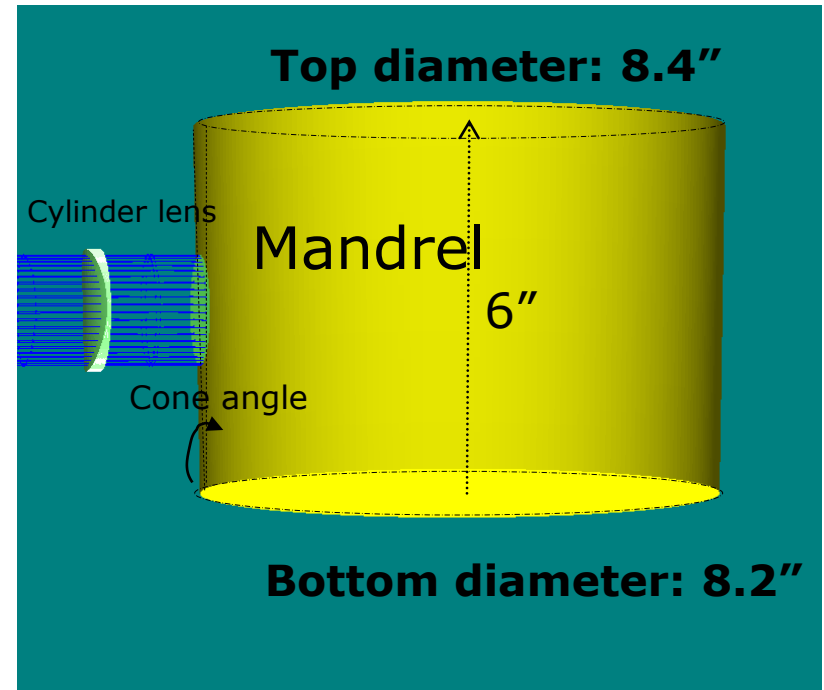


Mandrel Used for Demonstration

Photo of the Mandrel

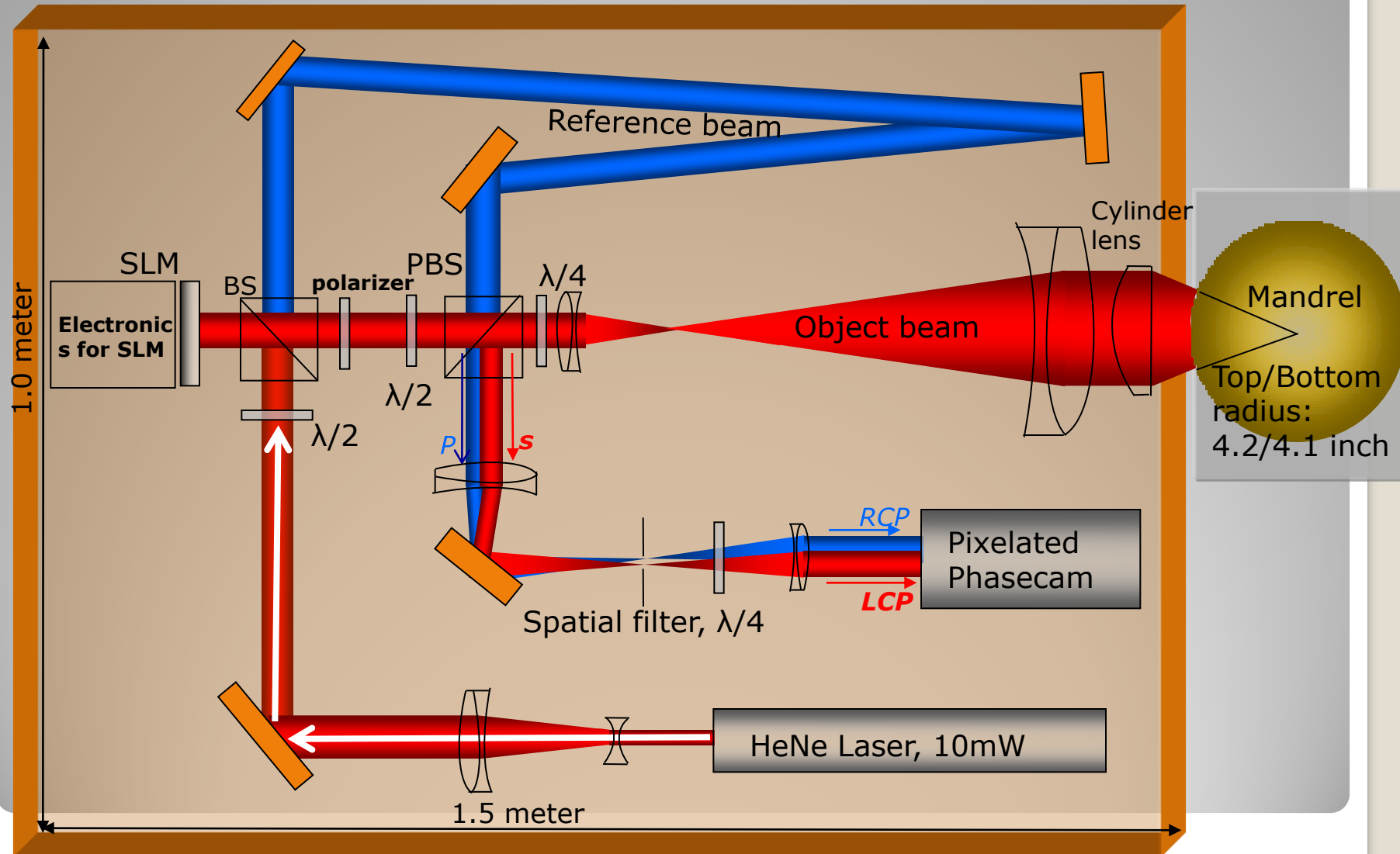


Geometry of the Mandrel

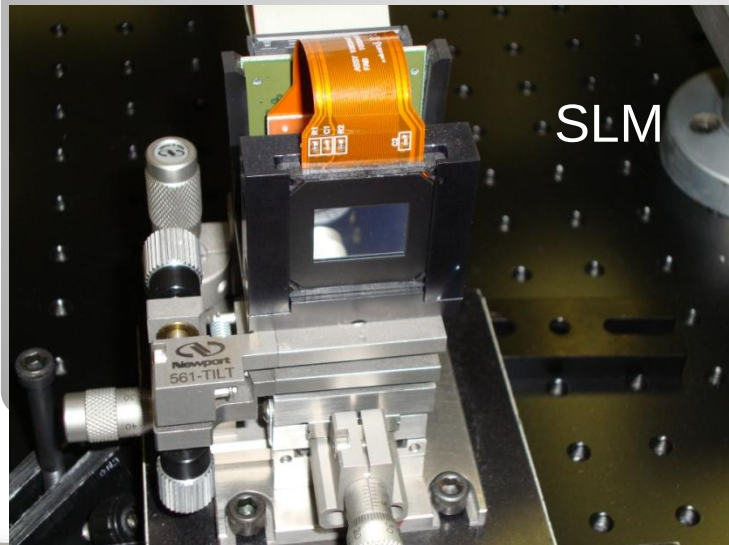
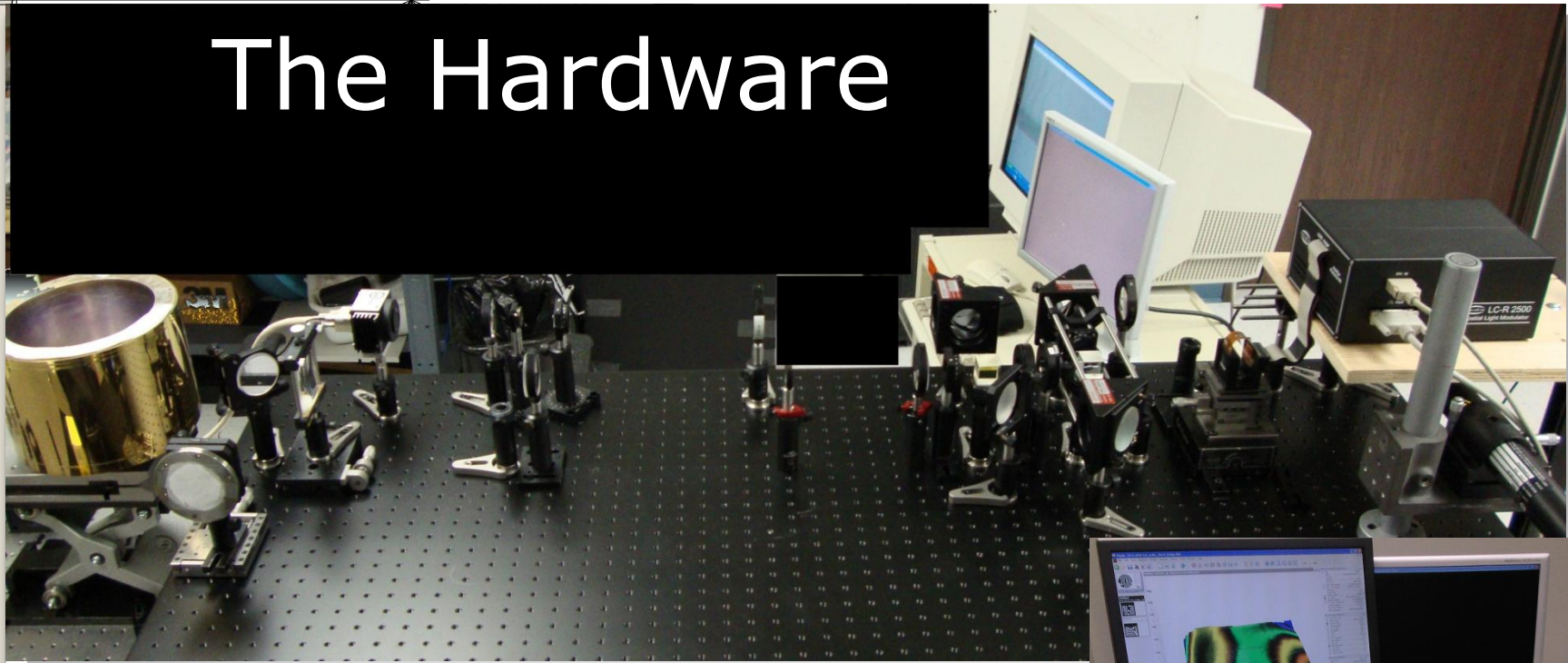


- Mandrel provided by NASA GSFC, 6-inches tall, top and bottom diameters are different, cone shape.

System Design (Top view)



The Hardware



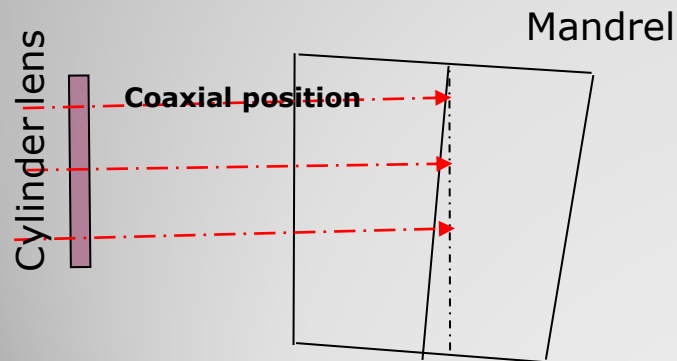
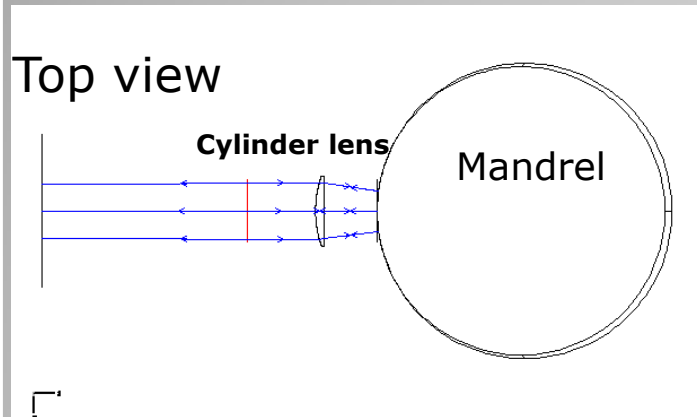
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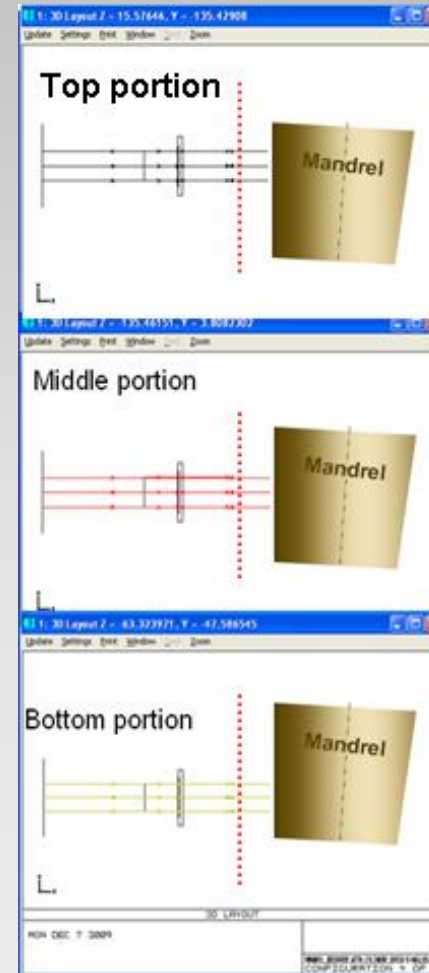
Simulation Results

Beam size at cylinder lens: ~ 1.8 inch
150mm fl cylinder lens

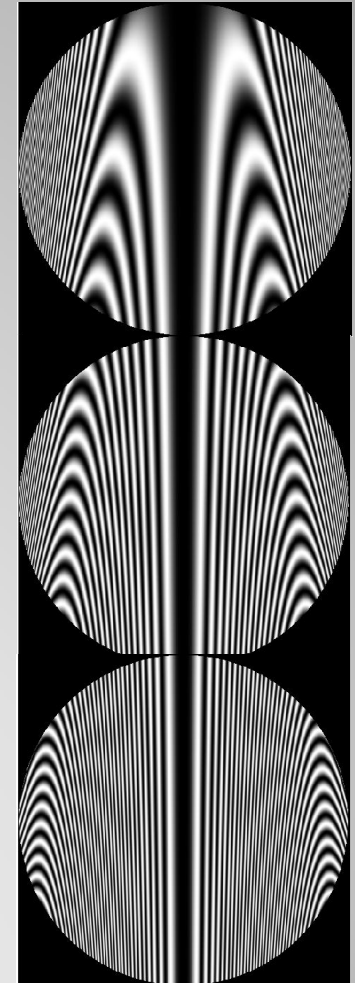
Top view



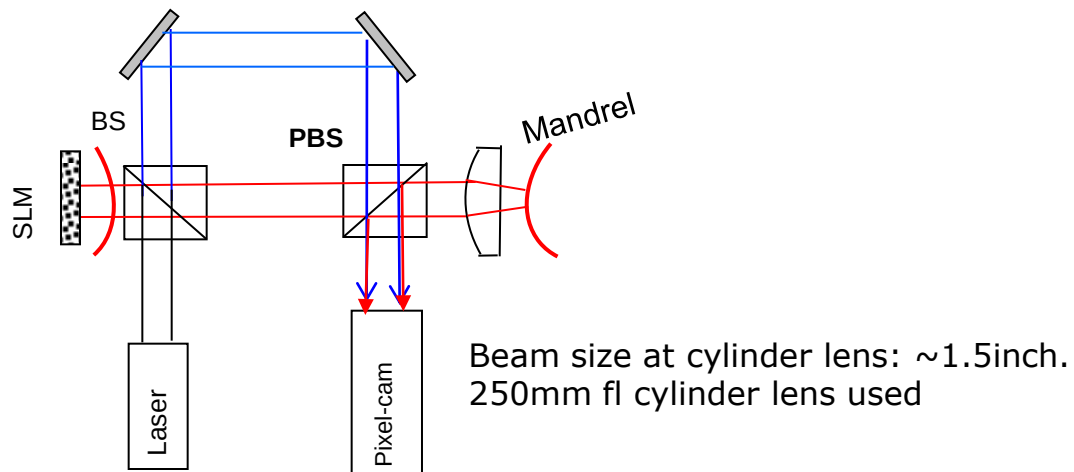
Side View



Interferogram

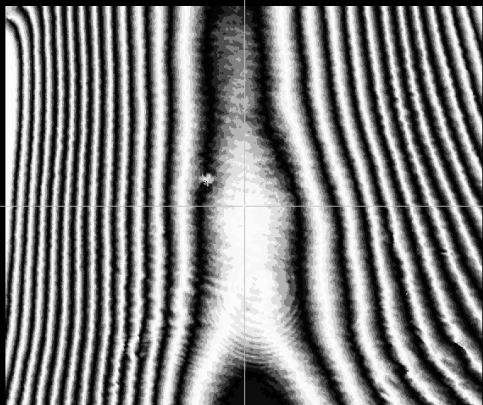


Reducing the number of Fringes from a mandrel

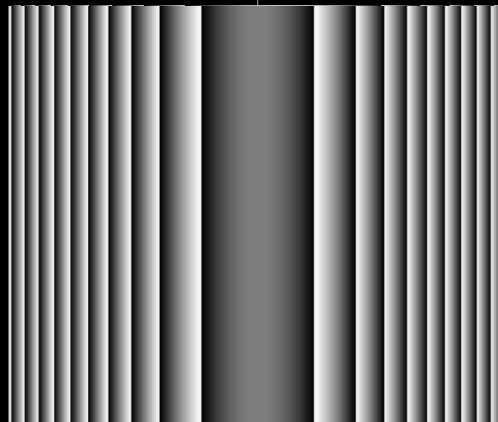


(Various waveplates and telescopes not shown)

Measured Interferogram



Applied Phase on the SLM



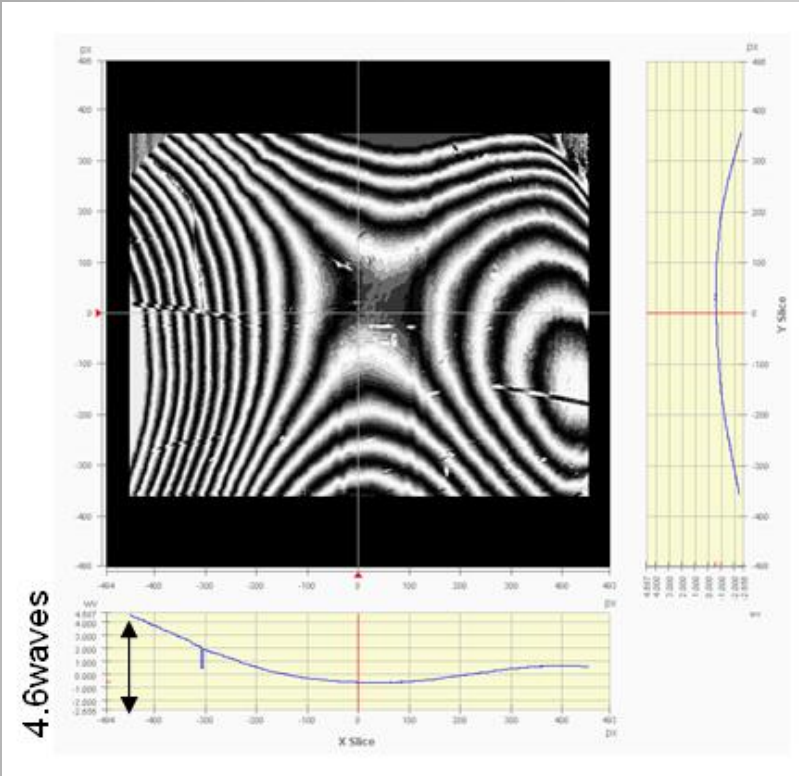
Residual Interferogram



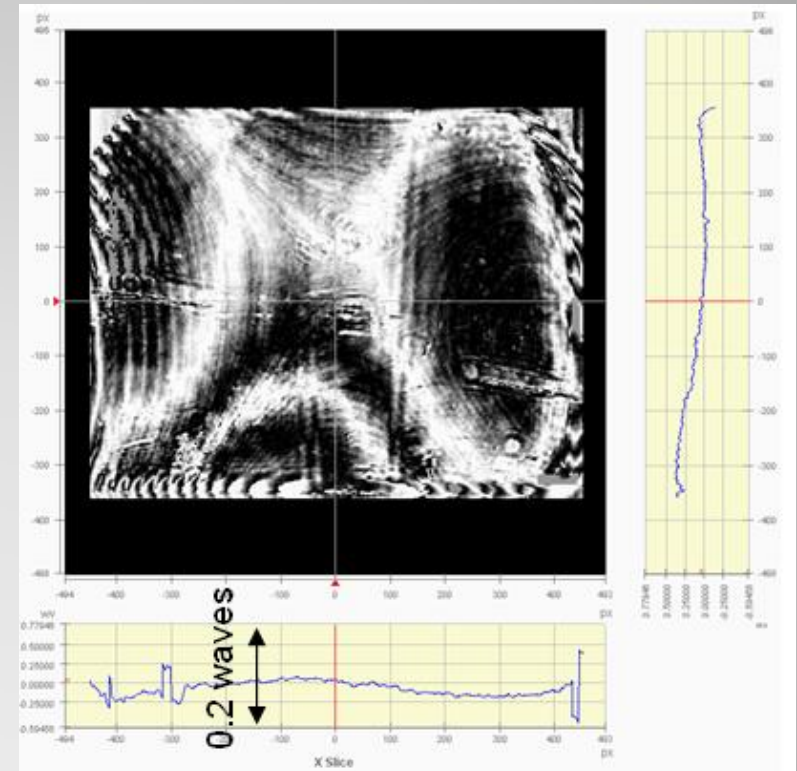


Compensation of one of Mandrel's wavefronts for a one inch beam

Before Compensation

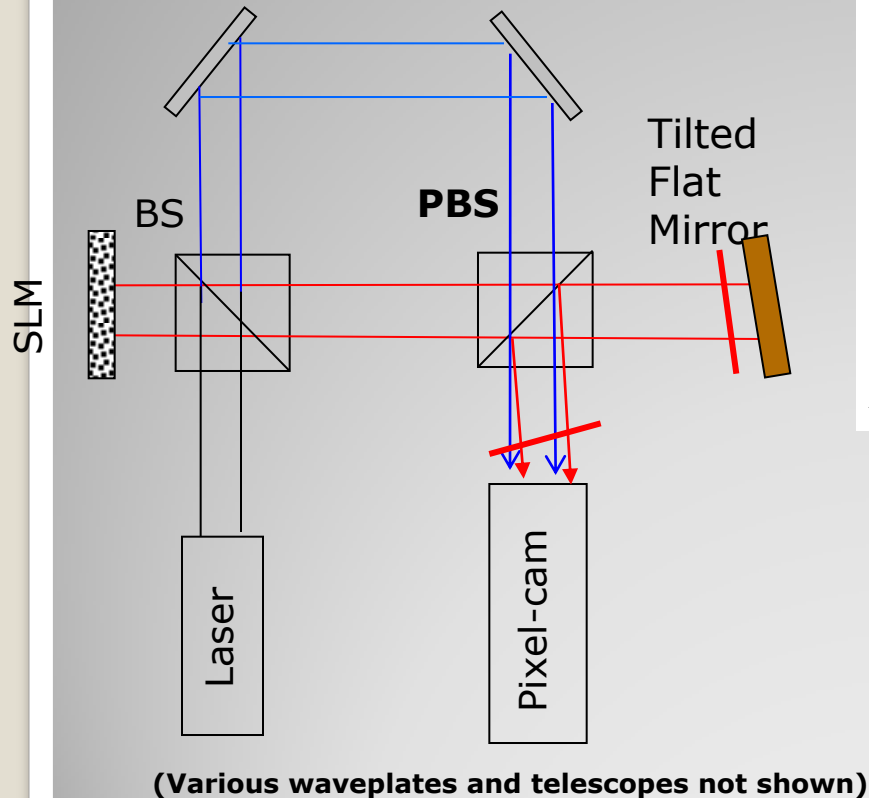


After Compensation

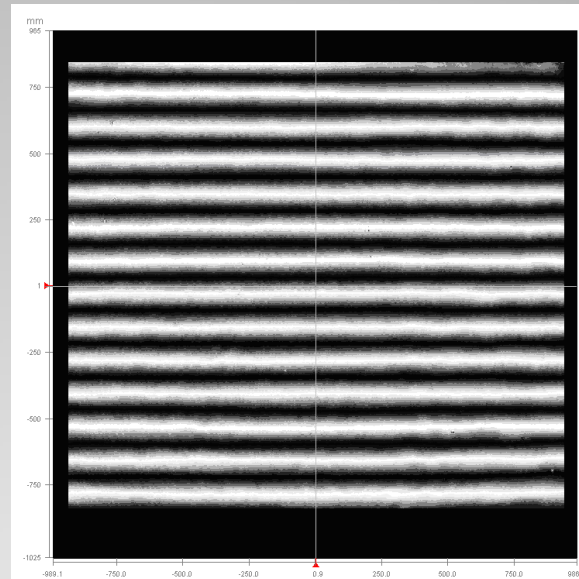


Resolution & Signal to Noise

-Tilt by the flat mirror, ~ 7 waves

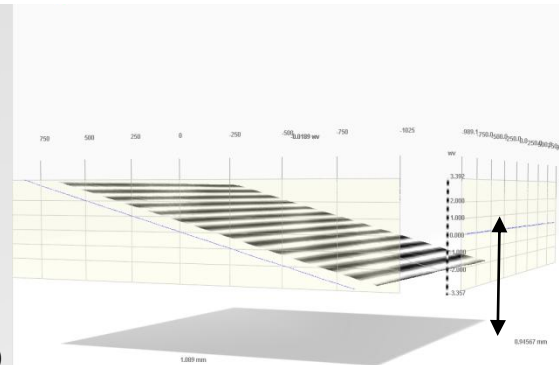


Tilt introduced by flat mirror



Interferogram

~ 14 fringes

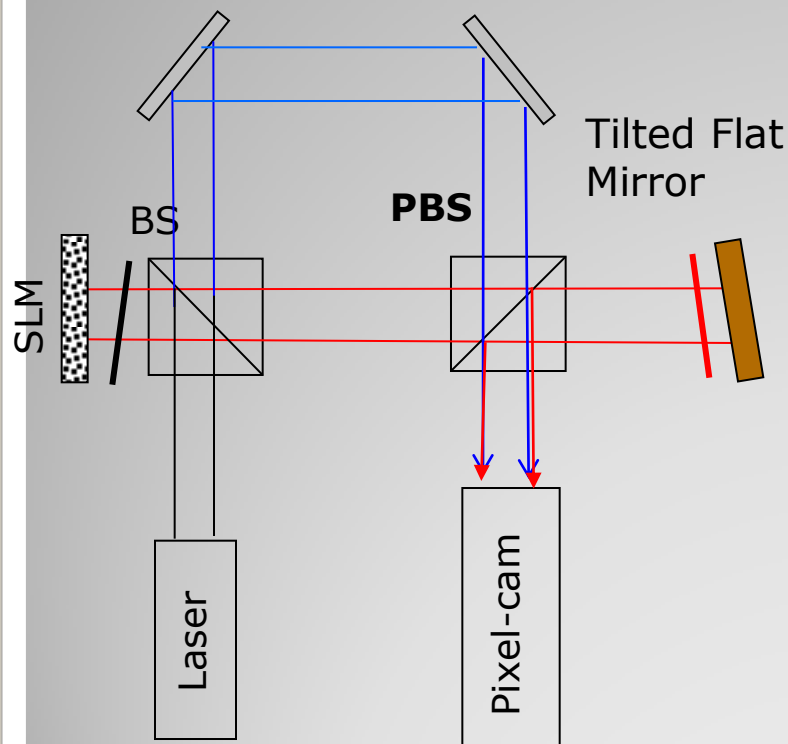


3D view

PV. $\pm 3.4 \lambda$

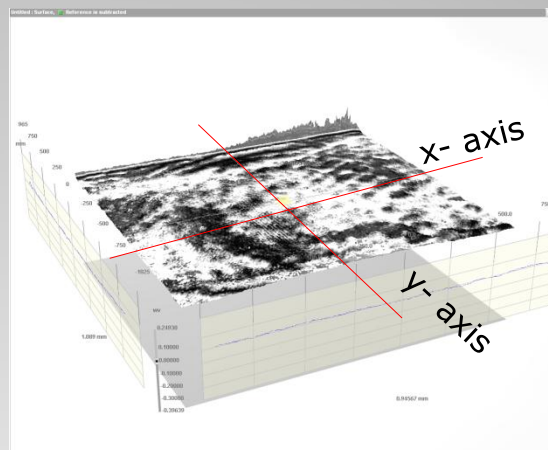
Resolution and Signal to Noise (Cont.)

- Residual wavefronts, less than 0.05λ

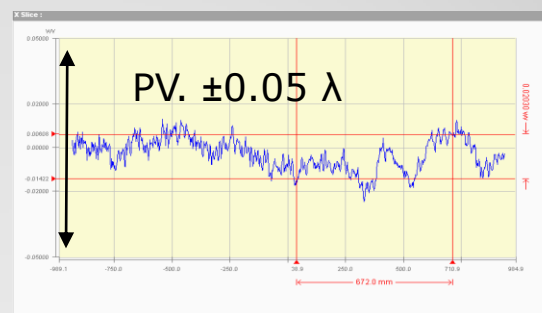


(Various waveplates and telescopes not shown)

SLM can compensate tilted wavefronts with an accuracy of 1/20th wave, or better



Interferogram
(3D view)

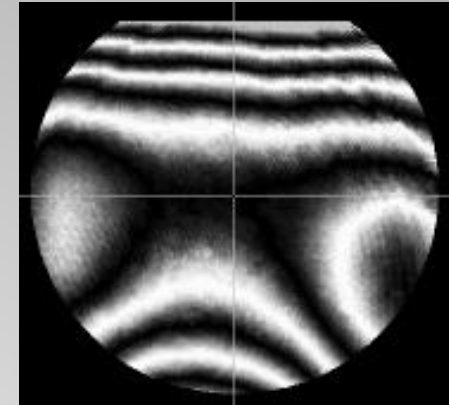
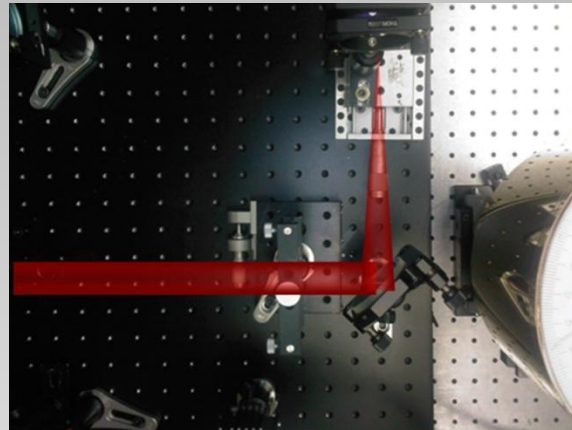
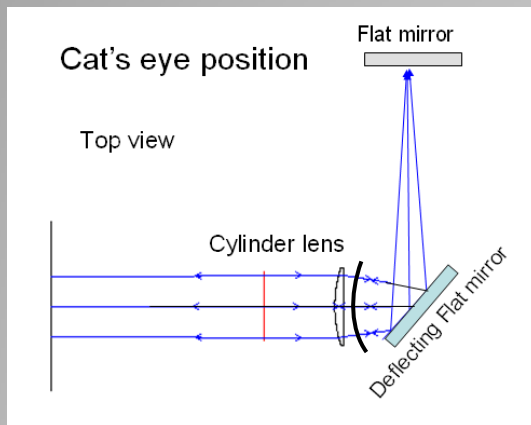


x-profile



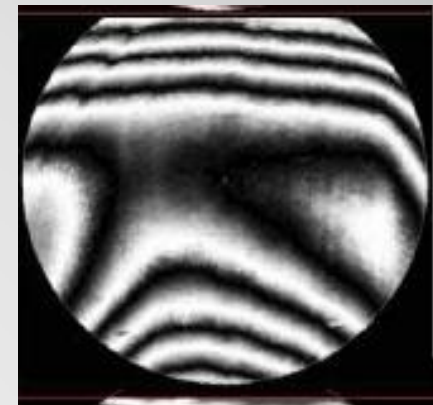
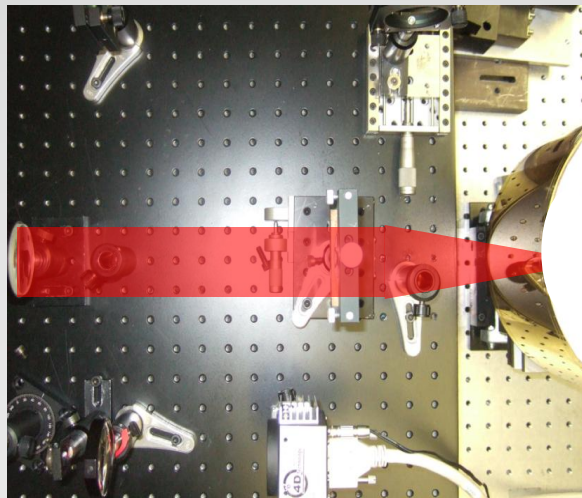
y-profile

Determining System Aberrations with a Cat's Eye



Due to SLM, cylinder lens, optical alignments, etc

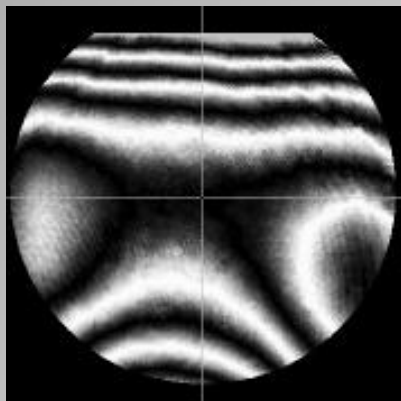
Coaxial position



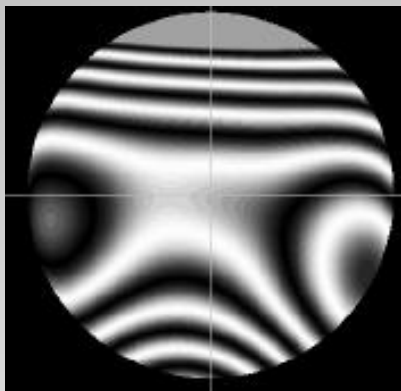


Cat's eye wavefronts by ZEMAX based on Experimental results

Experimental measurement



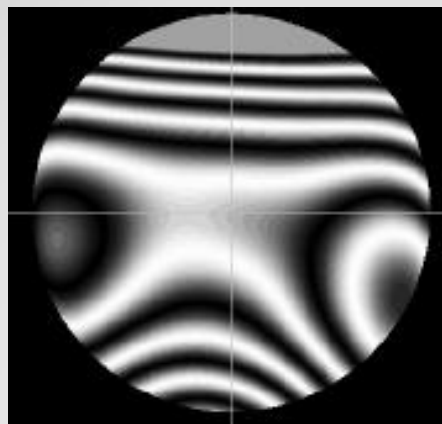
Simulated Fringes



Leftover < 0.1λ



Fitted Zernike Polynomials upto 36 terms, found and put into Zemax

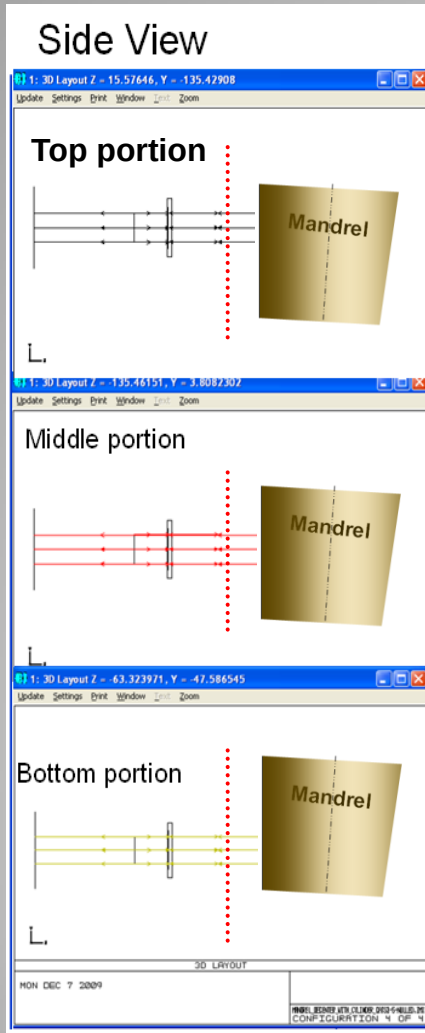


Simulated Cats' eye wavefronts by ZEMAX

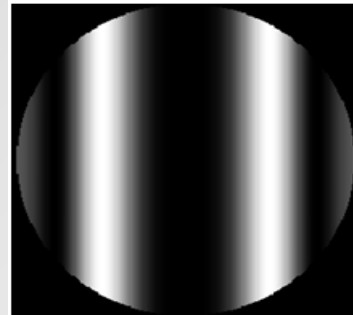
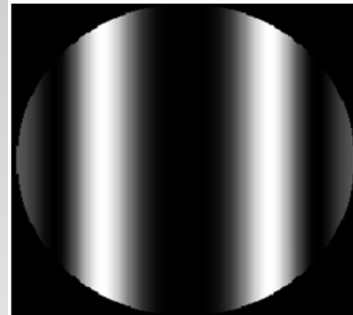
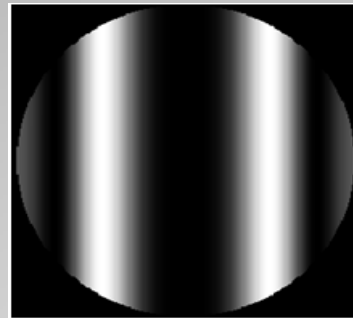


Simulated, No System Aberrations

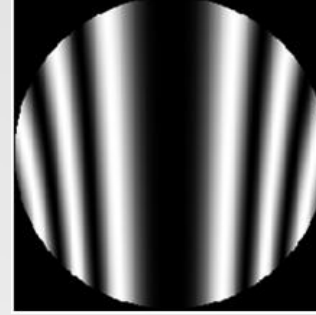
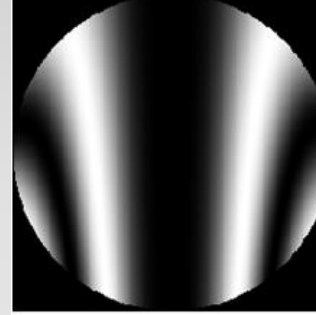
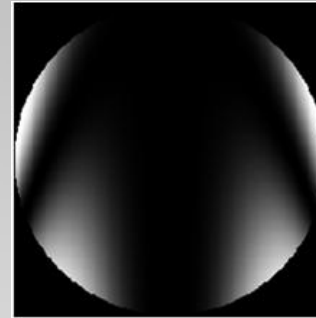
Measured



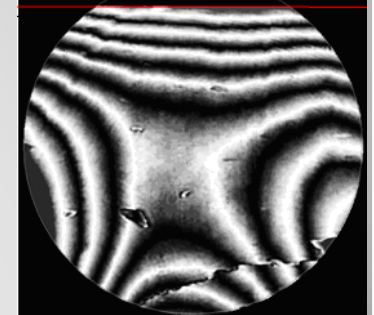
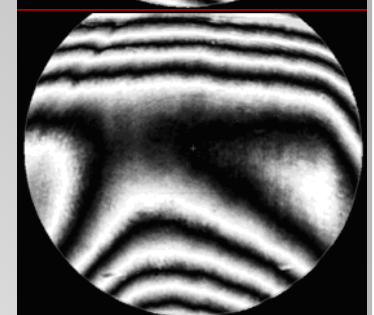
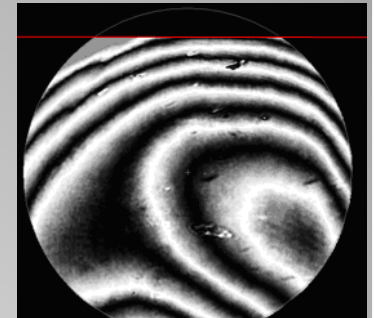
Beam : ~27mm



Wavefronts at cat's
eye position at each
height

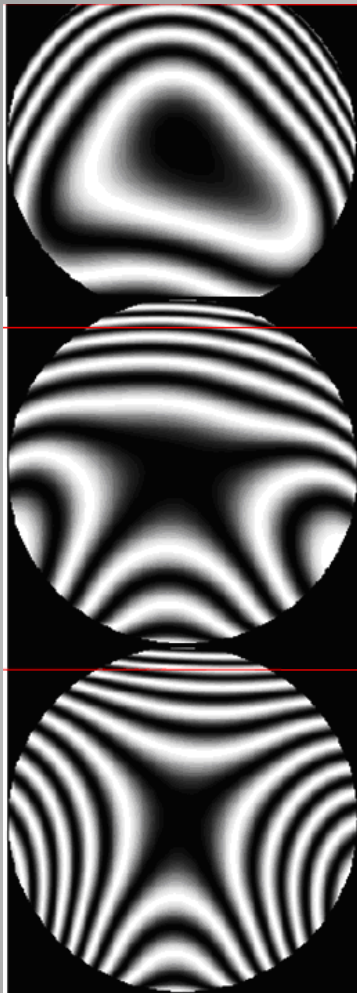


Wavefronts at
coaxial position at
each height

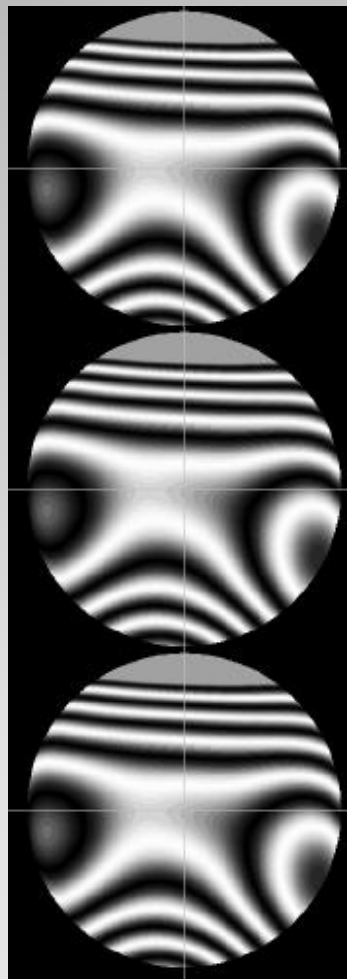


Wavefronts at
coaxial position at
each height

Net fringe due to a Mandrel



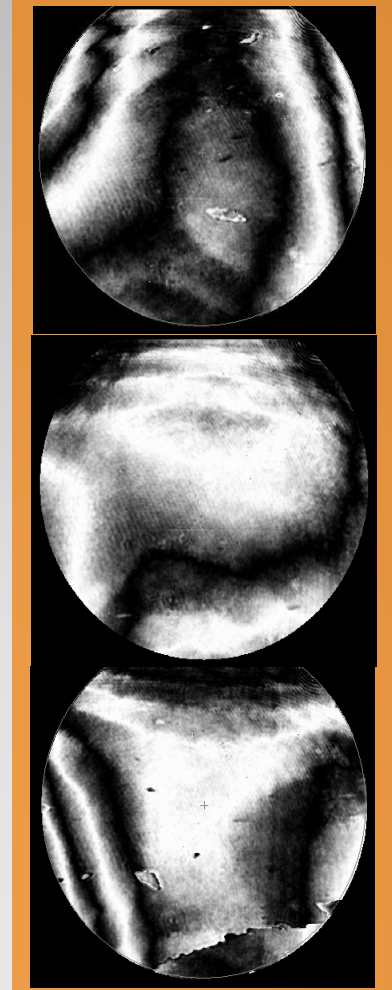
Coaxial position



Cat's eye position



Simulated Net Fringes

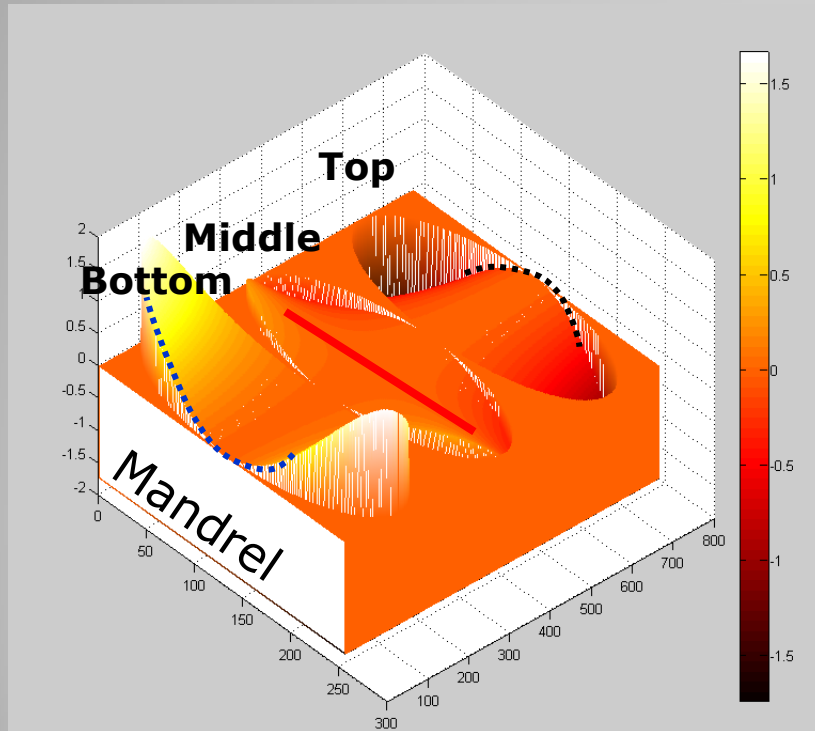


Experimental Net Fringes

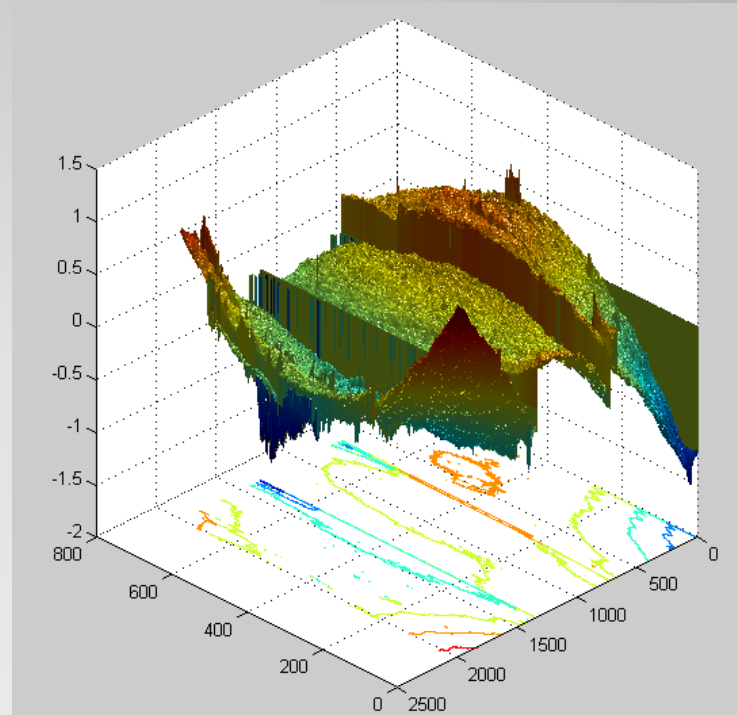


Wavefronts from the Mandrel, II

Simulated Wavefronts
of the Mandrel



Experimental Wavefronts





Unique Features & Applications

- Very wide dynamic range
- Applicable to aspheres and non axis-symmetric optics, i.e. freeform optics
- Enables null point testing
- Enables removing system aberrations

Summary and Conclusions

- Demonstrated wide dynamic range digital interferometry/Hartmann for advanced optical components using an SLM to:
 - provide wavefront preconditioning.
 - Hartmann & Interferometry in the same instrument
 - Scanning Hartmann
 - extended dynamic range.
 - Null point testing.
- Hartmann provides information needed to program SLM for wavefront preconditioning
- SLM Calibration procedure to produce Gamma curve.
- Procedure to subtract system aberrations.
- Accuracy of $\lambda/20$ with SLM in system.
- Concept extends dynamic range by more than 150λ



Future Work

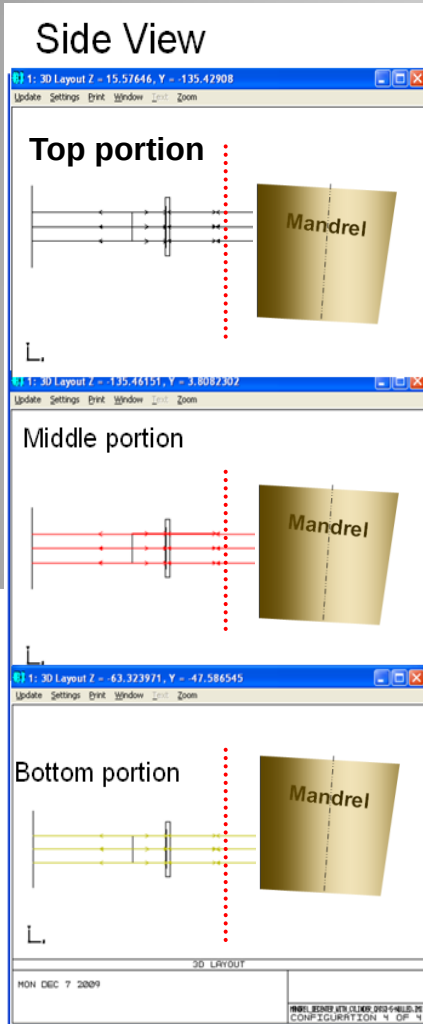
- Software Development
 - Transforming Hartmann data into wavefront Preconditioning data.
 - Automating Calibration
 - Interferogram stitching
 - System automation
- Hardware improvements
 - Calibration to push system accuracy
 - Incorporating improved SLM's
 - System packaging

Back up Slides

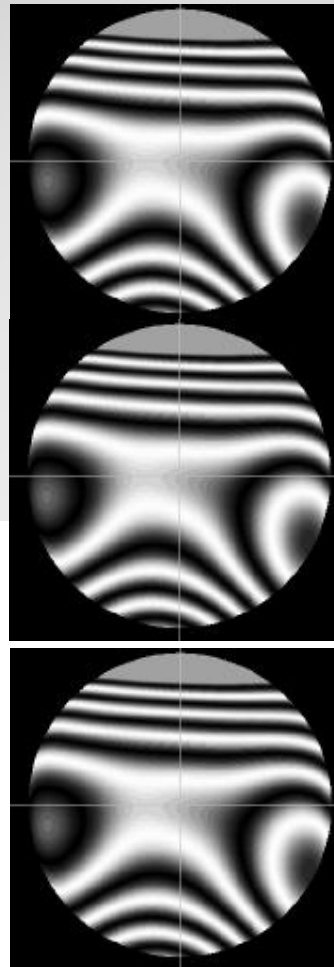
Including System Aberrations

Simulated

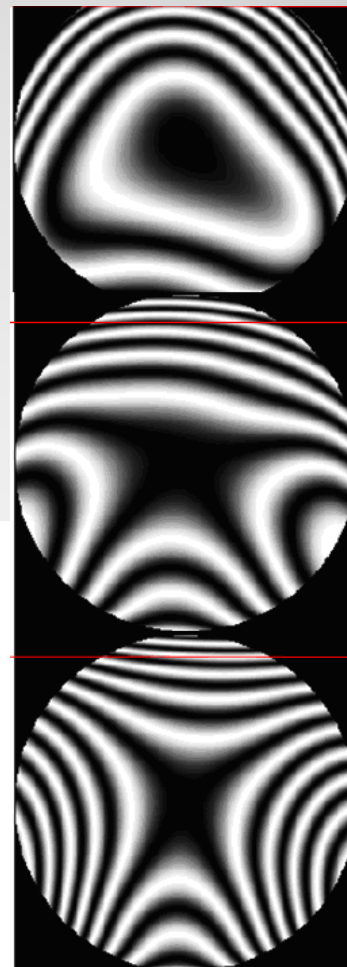
Measured



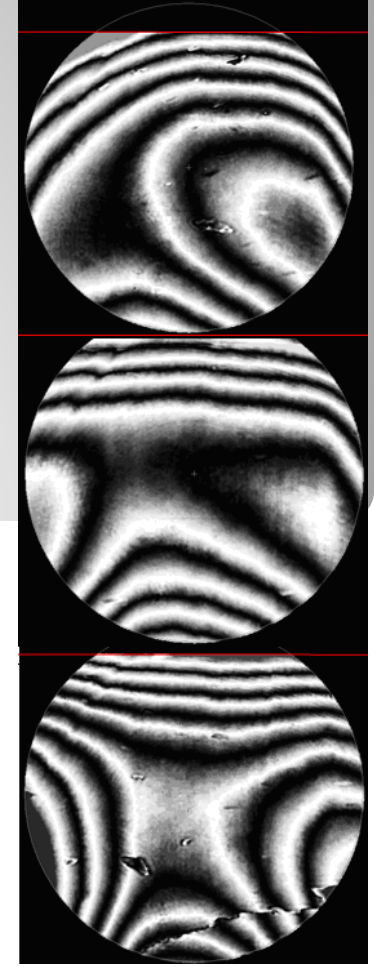
Beam : ~27mm



cat's eye position



coaxial position

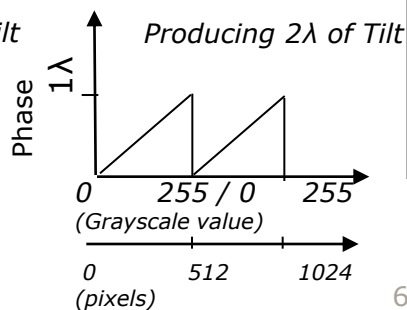
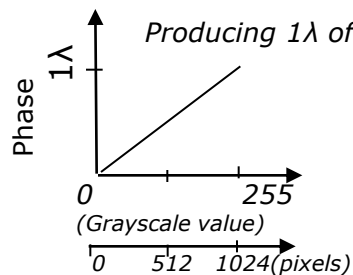


coaxial position

Key Components: SLM, Pixelcam*

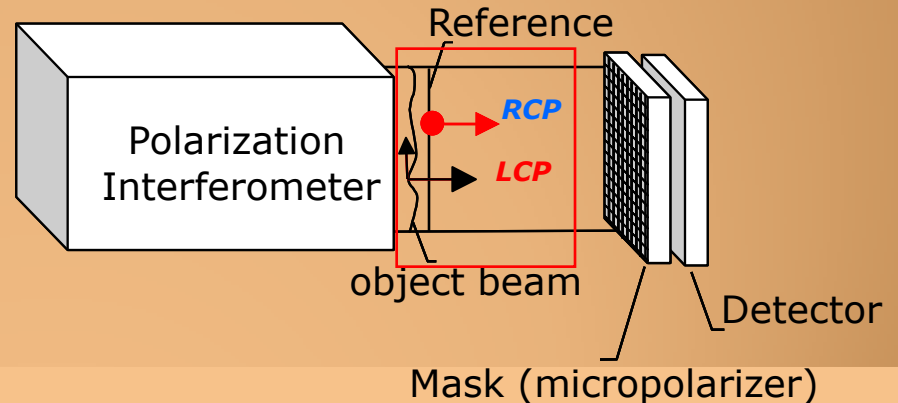
SLM for a compensator

- Programmable holographic optical element
- Produce wavefronts of any shape and can simulate freeform optical surfaces
- Holoeye SLM
 - SLM can produce a phase up to 2π at 632.8nm
 - Assign 0 to 255 grayscale values to 0 to 2π (or 1λ)
 - Can generate higher phases by wrapping phase
 - Can provide more than 150 wave tilt
 - Pixel size ~ 19 microns.
 - 19.5 x 14.6mm size (1024 x 768 pixels)



Pixelated Phasecam for a detector

- Spatial phase shifting interferometer
- Single shot, insensitive to vibration



0	90
270	180

Phase information of the object, $\Delta\phi(x, y)$ can be obtained from the 4 intensities on each unit cell.

$$I(x, y) = \frac{1}{2} (I_r + I_s + 2\sqrt{I_r I_s} \cos(\Delta\phi(x, y) + 2\alpha_p))$$

*Produced and Trademark by 4D Technologies, Inc, Tucson, AZ <http://www.4dtechnology.com>

Enhanced Interferometry with a Programmable Spatial Light Modulator

MetroLaser Incorporated
Irvine, CA

INNOVATION

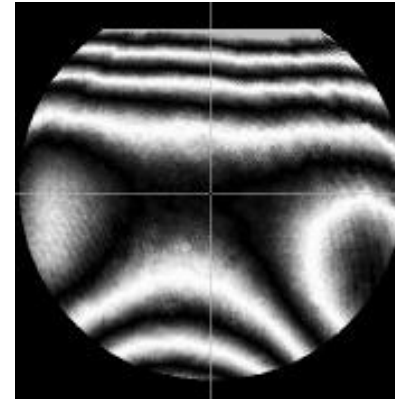
A unique optical inspection system incorporates a dynamic holographic optical element that combines and extends both Hartmann and digital interferometry with preconditioned wavefronts. The resulting system exhibits a wide dynamic range and will be especially useful for inspecting aspherical and free form optics

ACCOMPLISHMENTS

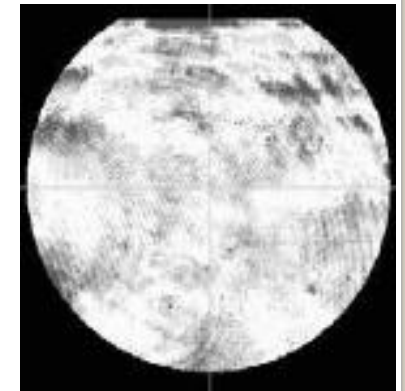
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- ❖ Accuracy of $\lambda/20$ with SLM in system.

COMMERCIALIZATION

- ◆ Basic concept patent application has been submitted
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Before Correction



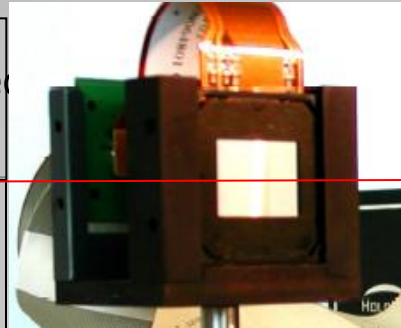
*After Correction
Wavefront Preconditioning*

APPLICATIONS/FUTURE WORK

- ◆ Government and Commercial Applications
 - X-ray telescope mirrors and mandrels
 - Free form Optics
 - Aspherical Optics
 - SiC telescope Mirrors
 - Ogive windows
- ◆ Future Work
 - Additional Software for System Operation and Automation
 - System Packaging
 - Noise floor and resolution improvement

SLM Calibration

SLM



Gray scale value is varied on upper half of SLM

Gray scale value is held constant at 0 on lower half of SLM

Applied gray values on the SLM

Gray value: 63

...

Applied gray values on the SLM

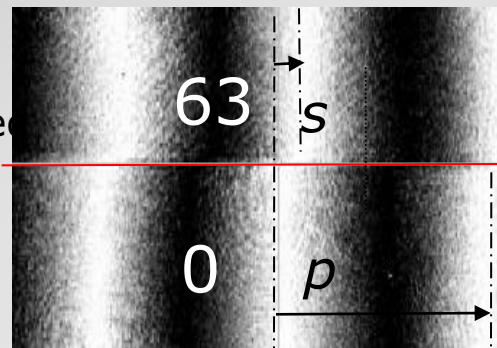
Gray value: 127

Gray value: 0

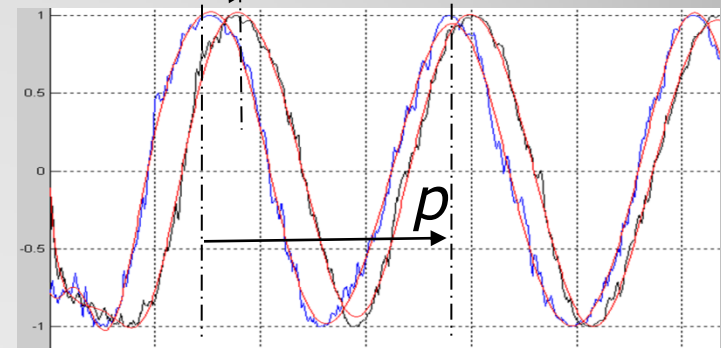
Gray value: 0

Interferogram Result

Phase shift vs. gray scale value is measured interferometrically

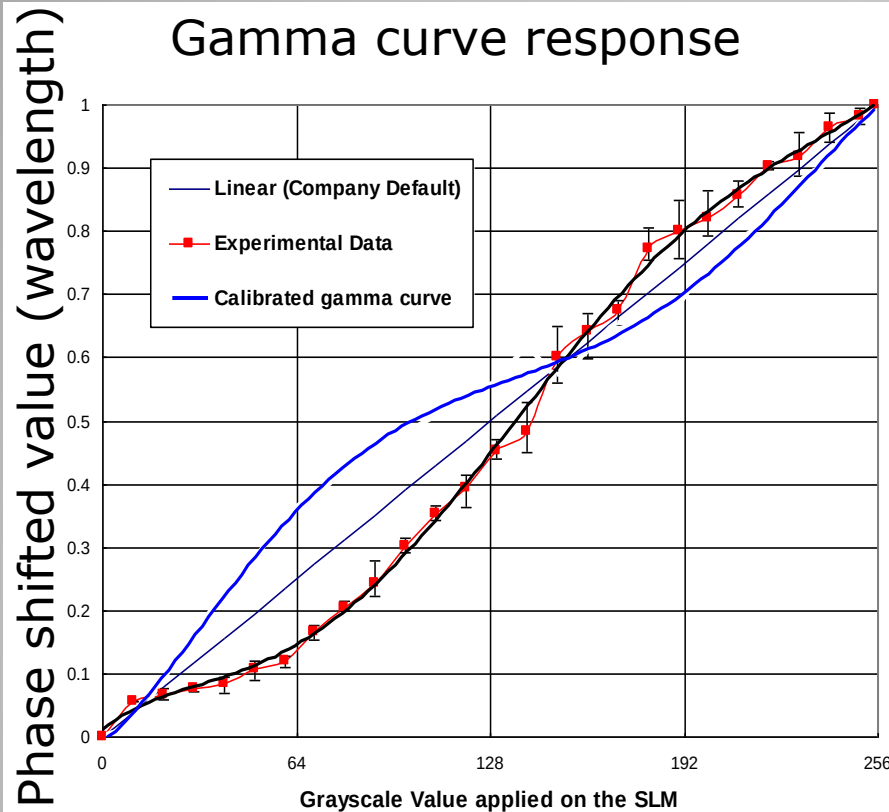


S: Phase shifted amount
P: Period of fringes



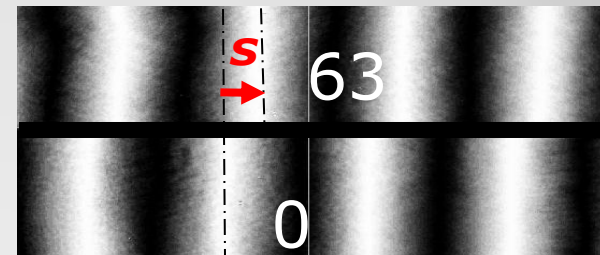
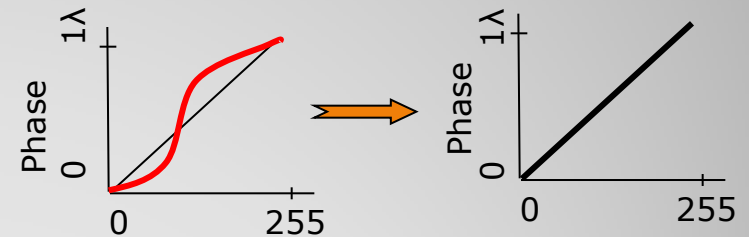
- Relative phase shift recorded to achieve an accuracy of data
- Gray value was varied from 0 to 255 on the top half, while it was held constant (0) on the bottom half
- Phase shifted vs applied gray value on the SLM produces gamma curve

Corrected Gamma Curve



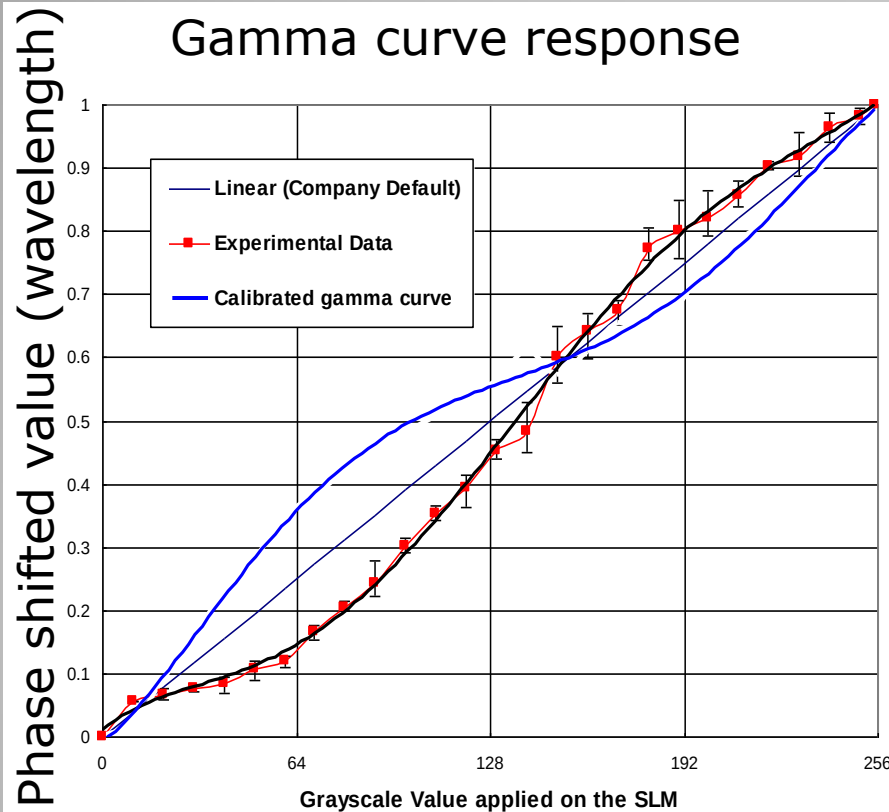
The calibrated gamma curve (blue color), is used to linearize the phase vs. gray scale response (red curve)

Errors when producing 1λ of tilt and corrected one after calibration



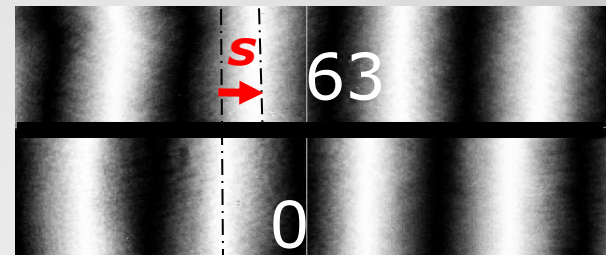
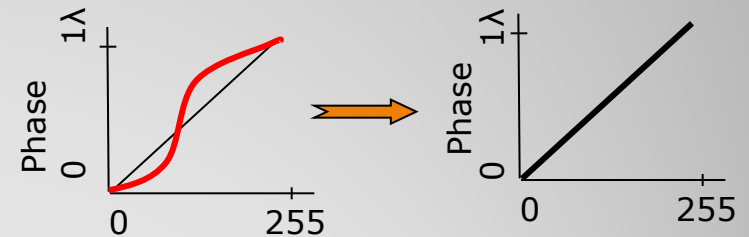
$S = 0.25\lambda$ shift when 63 value applied

Corrected Gamma Curve



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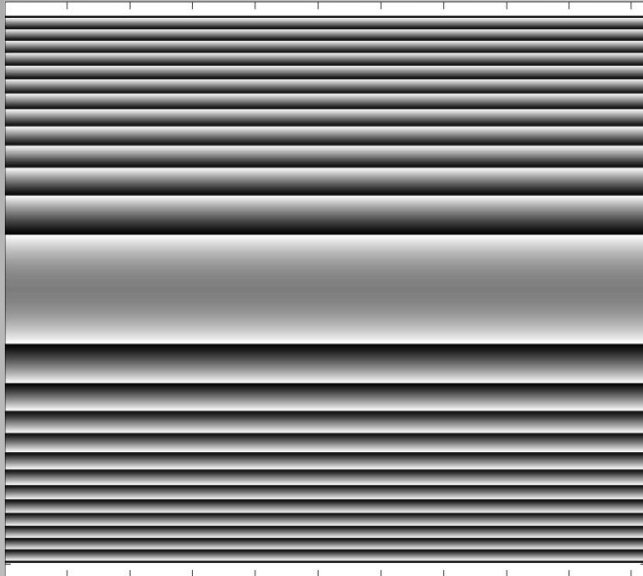


$S = 0.25\lambda$ shift when 63 value applied



Producing cylindrical wavefronts

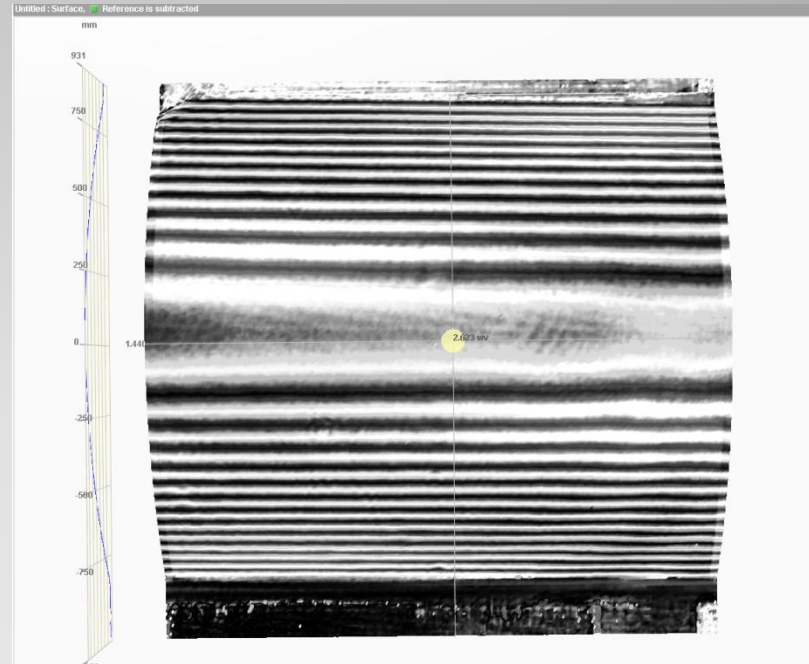
Applied Phase on the SLM



$$\Phi \sim a y^2$$

Matlab formula used to generate the phase.

Interferogram on the pixelcam

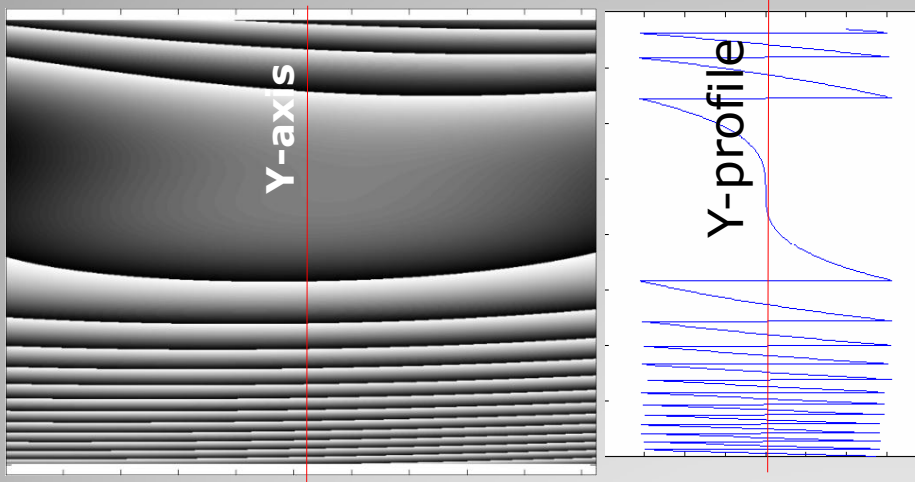


- Cylindrical wrapped phase written by Matlab code and applied on the SLM, which is good candidate compensating phase for a Mandrel.
- Phase produced and shown by the interferogram on the pixelcam

Producing arbitrary wavefronts

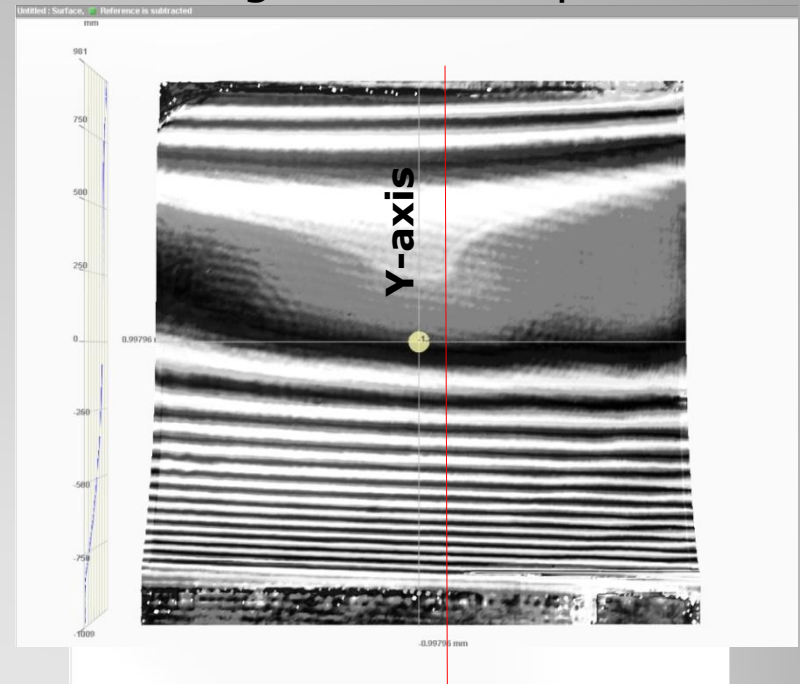
Interferogram on the pixelcam

Applied Phase on the SLM

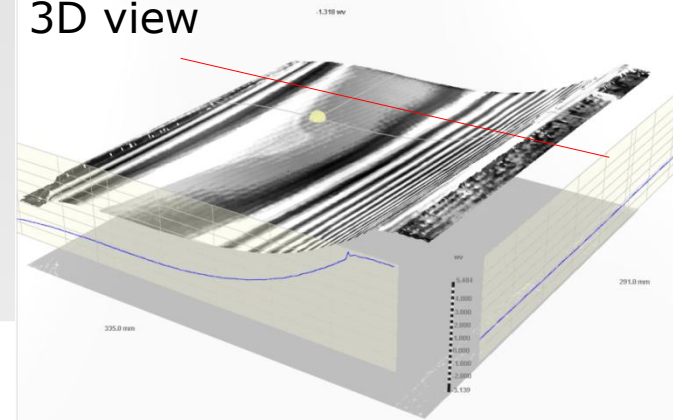


$\Phi = 20 \cdot (y - 0.25)^3 - x^2 + x \cdot y + x \cdot y^3$
 Matlab formula used to generate the phase

- Arbitrary wrapped phase written by Matlab code and applied on the SLM
- Phase produced and shown by the interferogram on the pixelcam

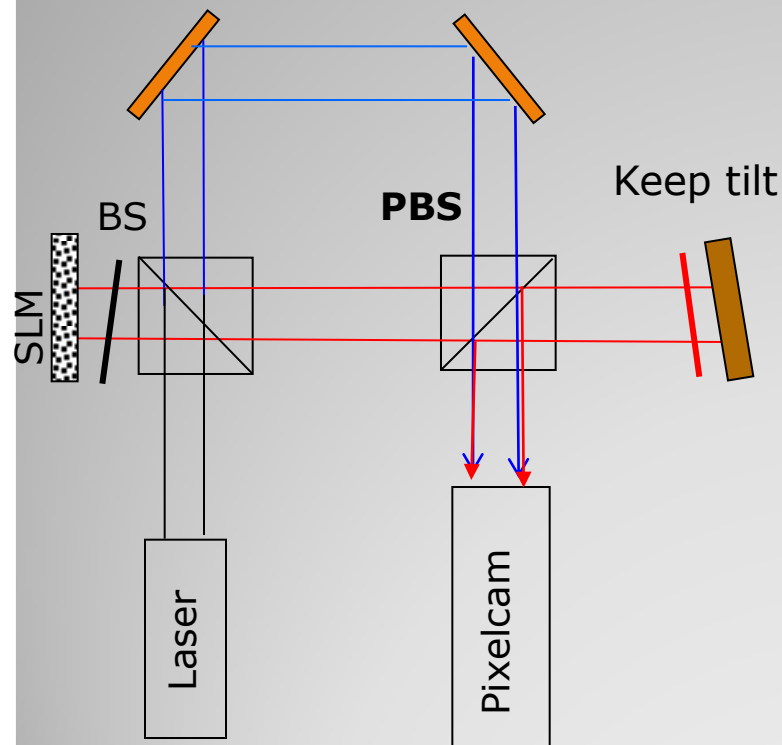


3D view



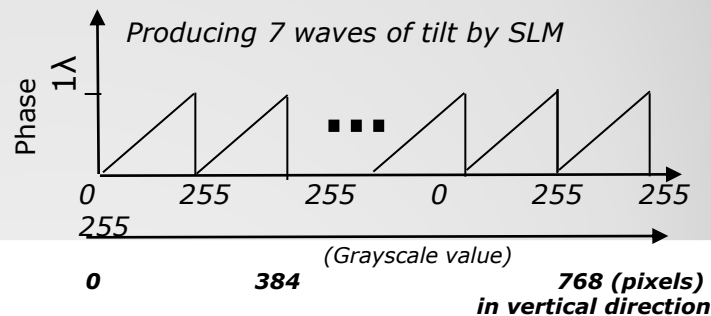
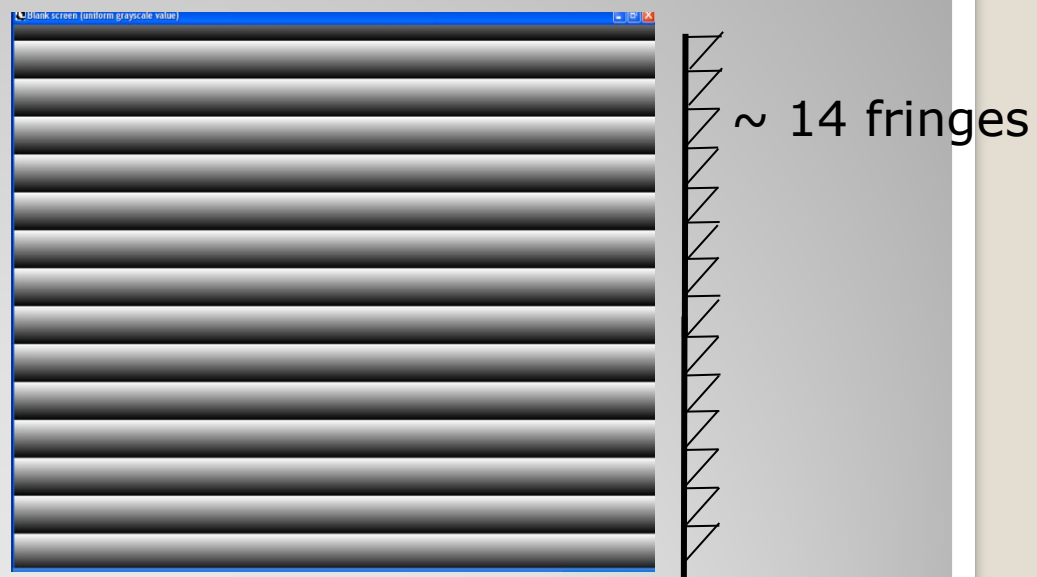
Compensation of Tilt

-Applied compensating tilt by the SLM



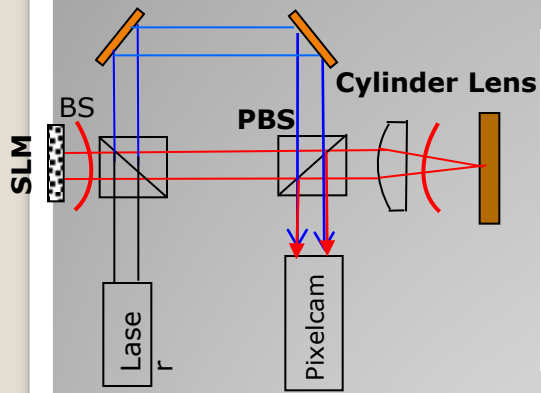
(Various waveplates and telescopes not shown)

Wrapped phase applied on the SLM

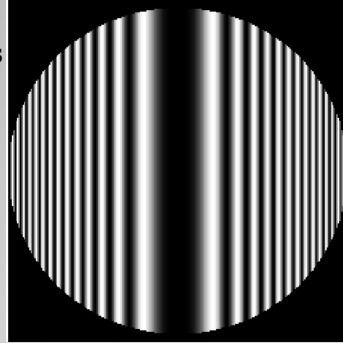


Compensating wrapped phase applied on the SLM

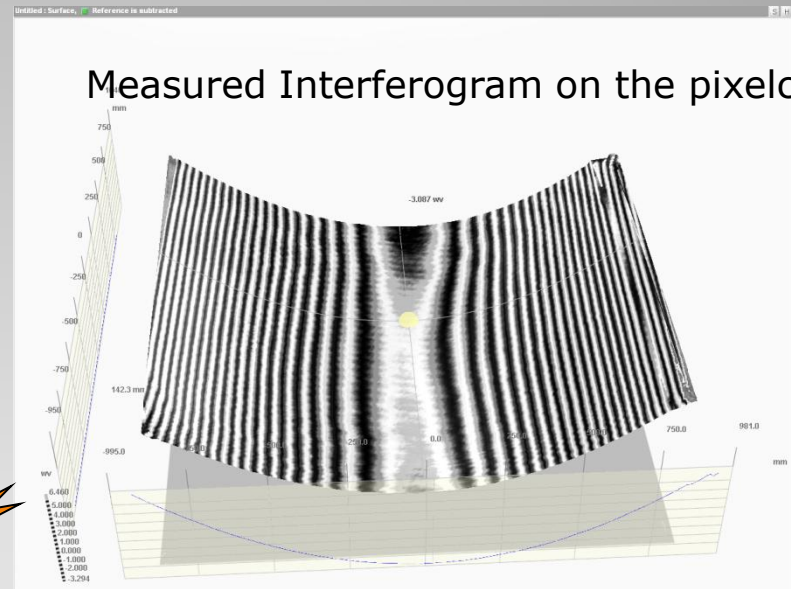
Compensation of Cylindrical wavefronts



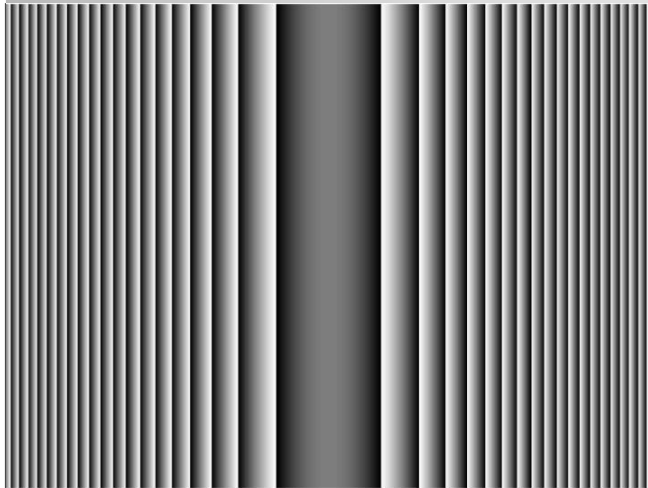
Simulation



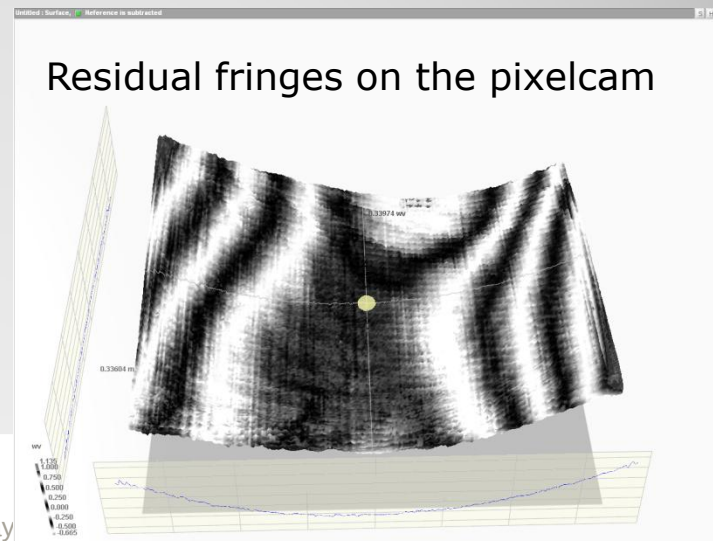
Measured Interferogram on the pixelcam



Applied Phase on the SLM

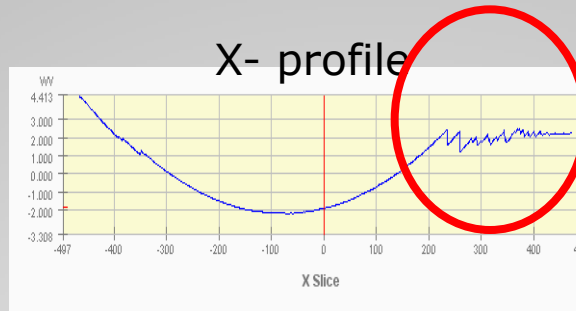
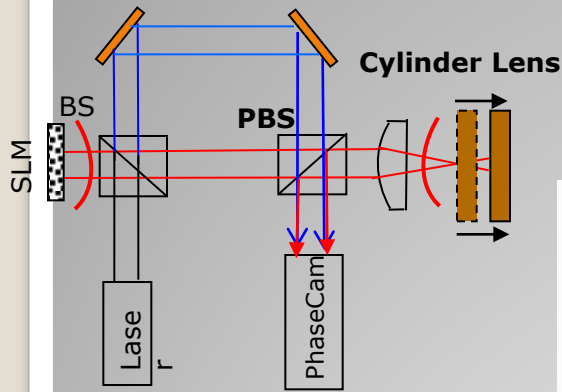


Residual fringes on the pixelcam

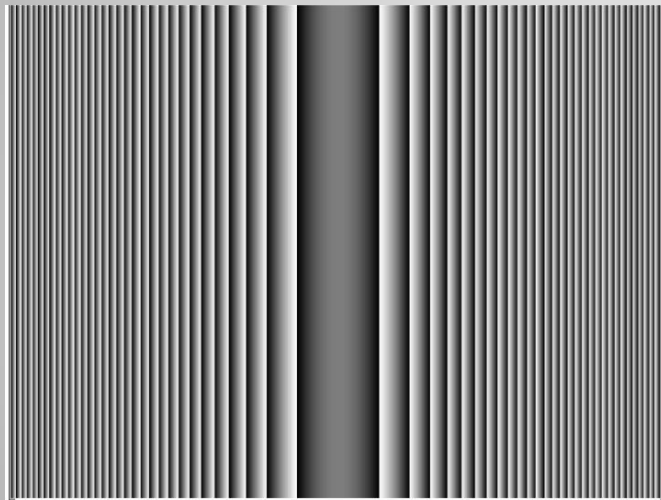


Compensation of Cylindrical wavefronts

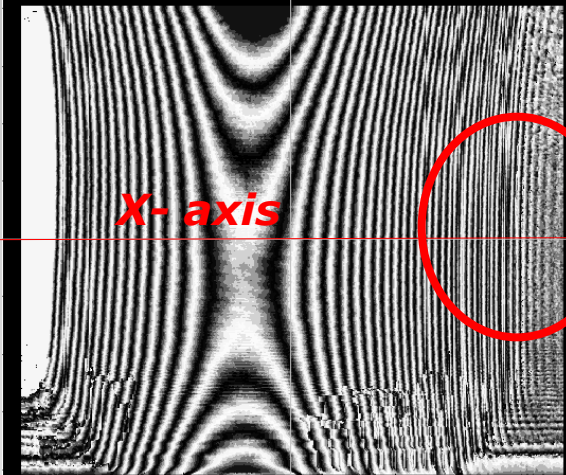
Extreme Case



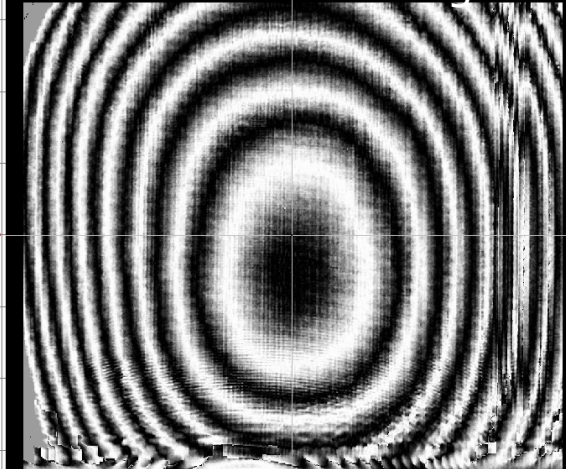
Applied Phase on the SLM



Measured Interferogram



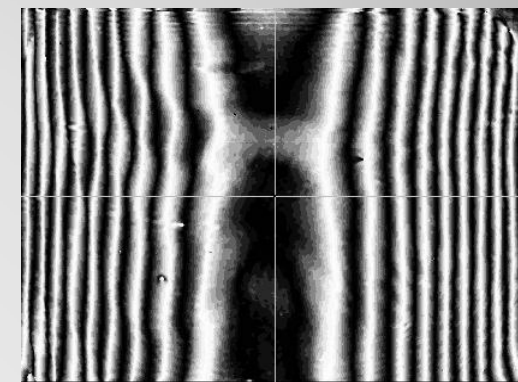
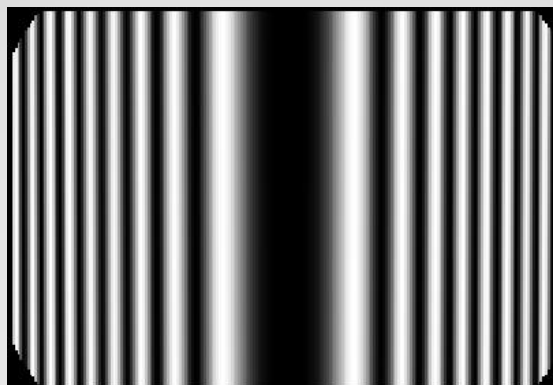
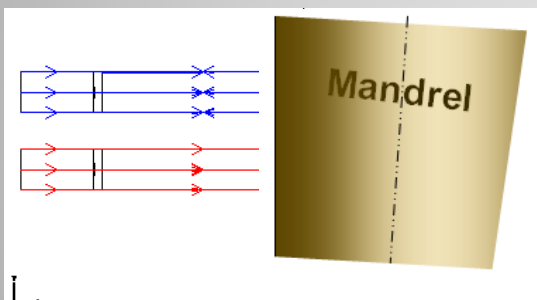
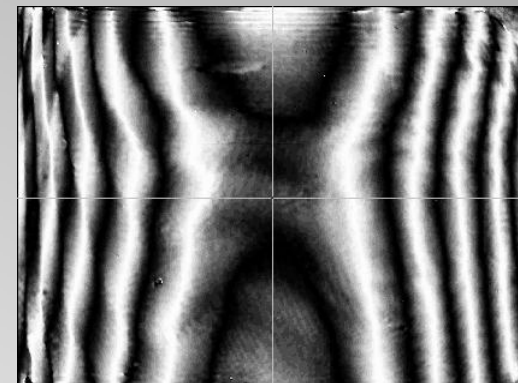
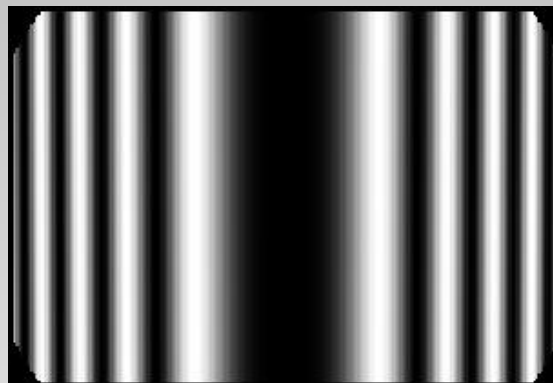
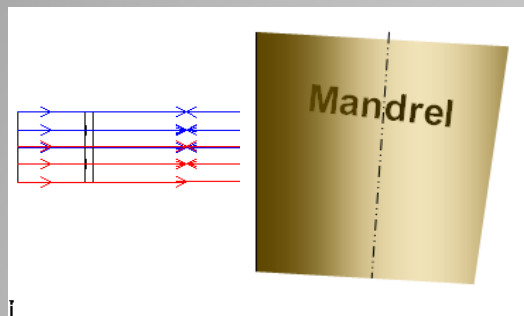
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The phase on SLM is used to bring the dense fringes down into the measurement range of the pixelcam. i.e: extended dynamic range.



Wavefront Differences at Different Heights of Mandrel



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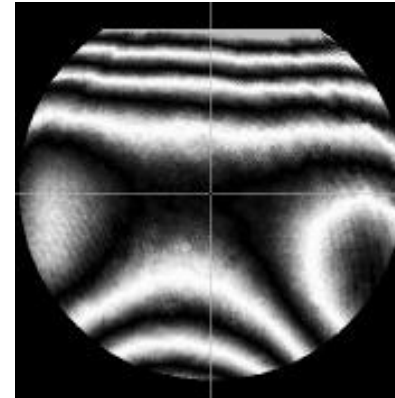
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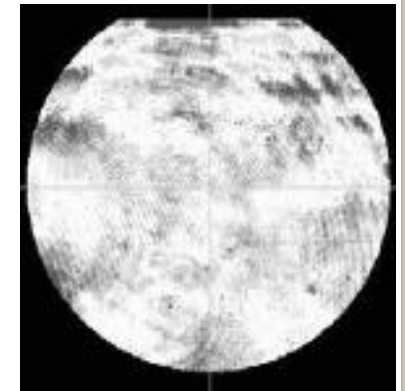
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